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*Synthesis and physicochemical properties of novel,
tungsten-oxide double perovskites and their derivatives
doped with selected rare-earth ions.*

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List of abbreviations

XRD - X-ray Diffraction,	TEY – Total Electron Yield
NPD – Neutron Powder Diffraction,	SDD – Silicon Drift Detector
XPS - X-ray Photoelectron Spectroscopy,	HT- High Temperature,
XAS - X-ray Absorption Spectroscopy,	BCW – Ba ₂ CeWO ₆ ,
EPR - Electron Paramagnetic resonance,	CCWO – Ca ₃ Ce ₂ W ₂ O ₁₂ ;
FTIR – Fourier Transform Infrared,	BPW – Ba ₂ PrWO ₆
DSC – Differential Scanning Calorimetry,	SCWO – Sr ₉ Ce ₂ W ₄ O ₂₄
TG – Thermogravimetry,	SG – Space Group,
DPs – double perovskites,	ICSD – Inorganic Crystal Structure
GS – Goldschmidt factor,	Database,
mGS – modified/modern Goldschmidt	CCDC – Cambridge Crystallographic Data
factor by Bartel et al.,	Centre
NUV – Near Ultraviolet,	JCPDS – Joint Committee on Powder
NIR – Near Infrared,	Diffraction Standards,
CCD – Charge Coupled Device,	RE – rare-earth (ions or atoms),
NA – Numerical aperture,	PL – Photoluminescence,
FWHM – Full Width at Half Maximum;	IR – Infrared,
Cp – Heat Capacity,	DFT – Density Functional Theory,
SEM – Scanning Electron Microscopy;	BE – Binding Energy,
HP – High Pressure,	E _g – Energy gap
DAC – Diamond Anvil Cell,	E _a – Activation energy
V _o – Oxygen vacancy,	E _F – Fermi energy
O _i – interstitial oxygen,	CB – Conduction Band
PTM – Pressure Transmitting Medium,	VB – Valence Band
RT – Room Temperature	CTB – Charge transfer band
LT – Low Temperature,	SSR – Solid State Reaction;

Abstract (Angielski)

This comprehensive work showcases four novel rock-salt type compounds as amphoteric rare-earth tungstates $A_xRE_yW_zO_{6z}$. Two of them belong to classic double perovskites, and the other two are more related to so-called ilmenites or garnets. All of them are acquired as microcrystalline powders via high-temperature solid-state reactions of pellets in inert atmospheres. This work aims to create and thoroughly describe novel materials that could possess superior optical or scintillating properties to their scheelite predecessors, like $BaWO_4:Ce/Pr$. Furthermore, due to precursors' unique PL behavior & polymorphism, a high chance of finding an efficient energy convertor or radiation temperature sensor was anticipated. Therefore, 3 of the investigated compounds were chosen to host cerates at their RE occupation sites (RE=Ce) and the remaining one in the form of praseodymate (RE=Pr). The former group was targeted to be double perovskites varying only by alkaline 2+ atoms (A= Ba, Sr, Ca^{2+}). However, due to significant ionic radii differences, the last two turned out to be rather large, tilted, and less ordered structures. Ilmenites and garnets are popular side products during the double perovskite production process but are also worth investigating. Regarding praseodymium, only the barium variant was attempted – mainly to comprehensively compare original properties and unique phenomena acquired later during the investigation process from the nearest isostructural member. Also, barium turned out to be the only large-enough ion to produce good quality, double perovskite matrix.

Presented studies will have fundamental meaning and focus on synthesis descriptions and experimental details regarding used methods and related equipment. The second chapter will present an essential physicochemical characterization of undoped structures at ambient conditions using: powder XRD patterns & SEM projections, FTIR with Raman spectra, and XPS joined with complementary XAS data. Thirdly, results from non-ambient (temperature versus pressure) conditions will be shown concerning possible polymorphism (Raman spectroscopy), structural diversity (XRD, XPS, XAS), stability (in inert gases), and decomposition (air-sensitivity) using DSC and TG. Lastly, EPR measurements concerning magnetically-active content properties will open a gateway into the fourth chapter concerning optical measurements and unexpectedly decaying emission spectra of barium double perovskites related to charge transfer and interstitial oxygen evolution. Other results regarding the time-resolved photoluminescence of doped and undoped samples, excitation, absorption, theoretical and experimental band-gap form estimations are also present. Chosen dopants concern mostly Yb^{3+} and Ce or Pr^{3+} depending on discussed material – cerate-tungstates hold Pr, praseodymates Ce. All were selected in favor of creating a friendly environment for long-range, far down-converting energy transfer towards infrared and Yb^{3+} : (sensitizer Ce/Pr) $NUV \rightarrow UV-VIS$ (mediator WO_x) $\rightarrow NIR$ (activator Yb).

The final verdict about this research, based on all justifiable evidence collected so far, is that all these materials are relatively weak phosphors – barium praseodymate being the most photosensitive in matters of intensity but short-lived just as his cerate cousin; Sr and Ca-related cerium-tungstates being, on the contrary, weak but stable and persistent sources of light. They do not possess good enough capabilities to function as efficient, long-term down-converters, especially if they overheat and decompose in the air due to the extensive power applied by the lasers. They carry, however, very interesting charge-transfer capabilities related to photobleaching and internal oxygen evolution that might be at value for some NUV or deeper radiation one-time sensors - they are unfortunately unrecoverable. Use-and-forget manner makes them a one-way sensor. Nevertheless, these new matrices open many new possibilities and routes for future scientific research, i.e., in terms of magnetism or other RE dopants if they possess such shallow charge-transfer states just below the 5d conduction band. This work may continue beyond the scope of a Ph.D. dissertation or serve as an inspiration for others later on.

Streszczenie (Polish)

Praca ta w lapidarny sposób przedstawi 4 nowe materiały w postaci amfoterycznych proszków wolframianowych opartych na ziemiach rzadkich - o wzorze ogólnym $A_xRE_yW_zO_{6z}$. Dwa z nich klasyfikują się do klasycznych, podwójnych perowskitów, a dwa pozostałe to są bliżej spokrewnione z tzw. ilmenitami czy granatami. Wszystkie są uzyskiwane w wyniku wysokotemperaturowych reakcji w ciele stałym przeprowadzanych na skompresowanych pastylkach umieszczonych w atmosferze gazów obojętnych. Pierwotnym celem badań było stworzenie i dokładne opisanie owych materiałów, które powinny posiadać względnie lepsze właściwości optoelektroniczne, czy nawet scyntylacyjne w stosunku do swojego pierwowzoru jakim był $BaWO_4:Ce/Pr$. Co więcej, ze względu na unikalny polimorfizm oraz świecenie owego prekursora, istniała również duża szansa na wynalezienie efektywnego konwertera energii czy też wytworzenie radiacyjnego sensora czułego w niskich temperaturach. W związku z tym 3 z badanych związków zostały wybrane jako nośniki ceru w miejscach oznakowanych jako RE (cerany), a czwarty jako host prazeodymu (prazeodymiany). Pierwsza grupa docelowo miała należeć do rodziny podwójnych perowskitów różniących się jedynie jonami alkalicznymi ($A = Ba, Sr, Ca$)²⁺, jednak ze względu na duże różnice w promieniach jonowych, dwa ostatnie kationy okazały się tworzyć duże, pochyłe i mniej uporządkowane struktury. Są one popularnymi produktami ubocznymi w procesie produkcji ww. związków i też szeroko bada się ich naturę z różnych powodów (głównie w celach eliminacji). W przypadku prazeodymowego związku, stworzono go, gdyż chciano podjąć próbę potwierdzenia - na zasadzie kompleksowego porównania - oryginalnych właściwości ceranu baru z czymś potencjalnie izostrukturnym, co mogłoby wykazać się czymś o równie unikatowych właściwościach, Bar jest też jedynym pierwiastkiem, który okazał się być wystarczająco duży by tworzyć struktury podwójnych perowskitów zatem wybór dla kolejnego sąsiada w układzie chemicznym wydawał się być oczywisty.

Prezentowane badania mają fundamentalny charakter i pierwszy rozdział skoncentruje się głównie na szczegółowych opisach syntez i eksperymentalnych metod oraz powiązanego z nimi sprzętu. W drugim rozdziale zaprezentowana zostanie podstawowa charakterystyka fizykochemiczna struktur niedomieszkowanych w warunkach otoczenia z wykorzystaniem: proszkowej dyfrakcji XRD i zdjęć SEM. Spektroskopia FTIR zestawiona z komplementarny sposób wraz z widmami Ramana posłuży do stworzenia tzw. odcisków palca – charakterystycznych widm używanych w toku dalszych analiz strukturalnych. Dane z XPS oraz XAS ukażą równocześnie dualistyczny charakter jonowy wszystkich związków w obrębie typowych podpowłok i ich krawędzi absorpcji (głównie RE i W). Rozdział trzeci przedstawi wyniki z rozmaitych technik uzyskanych w warunkach innych niż pokojowych – głównie w funkcji temperatury oraz ciśnienia – by zbadać złożone kwestie polimorfizmu (spektroskopia Ramana), spornego zróżnicowania strukturalnego (XRD, XPS, XAS), czy stabilności (w gazach obojętnych) i ewentualnego rozkładu (wrażliwość na powietrze) przy użyciu technik DSC i TG. Ostatecznie niecodzienne wyniki pomiarowe EPR, zainspirowane istnieniem podstawowych form paramagnetycznych w związkach, otworzą drzwi ku czwartemu rozdziałowi, który dotyczyć będzie pomiarów optycznych. Tam zostanie uwypuklone zjawisko czasowo-zależnego zaniku emisji ziem rzadkich wśród barowych podwójnych perowskitów związane najpewniej z przenoszeniem ładunku do W czy międzywęzłowego tlenu pod wpływem promieniowania. Znajdują się tam również inne wyniki dotyczące porównania fotoluminescencji próbek domieszkowanych i niedomieszkowanych, absorpcji i emisji, teoretyczne i eksperymentalne przewidywania szerokości domniemych (prostych) przerw energetycznych. Wybrane domieszki to głównie Yb^{3+} i Ce lub Pr^{3+} w zależności od omawianego materiału – wolframiany cerowe zawierać będą Pr, prazeodymiany zaś Ce. Wszystkie dobrane tak by stworzyć korzystne warunki do zaistnienia zjawiska przekazu energii dalekiego zasięgu w kierunku iterbu i podczerwieni: (donory $h\nu$ Ce/Pr) NUV \rightarrow UV-VIS (pośrednik WO_x) \rightarrow NIR (akceptor $h\nu$ Yb).

W świetle wszystkich wykonanych i zebranych do tej pory wyników badań można rzec, że wynalezione materiały są stosunkowo słabymi luminoforami – praeodymian baru jest najlepszy pod względem intensywności, ale jego emisja jest nietrwała, podobnie jak w izostrukturalnym ceranie; Obydwa związki wydają się być wysoce podatne na światło (nietrwałe). Wolframiany ceru związane ze strontem czy wapniem są bardziej trwałe i długoterminowymi emiterami, ale znacznie słabszymi. Związki nie wykazują zatem wystarczających możliwości czy też zdolności, aby funkcjonować jako wydajne konwertery promieniowania, zwłaszcza jeśli mogą ulec szybkiemu przegrzaniu i rozkładowi z powodu przyłożonej dużej mocy laserowej w atmosferze tlenowej. Posiadają one jednak inne, bardzo interesujące właściwości takie jak możliwość przenoszenia ładunku powiązaną pośrednio ze zjawiskiem fotowysyblania i ewolucji wewnętrznie zgromadzonego, międzywęzłowego tlenu. Może to nieść wartościowy potencjał dla konstrukcji czujników promieniowania NUV lub głębszego (bardziej energetycznego), ale w sposób nieodwracalny – niestety są one raczej jednorazowego użytku. Niemniej jednak wszystkie wyżej wymienione matryce otwierają wiele nowych możliwości i ścieżek dla rozwoju przyszłych badań naukowych, np. w zakresie magnetyzmu lub innych z zastosowaniem różnorodnych domieszek ziem rzadkich. Ta praca może być równie dobrze kontynuowana po doktoracie lub posłużyć innym jako inspiracja do dalszego rozwoju.

Introduction and Aims

This study thoroughly describes four novel RE-tungstate materials, their crystallographic structure, physicochemical properties in various conditions, polymorphism, stability, and some basic optical phenomena related to luminescence, like charge and energy transfer. At first, undoped materials will be shown to create a base for further deliberation. Dopants like Cerium, Praseodymium, and Ytterbium 3+ will be introduced later to assess the potential for creating efficient NUV (Ce/Pr) → UV-Vis (WO_x) → NIR (Yb) down-converters or NUV sensors related to thermometry far below 0°C.

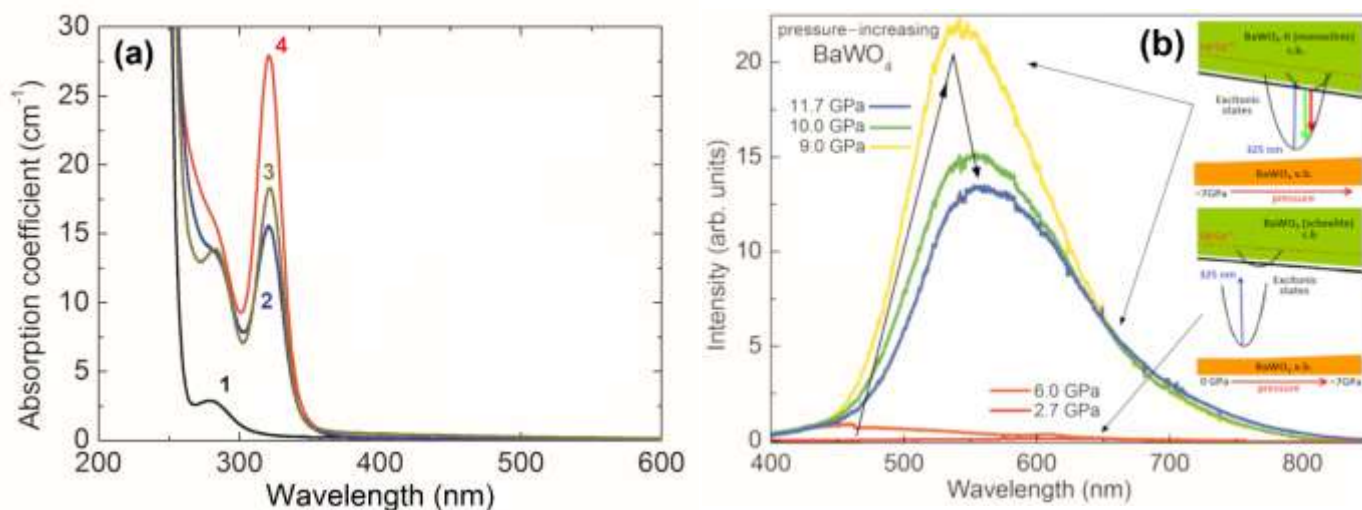


Figure 1 (a) Ambient absorption spectra of (1) pure BaWO₄, doped with 0.5 atom % Ce³⁺ along (2) [100] or (3) [001] crystallographic axes, and (4) additionally codoped with 1 atom % Na⁺ to compensate for any additional vacancies. Scheme (b) shows pressure-influenced PL properties of BaWO₄ crystals during excitation with 325 nm laser wavelength.

The research was inspired by a previous BaWO₄ study where this famous scintillator, known for elaborate polymorphism (& synthesized using the Czochralski method),¹ activates its optical capabilities during compression while being excited by NUV, 325 or even deeper 275 nm laser light – as *Figure 1*.² During that process emerging, intense, yellowish-green PL bands cover a broad visible range of emission making these crystals potentially attractive sensitizers for stimulating any artificially implanted NIR activators as shown in *Figure 2*.³

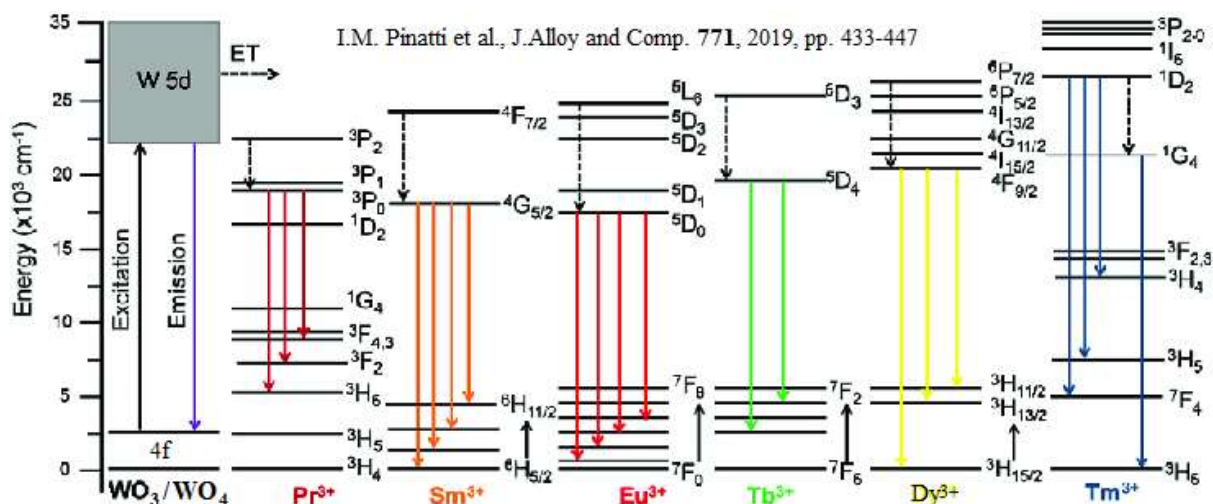


Figure 2 Diagram of an energy transfer mechanism for rare-earth doped tungstates excited with 350 nm or shorter laser light showing potential for future sensitizing ions.

On the basis of theoretical and experimental studies, several authors suggest that the origin of this greenish emission is an extrinsic mix related either to $(\text{WO}_3 + \text{F})$ centers⁴⁻⁷ or some oxygen-deficient complexes $(\text{WO}_3 + \text{V}_\text{O})$, which can emit orange light.⁸ Others consider the interference of defect centers containing interstitial oxygen $(\text{WO}_4 + \text{O}_\text{i})$ ⁹⁻¹¹ or intrinsic distortions in the $(\text{WO}_4)^{2-}$ tetrahedra as a source of the initially expected green light being artificially redshifted.¹² Theoretical studies performed by us,² Campos¹³ and Longo et al.¹⁴ attribute this PL response to various interstitial vacancies and subtle intrinsic oxygen distortions creating vast complexes within convoluted tungsten WO_x matrix. That is the most plausible answer, given the data gathered below. The activation mechanism of recorded emission initially presents itself by simple band-gap shrinkage (by lowering the top 5d W^{6+} state) while pressure is being applied - like in the pictograms shown to the right of *Figure 1b*.^{2,15,16} At those circumstances, the scheelite unit cell $(I4_1/a)$ ¹⁷⁻²¹ structurally changes at ~ 7 GPa through so-called second-order phase-transition to a metastable $I2/a$ fergusonite state,^{15,22-26} to finally form a monoclinic $P2_1/n$ phase above ~ 9 GPa,²⁷⁻³¹ as depicted in *Figure 3*.

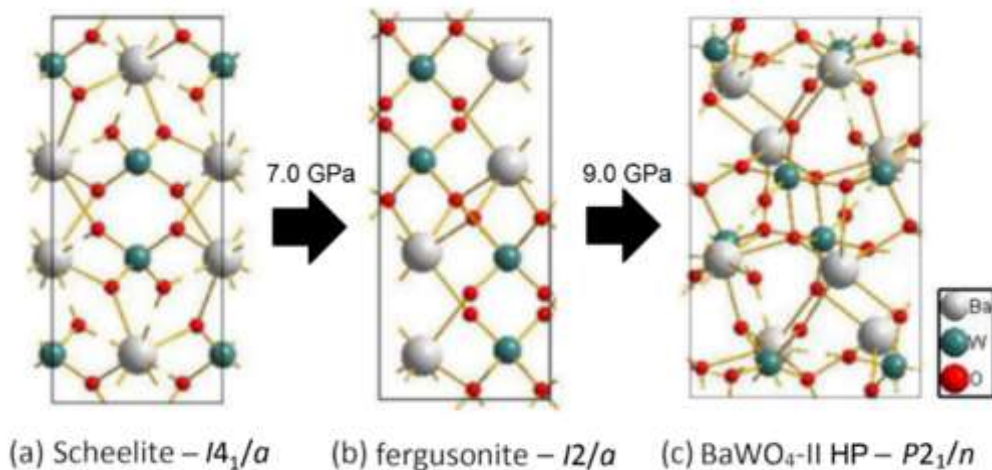


Figure 3 Three crystallographic structures of BaWO_4 : (a) ambient- and low-temperature scheelite; intermediate (7–9 GPa) fergusonite (b); high-pressure (+9 GPa) monoclinic BaWO_4 -II phase (c).

The evidence of such reversible behavior, although with large ~ 4 GPa hysteresis, was provided by Raman spectra in *Figure 4*. (Appendix *Figure A1* provides more insight about related peak versus pressure position dependencies during compression). Nevertheless, the emission of the scintillator is affected analogically – *Figure 5a,b*. The response was much more potent (regarding intensity) with a carefully chosen set of RE dopants, as seen in *Figure 5c,d*. Ideologically, Ba^{2+} should be substituted with RE ions (not W) hosting NUV sensitizers like Ce or Pr^{3+} ions, further reinforcing the PL-related effects by a couple of additional multisites. This must be done carefully since some additionally introduced (O-related) defects could hamper their intended influence. Already present, defective reddish component (band) from aforementioned V_O or O_i intrusions visualized in *Figure 6* might, however, come in handy later when exciting other NIR energy acceptors.^{2,32,33} To control their effective presence, smaller ions like Na^+ or Li^+ were introduced as co-dopants – mainly to compensate for differences in the Ba-RE ionic radii as convenient space fillers and to relieve the tension or negative internal stresses developing within matrices mentioned above.³⁴⁻³⁶

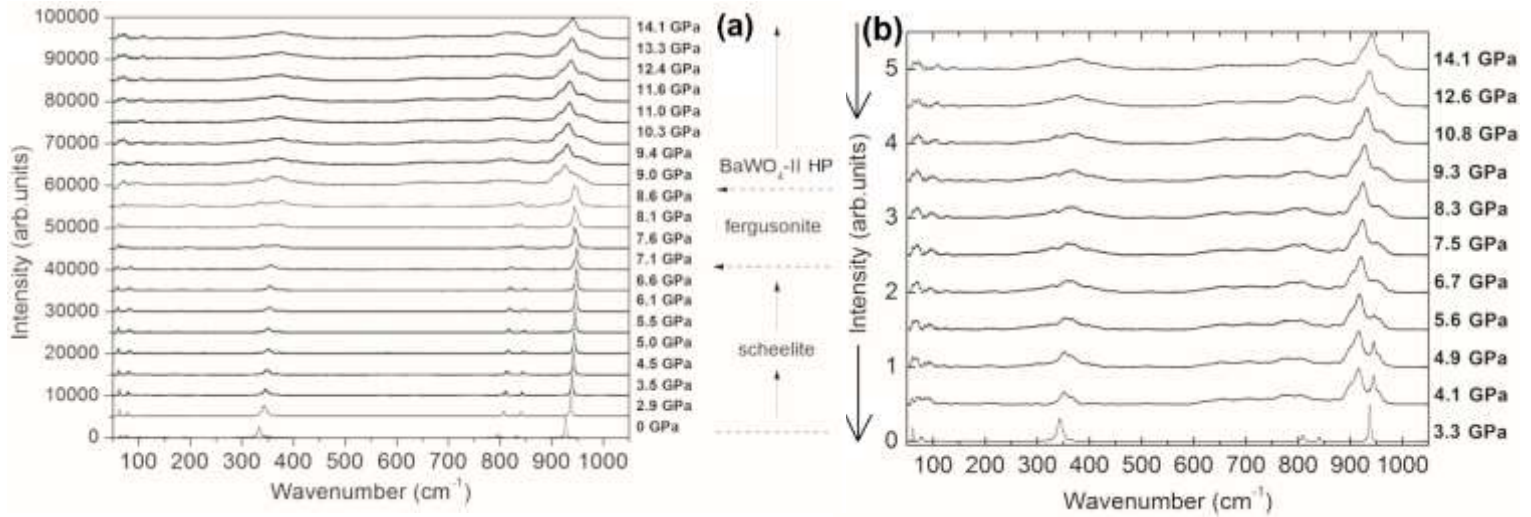


Figure 4 (a) Pressure-dependent, room-temperature Raman spectra of BaWO₄ doped with 0.5 atom % Ce³⁺ and 1 atom % Na⁺, oriented along [100] axis. Phase transitions are noticeable at 7.1 GPa (to the metastable fergusonite phase) and 9.0 GPa to monoclinic HP BaWO₄-II. Picture (b) contains data collected during pressure release. The reverse transition from the HP BaWO₄-II to the scheelite phase occurs between 4.9 and 4.1 GPa. Arrows guide the eye.

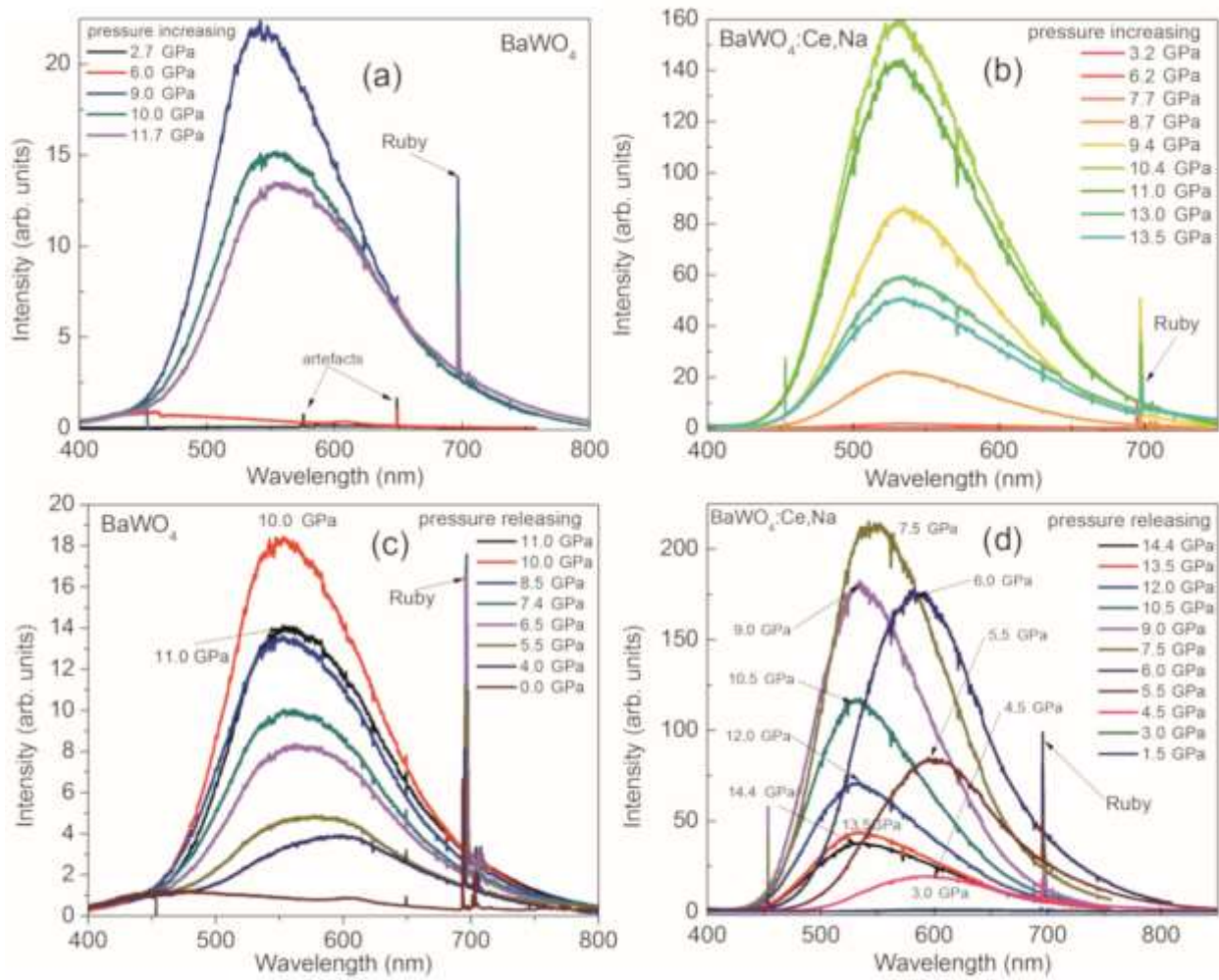


Figure 5 Luminescence of pure BaWO₄ and BaWO₄:Ce³⁺ (0.5 atom %), Na⁺ (1 atom %) as a function of pressure increase (a, b) and release (c, d) at 293 K. Excitation wavelength: 325 nm.

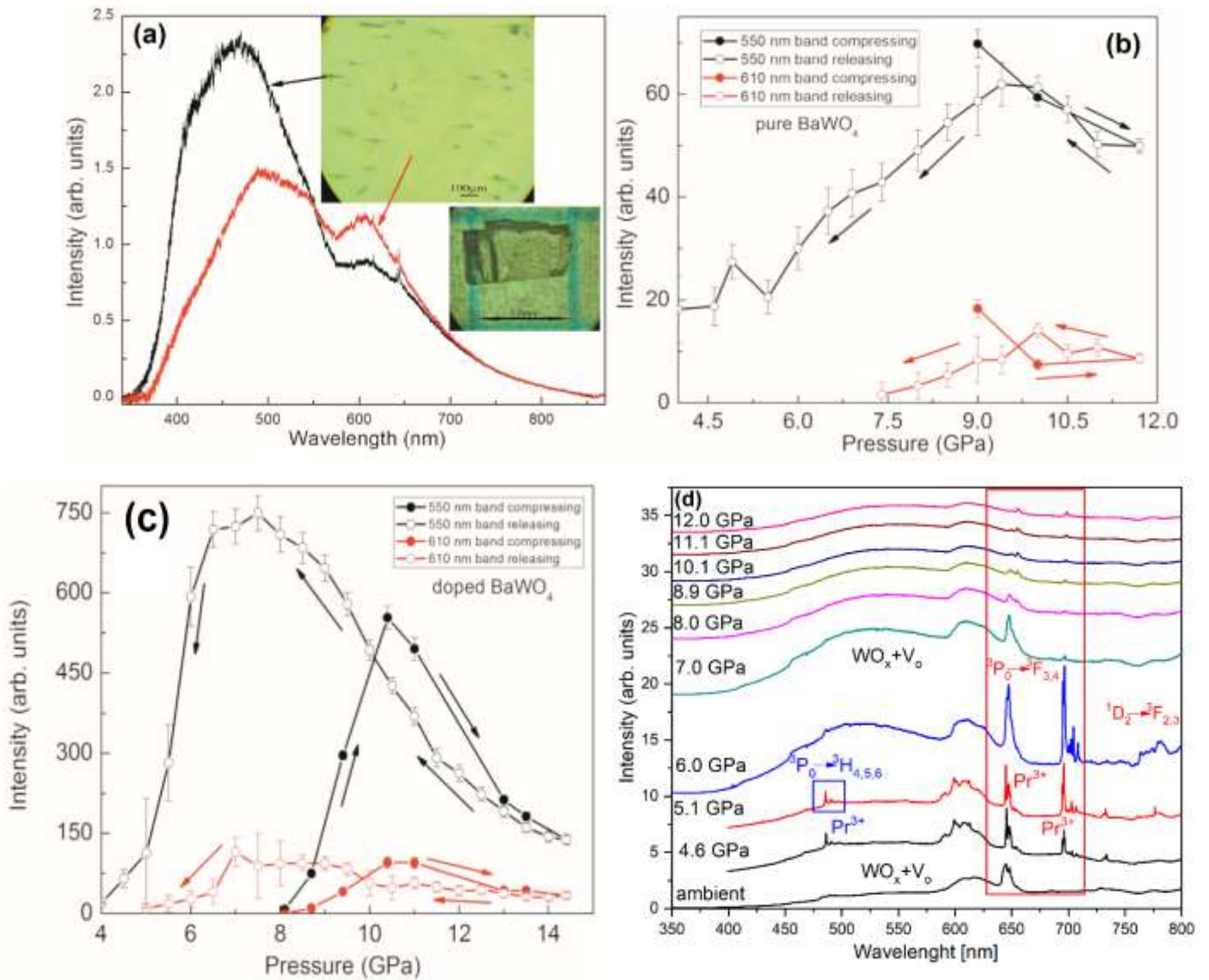


Figure 6 Luminescence of a BaWO_4 single crystal at 70 K under 275 nm excitation is shown in figure (a), where it reaches the maximum output. The red PL spectrum stems from blackish inclusions, which are actually empty regions, while the black spectrum is from a homogeneous crystal without any inclusions as visible in the insets. Intensity vs. pressure dependence of greenish-yellow 550 nm (black symbols) and weak, orange 610 nm (red symbols) PL bands are visualized in figure (b) for pure and (c) doped $\text{BaWO}_4:\text{Ce}^{3+}$ (0.5 atom %), Na^+ (1 atom %), single crystals measured at 10 K under 325 nm excitation. The direction of pressure changes is indicated using color-specific arrows. Figure (d) shows high-pressure luminescence spectra of $\text{BaWO}_4:\text{Pr}^{3+}$ (0.5%), Na^+ (1%) under 302.5 nm laser excitation. One can witness Pr and WO_x bands evolution during compression through metastable (fergusonite) phase above 6 GPa, where the emission briefly intensifies and later quenches after a brief transition to monoclinic $P2_1/n$ SG.

So, basing on initial $\text{BaWO}_4:\text{RE}$ data, aforementioned literature,¹⁻³⁵ and some theoretical projections presented below, an assumption was made. Naturally available materials and their chemical stoichiometry suggested that the formula would expand towards a double perovskite if RE were fully introduced into the matrix, not as dopants but as genuinely integrated

constituents.³⁷ Perovskites alone are a group of rock-salt type minerals discovered by Lev Perovski in late 1839 by the Ural mountain range.³⁸ Their natural occurrence in Earth's mantle was associated with primordial processes initially started in dimming stars. There, the depletion of Ti in the photosphere signified the transition from M to L-type brown dwarfs at their final stages of existence, which is also the primary ingredient of Perovskite predecessors.³⁹ Generally, they possess a straightforward, empiric formula – $X^{II}A^{VI}BX_3$, where A is a 12-times-coordinated, 2+ alkali ion. B is a 6-times-coordinated, d- or f-block cation of various charge states. X is obviously a charge-compensating anion that could vary depending on the environment or desired usage if synthesized artificially. It could be Br⁻, I⁻, Cl⁻, F⁻, S²⁻,^{40,41} but foremostly, popular O²⁻ - just like in its archetype, CaTiO₃.^{42,43} Barium titanate was later used by Megaw et al.⁴⁴ to describe the structure in detail using XRD – and so, the knopite was distinguished as a rare-earth-bearing sub-derivative bearing close resemblance to our barium compounds. Since many different cations could be embedded into those knopite-like structures, a large variety of materials could be theoretically tailored & further developed into specific applications. A popular subject of interest is nowadays: optoelectronics,³⁷ piezoelectric sensors;^{45–47} Widely considered magnetism^{48,49} associated with antiferro-^{50,51} or ferrimagnetism,^{52,53} memory-storage devices,⁵⁴ supported by superconductivity;^{55–58} photocatalysts^{59–61} used in energy conservation just like fuel-^{62–64} or solar cells^{65–67} are also at grasping range. Such broad technological scope withholds a large field of compounds ranging from metallic, semiconducting to even insulating properties using different dopants at adjustable temperatures.^{37,54,68} It would be a shame not to explore them.

Double perovskites function on the same or even broader scale. They are something in shape and form similar to a natural perovskite, but the unit cell size is usually doubled ($A_2BB'X_6$). The evenly spread, corner-shared clusters hosting central ions of various ionic radii (and type, B \neq B') surrounded by many electro-accepting ions keep them together.³⁷ Their numbers and general space group affiliation are usually dictated by several rules: ionic radii, bonding type/electronegativity associated with its coordination number, and an overall compendium of Glazer tilt systems that encompass registered distortions in a comprehensive encyclopedia.^{69,70} These systems could be further connected to one another by mutual structural dependencies based on group theory and DFT calculations suggesting possible coexistence or polymorphism if a proper stimulus is applied.^{71,72} Firstly, however, one must predict if considered formations are available at all. Several offhand methods were derived to accommodate such matters. The most common one is Victor Goldschmidt's factor from 1926.⁷³ Basing on chosen A, B, and X-site cations, their ionic radii and coordination number one can estimate (with ~75% accuracy) if the chosen material will fall into the favorable range of synthesis where a quite-ordered structure will manifest itself. Ideal theoretical values range from 0.85÷1.05 but of course, the closer one gets to the border, the more tilted structure will be, finally resulting in other, semi-related SG of a monoclinic or elongated rhombohedral kind – scientists usually strive to achieve a perfect cuboid. To compensate for subtle details, numerous discrete structural changes and dependencies now coming into view (thanks to modern techniques and statistical overview after almost a century) a more accurate (~92%) and sophisticated GS factor formula has been established by Bartel et al.⁷⁴ In theory, it should provide more information and distinguish better ordered-vs-disordered crystallographic

systems in the region of interest, but the ideology behind it is still the same. The desired acceptability values change from **0.825÷1.059** regarding (GS) to $\tau < 4.18$ (in mGS), but only if $r_A > r_B$ rule is still maintained. So, considering the general formula of our desired double perovskites to be $^{XII}A_2^{VI}BB'^{II}X_6$, where $^{XII}A = Ba, Sr \text{ or } Ca^{2+}$; $^{VI}B = RE = Ce/Pr^{3/4+}$; $^{VI}B' = W^{4/5/6+}$; and $^{II}X = O^{2-}$ the aimed structures present themselves theoretically in range of values shown in *Table 1* below.

Table 1 Classic (GS)⁷³ and modernized (mGS)⁷⁴ Goldschmidt parameters were fashioned into calculating theoretical factors predicting the existence of discussed, newly-developed materials. The ions have been paired in such a way to respect the preferable B-site (8+ or 9+) charge balance within those structures. *Red* color marks worse GS/mGS factor, *green* better values. **Bold** fonts mark the best (top) and worst (bottom) of them all.

$^{XII}A^{2+}$	Ionic radii [Å]		$^{II}X (O^{2-})$	t (GS)	τ (mGS)
	^{VI}B -site (Ce/Pr)	$^{VI}B'$ -site (W)			
Ba 1.61	Ce ⁴⁺ 0.87	W ⁴⁺ 0.66	1.35	0.989	3.421
	Ce ³⁺ 1.01	W ⁵⁺ 0.62		0.967	3.460
	Pr ⁴⁺ 0.99	W ⁴⁺ 0.66		0.962	3.474
	Pr ³⁺ 1.13	W ⁵⁺ 0.62		0.941	3.578
Sr 1.44	Ce ⁴⁺ 0.87	W ⁵⁺ 0.62	1.35	0.942	3.678
	Ce ³⁺ 1.01	W ⁶⁺ 0.60		0.916	3.829
Ca 1.34	Ce ⁴⁺ 0.87	W ⁵⁺ 0.62	1.35	0.908	3.940
	Ce ³⁺ 1.01	W ⁶⁺ 0.60		0.883	4.210

So for mGS ($= \frac{r_X}{r_{BB'}} - n_A \left[n_A - \frac{r_A}{r_B \times \ln(r_A/r_B)} \right]$), the lower the number, the better the result.

In old GS ($= \frac{r_A + r_X}{\sqrt{2}(r_{BB'} + r_X)}$), the further you go from ideal 1.00, the worse - more discrepancies in the form of alternating, non-cubic space groups might transpire.³⁷ That withholds the most truth for the lower border-line of classic GS where substitution with a smaller alkaline atom contracts the unit cell far more than B-site atoms exchange, creating, as a result, vast cationic vacancies. If they are, by chance, periodically ordered, several cases could, in proper circumstances, create an opportunity to produce compounds such as quaternary perovskites. Ideologically aforementioned materials host complex, $4a \times 4a \times 4a$ multilayered cubic (*Fm-3m*) superstructures where 'a' is denoted as a primitive cubic unit cell dimension.⁷⁵ Of course, since the GS parameter is lower in them, an enlarged number of defects changes A/B-site occupation to even below 50%. The mechanism derived from this archetype showcases a modification of most 6-times coordinated B-site octahedra to tetrahedra sites shared with chosen alkaline (A) atoms.⁷⁶ That is why most of these derivatives have been found in alternating tetragonal SGs. Furthermore, the scientific community is widely interested in them since they possess a large variety of yet unidentified capabilities as a new group of minerals. Some postulate intriguing relations between the elements of crystallographic structure and the emission spectra (concerning selectively chosen dopants & regularly placed defects that artificially modulate or even enhance the luminescence toward its persistent character); however, no concrete evidence towards a specific mechanism has been proposed yet.⁷⁷⁻⁷⁹ Fairly justifying other prosperities by their experimental results, some of them exhibit large magnetoresistivity,^{68,80} dielectricity,⁸¹⁻⁸³ superconductivity, or even antiferromagnetism in cryogenic temperatures.^{55,84} But first, let the community dispute whether or not these structures are really stable & belong to perovskites or other garnet-like SGs.

Various by-products could occur just above and below the classic GS comfort zone – from aforementioned garnets (i.e., $Mg_3Al_2Si_3O_{12}$) to a mix of unrelated salts and oxides.^{37,74,85,86} But, the most popular of them would probably be the ilmenites. They resemble a rigid titanium ore mineral with a more distorted, large, rhombohedral SG of tubular shape. Their idealized archetype would be trigonal $FeTiO_3$ with alternating layers tilted along the one (c) axis.^{87,88} It was originally discovered by William Gregor in 1791 in a river by the Russian Ilmansky mountain range.^{89,90} It could also be found (much later) on the Moon or Pluto.^{91,92} It is a widespread by-product of dying blue supergiant stars. Usually is highly anisotropic, brownish-gray, and exhibits weak magnetic properties. Due to immorphable, disordered structure riddled with many defects, ilmenites did not find many applications in today’s world – some are sparsely used in paints, inks, fabric and plastic dyes, paper, food coloring, sunscreen, and cosmetics.^{93–95} They are also used in heavy industry extreme refinement to extract Fe and Ti, but these processes seem to fade into history.^{96–99}

Table 2 $A_2BB'O_6$ materials with $A^{2+} = Ca, Sr, Ba$ composition, reported in the literature since 1950 by Vasala and Karppinen.³⁷ Left column represents the B cation, and the top row the B' cation. The green color indicates perovskite compounds successfully synthesized at ambient pressure. Purple - DP stabilized using either high-pressure / oxygen-partial-pressure conditions. Compounds with hexagonal non-perovskite structures like ilmenites or garnets are yellow. Cases where no coherent structures or phases related to the $A_2BB'O_6$ composition formed are marked as red. White boxes are yet uninvestigated options. X marks with possible color outcomes (according to the previous GS discussion) mark the spots of the work currently conducted work in this dissertation.

Ba	22	23	25	41	42	44	50	51	52	53	73	74	75	76	77	78	83	92	93	94
	Ti	V	Mn	Nb	Mo	Ru	Sa	Sb	Te	I	Ta	W	Re	Os	Ir	Pt	Bi	U	Np	Pu
3 Li																				
11 Na																				
12 Mg																				
13 Al																				
19 K																				
20 Ca																				
21 Sc																				
22 Ti																				
23 V																				
24 Cr																				
25 Mn																				
26 Fe																				
27 Co																				
28 Ni																				
29 Cu																				
30 Zn																				
31 Ga																				
32 Ge																				
38 Sr																				
39 Y																				
40 Zr																				
45 Rh																				
47 Ag																				
48 Cd																				
49 In																				
56 Ba																				
57 La																				
58 Ce																				
59 Pr																				
60 Nd																				
62 Sm																				
63 Eu																				
64 Gd																				
65 Tb																				
66 Dy																				
67 Ho																				
68 Er																				
69 Tm																				
70 Yb																				
71 Lu																				
72 Hf																				
80 Hg																				
81 Tl																				
82 Pb																				
83 Bi																				

Sr	22	23	25	27	40	41	42	44	51	52	53	72	73	74	75	76	77	83	92	
	Ti	V	Mn	Co	Zr	Nb	Mo	Ru	Sb	Te	I	Hf	Ta	W	Re	Os	Ir	Bi	U	
3 Li																				
11 Na																				
12 Mg																				
13 Al																				
20 Ca																				
21 Sc																				
22 Ti																				
23 V																				
24 Cr																				
25 Mn																				
26 Fe																				
27 Co																				
28 Ni																				
29 Cu																				
30 Zn																				
31 Ga																				
38 Sr																				
39 Y																				
40 Zr																				
45 Rh																				
48 Cd																				
49 In																				
50 Sn																				
57 La																				
58 Ce																				
59 Pr																				
60 Nd																				
62 Sm																				
63 Eu																				
64 Gd																				
65 Tb																				
66 Dy																				
67 Ho																				
68 Er																				
69 Tm																				
70 Yb																				
71 Lu																				
72 Hf																				
73 Ta																				
77 Ir																				
82 Pb																				
83 Bi																				

Ca	22	23	41	42	44	51	52	72	73	74	75	76	77	83	92
	Ti	V	Nb	Mo	Ru	Sb	Te	Hf	Ta	W	Re	Os	Ir	Bi	U
3	Li														
11	Na														
12	Mg														
13	Al														
14	Si														
20	Ca														
21	Sc														
22	Ti														
23	V														
24	Cr														
25	Mn														
26	Fe														
27	Co														
28	Ni														
29	Cu														
30	Zn														
31	Ga														
39	Y														
40	Zr														
46	Pd														
48	Cd														
49	In														
57	La														
58	Ce														
59	Pr														
60	Nd														
62	Sm														
63	Eu														
64	Gd														
65	Tb														
66	Dy														
67	Ho														
68	Er														
69	Tm														
70	Yb														
71	Lu														
81	Bi														

Given all the by-products that can transpire during sole double perovskite synthesis while GS factor lowers down, a brief statistical review of available space groups^{37,41,71,72} and Glazer tilt systems⁶⁹ is required to understand the current product placement of newly created materials as shown in *Table 2*. There, a summary of all double perovskite creation attempts is gathered to the best knowledge since late 1950 to the year 2015. Amongst all the green and violet squares, meaning successful and HP-aided syntheses, and yellow with red marks, suggesting the creation of other phases (such as ilmenites) or total failures in the form of mixed, inhomogeneous salts and oxides, black X's are placed. These were the indigenous spots that this research is trying to fulfill. Ba, Sr, and Ca (double) perovskites of Ce/W or Pr/W. The actual results will be discussed in Research and Discussion *chapter 2*, however, at this rate, one can estimate, by calculated GS, that their outcome would be as suggested by colors provided in the insets near inner X edges. Barium Ce, Pr/W compilation having $GS > 0.94$ should appear as relatively good DPs. Strontium Ce/W pairing is still above 0.90, so we could assume some type of order in the form of perovskites – maybe quaternary or more disordered shapes. Calcium is relatively small

and operates at the lower borderline of the tolerance factor, where statistics are not that promising. In terms of sheer numbers, they could turn out as ilmenites or, at worst, a mix of oxides and tungstate salts. Nevertheless, omitting statistical discouragement presented in *Figure 7*, one should still aim for the best outcome. BCW and BPW are relatively close to the ideal $GS = 1.00$ and are in the most packed, SG-rich range. Here cubic ($Fm-3m$), rhombohedral ($R-3$), tetragonal ($I4/m$), and monoclinic groups transpire – a little less $I2/m$ or more tilted $P2_1/n$. Since shrinkage of the unit cells (lowering GS) promotes $P2_1/n$ distortion and inflation of cubical $Fm-3m$ SG, all future projections should be focused solely on the remaining rhombohedral, tetragonal, and monoclinic groups. The lower we go (GS-wise) towards SCWO and CCWO, the monoclinic systems seem to be more prevalent. While $I2/m$ is visually set in the background, the chance for a cubic unit cell does not fade as rapidly as for the other SG. The pattern is, however, rather distinct and not likely. CCWO has clear dominance over $P2_1/n$ SG chance but with such low GS, materials could easily turn into other disordered systems falling out of double perovskites charts - *Figure 8*.

A wider scope of DP SG projection parameters should be considered, not solely based on simple GS values. Mainly speaking, the difference between inserted BB' -site ionic radii ($\Delta r_{BB'}$) and cationic charge states ($\Delta Z_{BB'}$) which have further connections to such factors like: Degree of order (S) being defined as $2g_{BB'} - 1$,¹⁰⁰⁻¹⁰² where $g_{BB'}$ is the occupancy declared by XRD measurements at B-site cations (correctly occupying designated sites without any sharing). (S=1 means complete order; 0 long-range disorder); and generally known as

Madelung's Free Energy.^{103,104} Madelung energy of ordering is an empirical value (not only for perovskites), which states that with the increasing charge difference and change in central ions ionic radii, the electrostatic repulsion should overcome the entropic issues inside any unit cell and cations should order properly given the exact amount of free interatomic space.¹⁰³⁻¹⁰⁵ This parameter is also proportional to $\sim \Delta Z_{BB'}$, and $\Delta Z_{BB'}$, in consequence, imposes a direct dependence on $\Delta r_{BB'}$. The larger the ΔZ , the better internal order one will get. To have a proper perspective, if $\Delta Z=0$, it is highly inconvenient. This value upsets the stereotypical 8+ balance of DPs & there is no energy gain between cations. They repulse each other pretty strongly, prompting higher disorder.^{106,107} Ionic radius is the only directly-related parameter that could strongly correlate with changing cationic charge/structure imbalance. While changing, it behaves in a proportional manner when quantified with some structure-specific constants. Generally, the larger the difference between central B & B' sites, the more probable the long-range ordering is.^{107,108} According to provided statements, the $\Delta r_{BB'}$ difference higher than 17% (or rather 20% given the statistics) would increase lattice strain that, in consequence, would overcome entropic influence, and the structure would come out genuinely ordered. *Table 3* contains all possible options and annotations from literature gathered by Vasala & Karppinen, with red marks pinpointing the range of existence for our newly-developed materials.

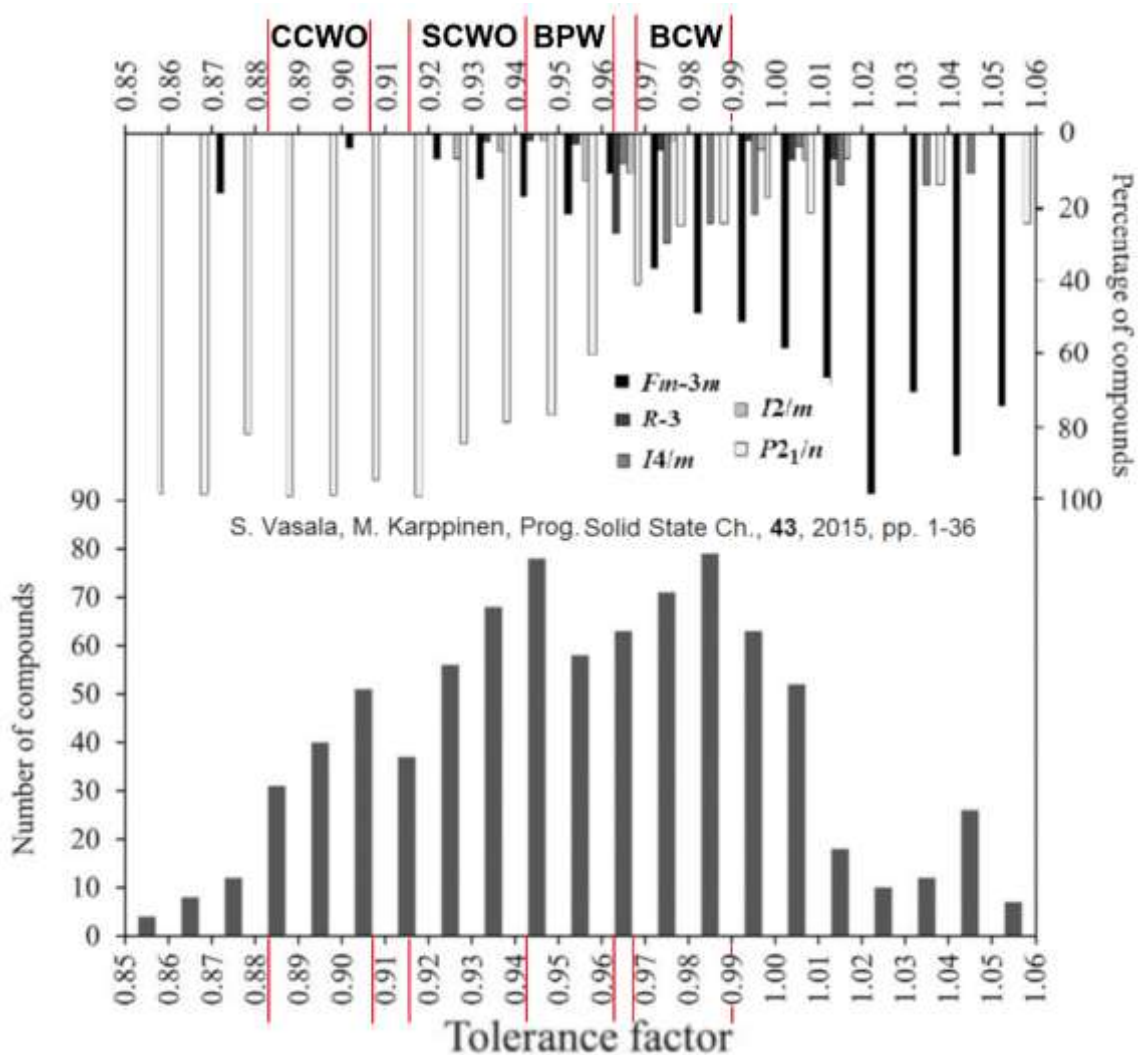


Figure 7 Number of $A_2BB'O_6$ compounds reported with different values of GS factor and percentages of ordered compounds hosting different space groups merged together from graphs presented in Vasala's & Karppinen's review.³⁷ Red color lines mark the borderline values of each investigated material, taking into account different B-site ions mentioned in *Table 1*.

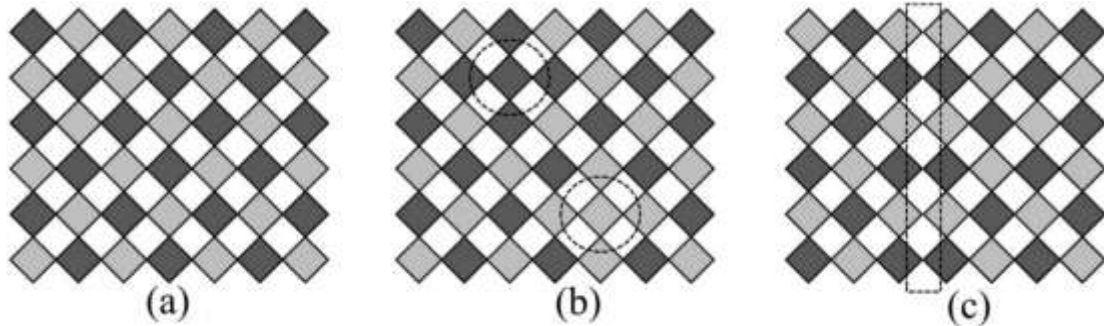


Figure 8 Schemes of probable B-site cation order in perovskites: (a) completely ordered, (b) antisite disorder, and (c) the antiphase boundary case.³⁷

Table 3 Summary of cation orderings found in different $A_2BB'O_6$ double perovskites made by Vasala & Karppinen.³⁷ Red color marks ideologically desired ionic structures.

$\Delta Z_{BB'}$	Ionic structure	Ordering	$\Delta r_{BB'}$ [\AA]
0	$A^{2+}_2B^{4+}B'^{4+}O_6$	Disordered with $\Delta r_{BB'} < 0.17 \text{ \AA}$	$\Delta r_{CeW}=0.21$; $\Delta r_{PrW}=0.33$
	$A^{3+}_2B^{3+}B'^{3+}O_6$	Disordered	
	$A^+_2B^{4+}B'^{6+}O_6$	Ordered	
2	$A^{2+}_2B^{3+}B'^{5+}O_6$	Mostly disordered, ordered if $\Delta r_{BB'} \gg 0.17 \text{ \AA}$	$\Delta r_{CeW}=0.39$; $\Delta r_{PrW}=0.51$
	$A^{3+}_2B^{2+}B'^{4+}O_6$	Most partially ordered	
4	$A^{2+}_2B^{2+}B'^{6+}O_6$	Mostly highly ordered	
	$A^{3+}_2B^+B'^{5+}O_6$	Ordered	
6	$A^{2+}_2B^+B'^{7+}O_6$	Ordered	

Researchers were fully aware of unfavorable (low) $\Delta Z_{BB'}$ values making their choices during pure-lattice modeling, as seen in *Table 3*. However, future experimentations showed that high enough $\Delta r_{BB'}$ with additional inclusions hosting mixed $\Delta r_{Ce^{4+}/W^{5+}} = 0.25 \text{ \AA}$ and $\Delta r_{Ce^{3+}/W^{6+}} = 0.41 \text{ \AA}$ pairs (appearing spontaneously) hold great promise in terms of autoregulation. That phenomenon bestows more promising GS factors for the overall matrix but is challenging to control. That is probably why no one has yet divulged into this group of compounds considering problematic synthesis and its incoming, meticulous optimization/control.

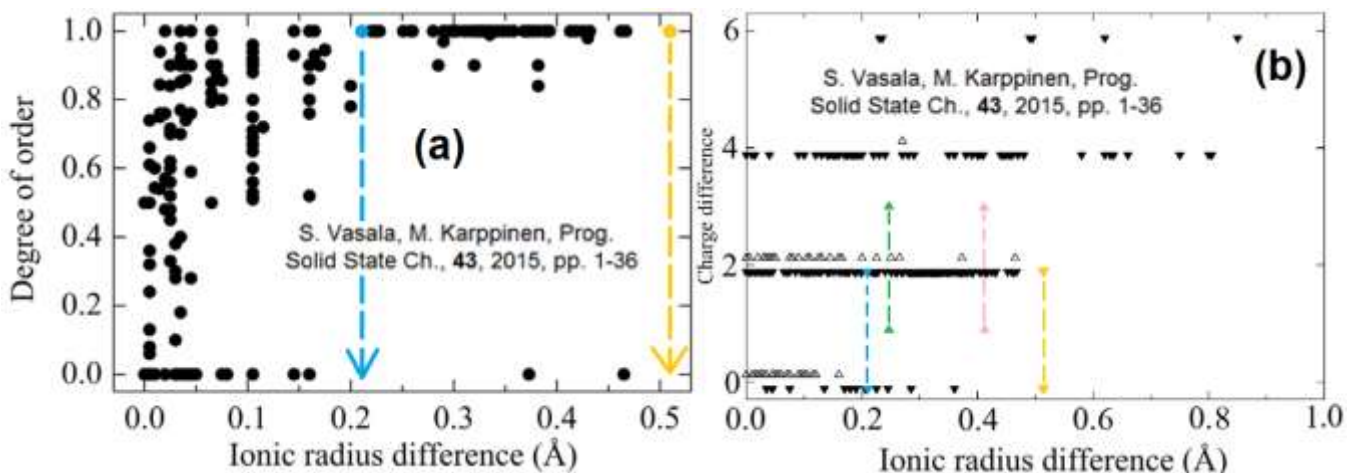


Figure 9 Statistical overview of (a) B-site cation's degree order for $A^{2+}_2B^{3+}B'^{5+}O_6$ perovskites, and (b) charge difference as a function of B-site cation radius divergence made by Vasala and Karppinen.³⁷ In Figure (b), disordered compounds (hollow triangles) have been moved up, and ordered (black triangles) slightly lowered for clarity/legibility reasons. Blue and yellow symbols with dashed arrows guide the eye in the range of possible placement of our newly-discovered materials – their stretch depends on Ce^{3+}/W^{5+} and Ce^{4+}/W^{4+} fractions/ ratios of intermediate pairing. Green and pink marks account for unusual, mixed ion couples that might transpire between the two sets mentioned above, possibly creating other disordered compounds like ilmenites. The 'pairing' refers to the batch's whole (scattered) volume, not the immediate crystallographic neighborhood.

Knowing all three Madelung parameters, one can now crudely project newly modeled materials into the compiled statistical charts montaged by Vasala and Karppinen³⁷ in *Figure 9*. There, some ideal, borderline points of order have been displayed according to currently chosen, 100% pure cationic options (Ce blue and Pr yellow). A broader range should be set regarding possible intermediate outcomes emerging from the chaotically mixed charge states (green Ce and Pr pink) due to the unideal dispersion of ions naturally upsetting the perfect balance. This uncertainty might be narrowed down by carefully-planned XRD, NPD, XPS, or XAS experiments; however, considering synthesis conditions, everything needs to be perfectly repetitive and stable, which is often not the case.³⁷ Charge disproportionation^{109–113} and A/B-site mixing^{114–117} are phenomena that are not easily controlled but might lead to some exotic oxidation states or properties that are not conventionally available. However, discretely balanced and stable ions inside a unique matrix are something to strive for. Compared to cation ordering, the charge balancing process is quite often quick & chaotic in rock-salt type materials – many factors can upset it. Sometimes fast cooling (freezing) is required to prevent minor phases from appearing in more significant amounts – kinetic barriers often limit such $A_2BB'O_6$ perovskites from appearing (influencing their $ABO_3 \times AB'O_3$ cluster fusion).^{118–121} This procedure often leads to discrete, temporary-frozen B-site balance (sometimes uneven) but is necessary to maintain the delicate 8+ charge state of internal DP ions. This procedure is, unfortunately, not typically recommended since it counters slow (but crucial) ionic diffusion during solid-state reaction & has a bad influence on the overall mix - even a small anion displacement can leave lattice strains & empty spaces for interstitial inclusions (like O). These phenomena would usually lead to later oxidation & facilitate unnecessary electron transfer from unknown sources. Some say that freezing is cheating nature by compensating the difference in space with artificial defects – the material should be freely self-balancing from a long-range perspective. Electron movement should not be obstructed then, but it is a highly disputed subject here.^{53,71,82,101,102} Schools of synthesis are different because charge ordering is not always complete, and ideal chemical formulae are not ideal integers – usually, fractions with a surplus of anions are added to meet the needs of many just to have satisfying purity.³⁷ This ideology is highly dependent on synthesis temperature, as one will see in upcoming experiments in the air or our syntheses' based on this 'so-called' rapid cool-down.^{122–125} High-pressure, of course, plays a role in it too but is not as prominent – more in terms of moderating such properties like excessive evaporation and final SG duality (RT polymorphism).^{126–128} The last subject is a fight between 2 competing processes associated with mixed valency^{105,110} and partial substitution (A/B-site crossover), as shown in *Figure 10*. That causes a lot of mayhem but also the emergence of interesting properties in terms of electric conductivity.^{56,129} Mixing of A/B states is slowed chiefly down but still transpires at different rates at each occupation site. Trans ionic migration is not commonly desired since it is the primary perpetrator of minor SG inclusions or even total failure in achieving a clean, homogenous product by breaking DP into oxides and salts/garnets – as portrayed in *Table 2* by all visible red spots. However, as one will see in matters of A=Ca, Sr, and B=Ce, even with such processes, compounds such as quaternary perovskites or ilmenites could still commonly appear.

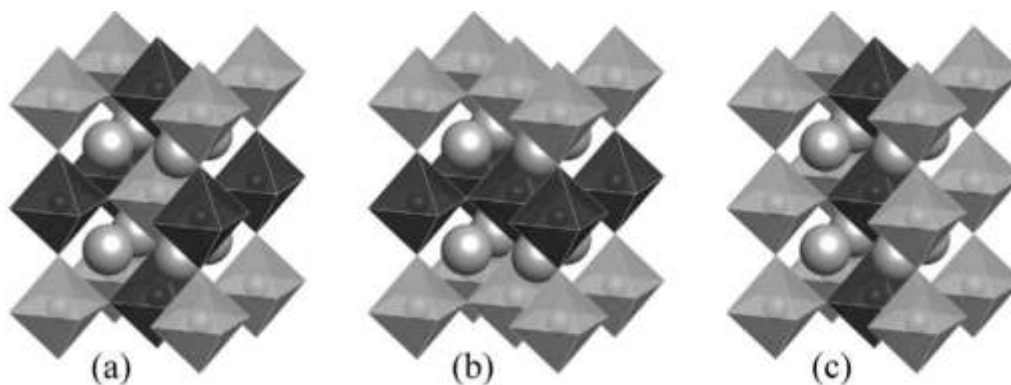


Figure 10 Different A/B-site cation orderings found in $A_2BB'O_6$ DPs: (a) rock-salt, (b) layered, and (c) column types as foretold by Vasala and Karppinen.³⁷

Of course, other perovskite systems exist, like so-called triple-deckers. Very useful, layered heterostructures of semiconductors (inorganics) enclosed by atomically thin organic (mainly insulating) sheets usually grown from solid solutions and withheld by van der Waals forces could be tailored into a wide range of optoelectronic applications.¹³⁰ Inspired by the discovery of 2D graphene and efficient encapsulation, those crystalline solids could also be used for energy storage and IT technologies. That is, only if interfaces are neatly stacked & well-defined thin films, which is a time-consuming and tedious procedure to maintain.¹³¹ By contrast, layered perovskites could also be seen as analogs of weakly bonded 3D lattices stabilized by organic molecules in Pb-based solar cells.^{132,133} Authors like Aubrey¹³⁴ have already noticed that, with current (organic perovskite) fabrication technology, this family of compounds could be exploited as a template for modeling white light emission with control of electrical conduction on a macro scale. Their optical experiments and computational modeling have also shown nice quantum effects worth further pursuit in ‘topological’ electronic states leading to spatially dependent conductivity.^{130,133}

Our team worked once on substructures derived from triple-deckers. So-called Ruddlesden-Popper matrices - atomically thin inorganic sodium titanate 2D lattices co-assembled from different $Ln = La^{3+}$ ($x=0.005$), Gd^{3+} ($x=$) ions doped with Pr^{3+} ($Na_2Ln_{2-x}Pr_xTi_3O_{10}$) for self-calibrated, low-temperature, optical thermometry.^{135,136} This project has proven to be a valuable experience regarding synthesis regimes and foreseeable Pr radiative transitions. Despite my work being mainly restricted to high-pressure polymorphism and Raman spectra (peak vs. pressure/temperature relations – see *Figure A2* for reference) one could witness greenish-blue (${}^3P_0 \rightarrow {}^3H_5/{}^3F_2$) and, more important, orangish-red light (${}^3P_0 \rightarrow {}^3F_4$ which could be used towards Yb^{3+} NIR excitation. Behaviorism such as this during compression served as another inspiration for this work.¹³⁷ The idea of incorporating Pr^{3+} as an efficient transmitter was established having in mind the advantage of non-radiative ${}^3P_0 \rightarrow {}^1D_2$ overpass and controlled luminescence quenching via shallow traps (vacancies) – a possible transition mechanism suggested by Aumüller et al.¹³⁸ Nevertheless, those layered perovskite-oxide compounds, expressed by a general $A_2A'_{n-1}B_nO_{3n+1}$ formula, where the perovskite $A'_{n-1}B_nO_{3n-1}$ two-dimensional, micro-spaced decks possess octahedra formed by slicing typical, perovskite structures alongside one of orthogonal, cubic directions interleaved by A cations as shown in *Figure 11* were too complicated for commercial implementation, so I have decided to turn my sight to something simpler (maybe more profitable).^{139,140}

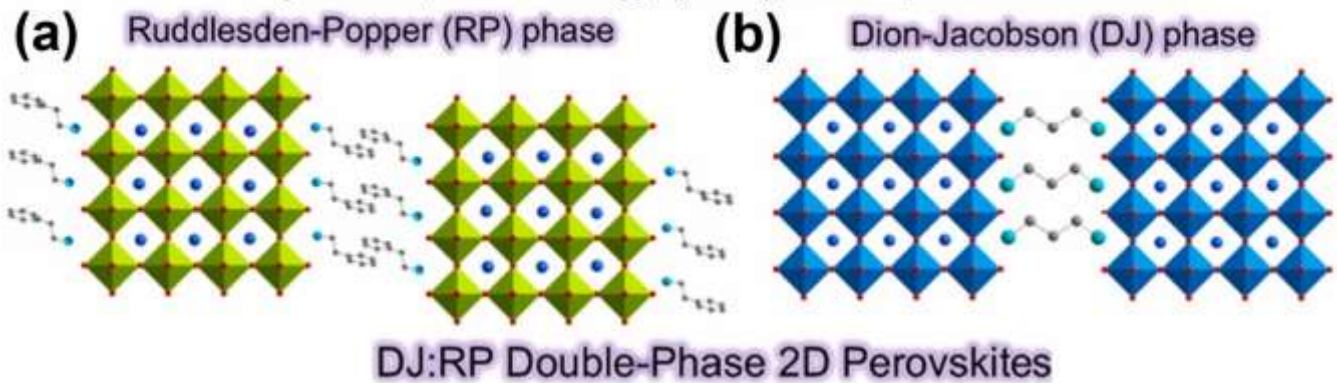


Figure 11 Structural manifestation of 2D-3D triple-decker perovskites phases – Ruddlesden-Popper (a), and Dion-Jacobson (b) as described by Ping Fu et al.¹⁴¹

Worth mentioning are the other Dion-Jacobson phase perovskites. They have also been proven to be more stable than 3D-based solar cells. These double-phased perovskites (not double perovskites) are composed of diammonium matrices often mixed with large organic ions (interlayered slabs), preventing local relaxation.¹⁴¹ Compared to mono ammonium counterparts, they suppress charge recombination and e^- transport (charge transfer) by improving energy converters' efficiency and stability, introducing some local distortions into the crystallographic structure.^{141–143} Those robust, harsh matrices have stressed lattice crucial in controlling cation migration by passive defects and distortions in 3D space. That is why those materials are considered mixed 2D:3D type cells, while the 2D aspect is rather crucial.^{144–146} Engineering such substances is rather complex and involves nucleation of the Ruddlesden-Popper phase first as an archetype due to lower formation energy. However, they possess immense potential in constructing stable, futuristic solar panels if only the designers choose proper spacers to effectively control thermal-related stresses in environmentally changing relative humidity.^{132,147} Now, one can see why some people (like me) decided to fast-cool (freeze) their materials during the last stages of solid-state reaction.

Research & Discussion

1. Experimental Section

1.1 Equipment & Instrumentation

Powder X-ray diffraction (XRD) data were collected on an X'pert MPD (Panalytical) diffractometer equipped with a Cu ($K\alpha$) lamp (40 kV, 30 mA) involving Johansson type germanium monochromator and semiconducting strip-detector of resolution close to 0.05° . Broad patterns (from 10 up to $150^\circ 2\theta$) were collected within a 3 to 12h window with setup working in Bragg-Brentano geometry. Rietveld refinement was performed on FullProf Suite.¹⁴⁸

Confocal micro-Raman spectra were registered on Monovista CRS+ spectrometer provided by S&I GmbH having: 0.75 m Acton-Princeton monochromator; back-thinned, deep-depleted PyLoN system, nitrogen-cooled (-125°C) CCD camera ($1340\times 100 - 20\mu\text{m}$ per pixel array); computer-controlled Olympus XYZ IX71 inverted stage equipped with Moticam (1280×1204) camera and long working distance objectives such as $50\times$ (NA 0.9, ambient conditions), $10\times$ (NA 0.23, high-pressures – HP, & temperatures - HT) and $5\times$ (NA 0.12, low temperatures - LT) magnifications were used during various studies. Used lasers were grey-filtered (power managed by automated Trivista software). Green, diode-pumped Cobolt Samba 04-01 series emitting 532 nm wavelength was used in junction with 2400 grooves per mm holographic grating in monochromator; and NIR Torsana StarBright L series 785nm laser was used together with 1500 grooves per mm grating. After passing through a 50/50 fused silica broadband beam splitter, the Rayleigh line was low-pass filtered from the Stokes signal up to $\sim 60\text{ cm}^{-1}$. The monochromator entrance slit had $100\mu\text{m}$, and the final resolution was 0.88 cm^{-1} for green, and 0.49 cm^{-1} for NIR laser, respectively. The acquisition time was growing from 40 min to 2h since both samples strongly absorbed the light and overheated quickly – in order not to burn them in varying gas and p(T) conditions, all samples' power density was set within the range of $0.217\div 1.12\text{ mW}/\mu\text{m}^2$.

Complementary Fourier Transform Infrared (FTIR) reflection spectra were collected using Nicolet 8700 spectrometer from ThermoElectron Corp working in Attenuated Total Reflectance mode. 64 scans, 1 s each, were registered between $400\text{-}4500\text{ cm}^{-1}$ with a resolution set of $\sim 1\text{ cm}^{-1}$.

Electron Paramagnetic Resonance (EPR) data was acquired via a Dual channel Bruker ESR300 spectrometer working on X band (9.33 GHz), with a precision of around 100 mG between 1 to 8.6 kGausses. Powders were placed in 4 mm thin wall quartz tubes of 250 mm length provided by WilmadLabGlass through Merck SA.

Synchrotron radiation experiments – X-ray Absorption Spectroscopy (XAS) - were performed on pellets at the SOLARIS NSRC facility in Poland, Cracow. The bending magnet (1.31 T) at the BL04 beamline used Total Electron Yield detection (TEY) mode at RT & Silicon Drift Detector (SDD) at LT. The synchrotron worked at electron energy of 1.50 GeV. An ultra-high vacuum of $\sim 1.5\times 10^{-9}$ mbar was applied. The beam spot size was roughly $\phi=2.5$ mm. No specific polarization was used. The resolution was set in between 75 (Ca $L_{2,3}$ -edge, O K-edge) to 150 meV (Ba, Ce $M_{4,5}$ -edges), depending on the choice of specific edge. The total working range was $150\div 1700\text{ eV}$. Samples in the form of compressed powder pellets

were placed on Mo holders & tighten up using screws with pins. Background measurements on empty holders were also performed.

X-ray photoelectron spectroscopy (XPS) measurements were performed after XAS by applying powders to Prevac setup equipped with Scienta R4000 hemispherical analyzer (pass energy 200 eV) and monochromatic X-ray tube (Al $K\alpha$ - 1486.7 eV source). The Full width at half maximum (FWHM) of the reference $4f_{7/2}$ Au line measured at room condition was 0.6 eV. The orbitals O 1s (K), W 4f ($N_{6,7}$), Ce 3d ($M_{4,5}$), Ba 3d ($M_{4,5}$) or Ca 2p ($L_{2,3}$), spin-orbit doublets were measured. Spectra were analyzed using the commercial CASA XPS software package (Casa Software Ltd, version 2.3.17) using Shirley background curves. Spectra were fitted with a mixed Gaussian-Lorentzian function. Mounting proper electrical contact between low-conducting powder and the instrument was ensured by the incorporation of carbon tape.

Differential Scanning Calorimetry (DSC) & Heat Capacity (C_p) were taken on Netzsch Phoenix DSC apparatus, model DSC 204 F1. Temperature scans ranged from room conditions to 873K at a heating rate of 10K per min. A three-staged, heating-cooling-heating system was set in inert-inert-air conditions to assess the particular behavior of each sample in specific environments. The heat capacity (C_p) reference was Sapphire. Crucibles hosting investigated powders were made out of concave Al pans.

Thermogravimetry (TG) was subsequently carried out in nitrogen and air conditions - 20 ml per min of nitrogen flow – using Netzsch Tarsus, model TG 209 F3 at temperatures ranging from 293K to 1273K. The heating/cooling rate was roughly 10K per min. The mass of each powdery sample placed in a small corundum crucible had approximately 10 mg.

A Hitachi SU-70 SFE scanning electron microscope with a dispersive radiation detector (EDX) and cathodoluminescence system GATAN Mono CL3 was used to take SEM images. Powders' images were additionally processed for statistical purposes using ImageJ software.

High-pressure (HP) measurements were performed up to 20 GPa via hydraulic drive in Almax easyLab diamond anvil cell (DAC) using diamonds with 0.45 mm II-as standard design culets. Gaskets with 0.15 mm holes were made from Inconel x750 alloy and filled up to 50% with powder. The applied pressure transmitting medium (PTM) was argon, and ruby (α - $Al_2O_3:Cr^{3+}$) was used as a pressure gauge applying the principles divulged by Syassen and Mao et al.^{149,150} An exemplary compression cascade showcasing R1/R2 modes used for calculations is shown in *Figure A3*.

Low temperatures (LT ~5K) were maintained using continuous helium-flow, Oxford Instruments cryostat model CF1204. High-temperature measurements (up to 873K) were performed in an N_2 and air atmosphere using the LINKAM FTIR600 stage working together with the T95 controller and LNP95 water pump operated via PC LINK software & Pt103 thermocouple. Both instruments were equipped with NUV-permeable quartz windows.

UV-Vis-NIR absorption & reflection spectra were collected in the 250-900 nm range using Cary 5000 spectrophotometer equipped with a photomultiplier, a PbS detector, and a Quanta-phi integrating sphere attached to measure powders and pellets. The sphere is coated with Spectralon® 95% reflectance material, 2.5 μm thick, coupled with a double Littrow monochromator, Schwarzschild optics, and tungsten, halogen, or deuterium arc source. Experiments were executable in a photometric system working in double-beam mode operated by WinUV software.

Emission and excitation spectra were measured on Horiba Jobin-Yvon Fluorolog-3 spectrofluorometer working in FL-3-12 configuration with FluorEssence software. 450 W Xenon

lamp, joined with double (emission) or single (excitation) grating monochromator, works in tandem with two possible detectors: single-photon counting mode PMT Hamamatsu R928P (operational range 260-850 nm) or PbS for NIR range.

The EKSPA photoluminescence detection system known as an optical parametric oscillator (OPO) excitation source joined with a SpectraPro SP-2500 spectrometer provided by Princeton Instruments was used to separate incoming UV light for dopant, LT excitation wavelengths: from 212 nm, through 232, 252, 275, 325 to 360 and 431 nm. The multichannel detector head provided by Hamamatsu Photonics was controlled by MCD controller c7557-01. Slit alternated between 1 (doped) and 1.5 (undoped) mm while acquisition time varied from 2×5 (undoped) to 10×0.5 (doped) seconds. If any, the commonly used filter was only low-pass 400 nm.

1.2 Synthesis conditions and reactants

After meticulous optimization, substrates chosen for solid-state synthesis are summarized in *Table 4*. Process specifics are crudely summarized in an infographic labeled as *Figure 12*, however, the whole proceedings and remarks regarding the research and development of those paths will be the subject of discussion in this paragraph as well as all of the tested options and disregarded intermediate reactants or solutions.

Table 4 Summary and characteristics of substrates used in solid-state syntheses with their respectable providers, purity, and other valuable remarks regarding their proper safe-keeping.

Substrate	Product	Purity [%]	Provider	Remarks
BaCO ₃	BCW & BPW	99.99	Strem Chemicals	Preheated to release H ₂ O
CaCO ₃	CCWO	99.95		
SrCO ₃	SCWO	99.94		
CeO ₂	BCW, BPW	99.995	Strem Chemicals	CCWO & SCWO failed the test Preheated to release H ₂ O
Ce ₂ (CO ₃) ₃ × n (6) H ₂ O	CCWO, SCWO	99		
(NH ₄) ₂ Ce(NO ₃) ₆	BCW	99.99		
WO ₂	BCW, BPW	99.9	Merck	Positively tested CCWO & SCWO tests failed
WO _{3-x} (W ₁₈ O ₄₉)	BCW	94	Self-made	Positively tested Unstable
WO ₃	CCWO, SCWO	99.995	Strem Chemicals	Pr ₃ O ₅ ~6%
Pr ₆ O ₁₁	BPW	99.99		
Pr ₂ O ₃	All doped samples	99.9	Chemat/Alfa Aesar	Unstable in air
Yb ₂ O ₃			Strem Chemicals	

The first syntheses were done using Pechini,¹⁵¹ hydro-thermal & solvothermal methods¹⁵² since they are much cleaner and provide better condition control with overall homogeneity within the structures. However, several issues led to a quick rejection of all previous attempts. Firstly, materials oxidized beyond desired proportions during calcination and drying in air, especially when polymer resin (mannitol) had to be thoroughly removed. Secondly, water is also a bad environment since it hosts a substantial amount of oxygen despite discovered substances being amphoteric -their precipitation was relatively easy, but elevated

O & OH content was substantial. Furthermore, in terms of protecting vital, not-fully-oxidized, amphoteric reactants like Ce^{3+} , W^{4+}O_2 , or $\text{W}^{4/5+}_{18}\text{O}_{49}$, acid was essential to be added, which often boosts the oxidation. Thirdly, using any good alkalized (NaOH) organic (alcohol) mixture with surfactants (like EDTA or DTPA) in mild temperatures (considering being stored below $180\div 200^\circ\text{C}$ in a pressurized autoclave to $6\div 8$ atm.) inevitably always results in the formation of high $20\div 30\%$ content of other, highly undesired luminescent-active impurities^{2,5,153,154} like A_2WO_4 , A_2WO_5 or even AWO_4 mixed with ($\text{A}=\text{Ba}/\text{Ca}^{2+}$) cerates or just simple ceria rendering further experiments useless. That led the team to believe that actual, slow migration control of well-mixed components during pressurized solid-state reaction in an inert atmosphere is actually the key.

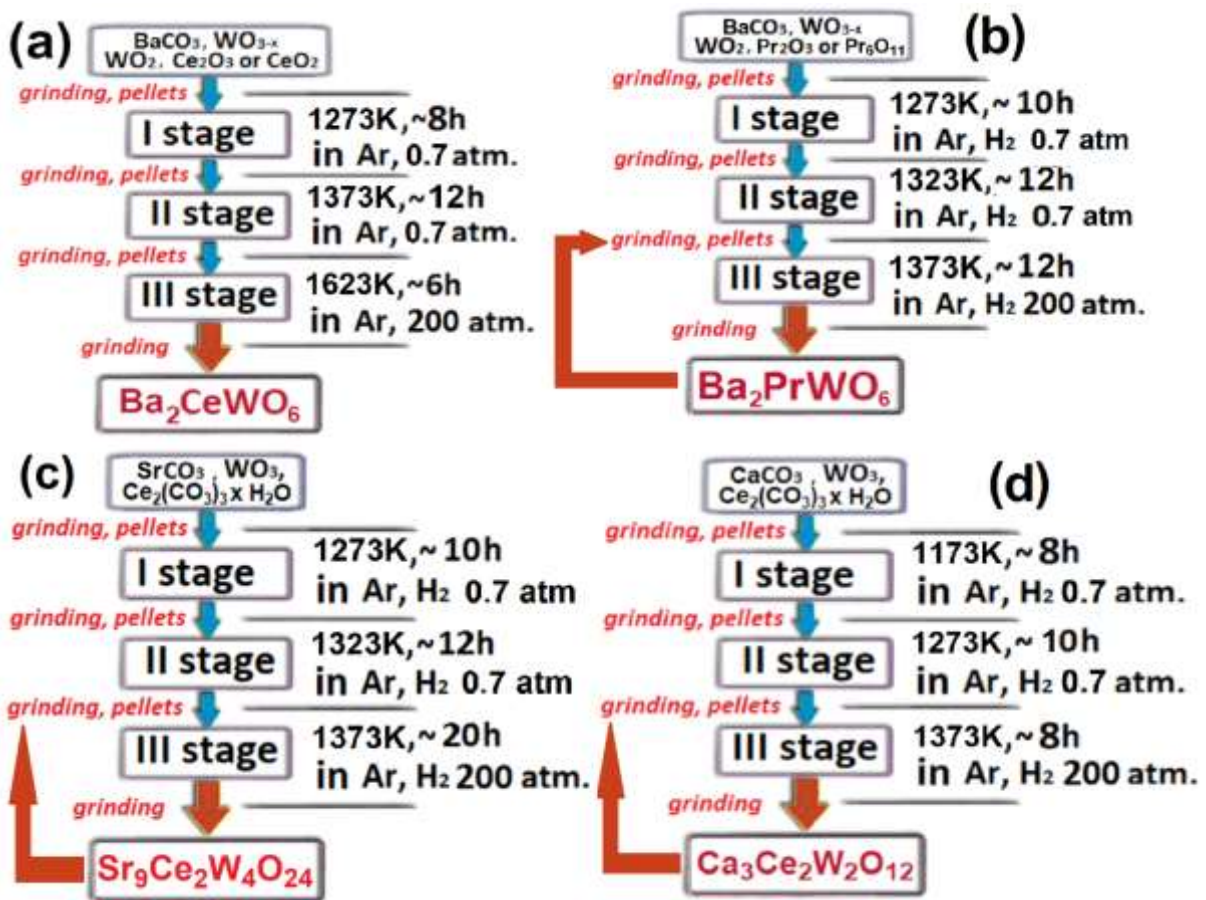


Figure 12 Ideological scheme depicting all optimized solid-state syntheses of investigated tungstates: (a) for BCW; and (b) BPW double perovskites, respectively; (c) for SCWO, which came out as something in the shape of quaternary perovskite; and closely related (d) CCWO ilmenite.

Therefore, all materials (in the form of powders) were synthesized using high temperatures in an oxygen-depleted/airless atmosphere via solid-state reaction in constant-flow reactors at temperatures above 1173K. Initially, the pressure was lowered to 0.7 atm. to improve the flow of the inert medium and desorption of gases coming out from decomposing reactants pressed earlier into tight pellets (to 10 MPa, $d\sim 13\text{mm}$). All of them were sintered in corundum crucibles after grinding with ethanol after each step. Later, the whole system was heated up (specifically to levels according to Tamman's rule)¹⁵⁵ and pressurized to ~ 200 atm. to increase the ionic diffusion rate of specific RE ions and ensure that the oxygen is being flushed out of the reactor. During 1st

stage, two intermittent (~15 min long) stops were made at 723 and 1173K due to the big release of H₂O, NH₃, and eventual NO_x, CO₂, respectively, depending on which reactants were used. The heat rate should not exceed ~10K/min, but cooling down below 1273K must be rapid to freeze the atoms in place so that oxides, cerates, and other simple A₂WO_{4/5} tungstates would not form. This regime should be maintained using an inert atmosphere down to at least 393-373K - mainly to avoid any oxidation and moisture absorption since these materials will turn out to be oxygen- (evolution with temperature) and light- (time-resolved, deep UV) sensitive. They turn from weak semimetallic phosphors to other insulating scintillators via decomposition. In order to prevent such changes & keep them as they truly are for more than eight months, a vacuum exicator in a dim room without any strong sunlight is recommended. One might also add 1÷5‰ of H₂ to the synthesis mixture to neutralize any possible free radicals during the synthesis.^{156,157} Sometimes, after grinding post-last stage and powder XRD check, a fourth round of synthesis was initiated, repeating everything mentioned in the last 3rd stage if some remaining reactants could be neutralized - depicted schemes were not just the bare minimum needed to achieve satisfying results. The crude nature of SSR already implies that sometimes a repetition might be needed even if all conditions are the same, however usually, it was enough – chemistry and internal processes could not be ideally controlled each and every time due to entropic reasons, but this was the average number of steps taken.¹⁵⁸⁻¹⁶⁰ If one is exemplary lucky, BCW can occasionally form a big, dark crystal with a violet tinge encased inside a tightly-pressed pellet. However, that result was achieved only once (not repeatable), sacrificing a lot of powdery exterior towards unwanted CeO_{2-x} & BaWO₄ impurities. Usually, a minor 1÷2% surplus of WO₂ and inert gas with 2500+ atm. pressure is also needed.

Products that were actually aimed at are: Ba₂CeWO₆ (BCW), Ba₂PrWO₆ (BPW), Sr₂CeWO₆, and Ca₂CeWO₆ in the form of double perovskites. Unfortunately, the latter two turned out to be something different (as seen in *Figure 12*) due to naturally caused entropy arising from apparent elongation or rather tilt of shrinking and twisting alkaline-oxygen polyhedral along one axis.^{37,69} First batches were inhomogeneous dualistic mixes of 50:50 desired phase versus those depicted in the aforementioned schemes. After many attempts, not much success was achieved in terms of thoroughly purifying them from all the extra content, but optimizing the conditions towards desired, 8+ BB'-charged double perovskites resulted in +93% purity – at least regarding barium. The other two, not so much. So, basing on pieces of information provided by Vasala and Karppinen,³⁷ a test was derived to see if the other 2 (Sr & Ca) compounds would be achievable in the other pure form given they naturally chose some different shape/stoichiometry. This side-tracking is common in DPs synthesis and is also worth pursuing according to science due to empirical investigations of structural disorder. The notes about the strontium-related compound at least state that it remains within the desired family of compounds but as a quaternary Sr₉Ce₂W₄O₂₄ (SCWO) perovskite.^{75,77-79} The last, Calcium one, formed so-called ilmenite Ca₃Ce₂W₂O₁₂ (CCWO) derivative.¹⁶¹⁻¹⁶⁶ Both with massive unit cells preferring 9+ B-site ion charge balance.

All the findings were documented (in upcoming *chapters 2, 3 & 4*) using methods mentioned before in the experimental *section (1.1)*, however, one peculiar thing stood out. All materials seemed to show dualistic BB'-site ion charge, which transfer was difficult to control using conventional stimuli like temperature, pressure, time-resolved PL, or reheating in reduction atmosphere/gases. So, an attempt was made to utilize alternative resources in the form of intermediate oxides, other than conventionally

known WO_3 , WO_2 , CeO_2 , or Pr_2O_3 , to see whether or not something more homogeneous (or even exotic) will be easier to obtain & naturally balance itself out. Consequently, WO_{3-x} (idealistically $\text{W}_{18}\text{O}_{49}$ closest to non-existent W_2O_5), heat-pretreated $\text{Ce}_2(\text{CO}_3)_3$ hydrate (as unstable Ce_2O_3 substitution), and Pr_6O_{11} (nearing perfect, hard-to-achieve PrO_2 structure) were involved but, as it turned out, to no avail. Cationic balance, later described in *chapter 2.3*, approached the same discrete values with passing time but from the opposite angle or charge-state boundary conditions. Meaning ions were seeking the same ratios from a different perspective. That could be somehow irreversibly manipulated (especially later using deep-NUV or X-ray light) but with some dire structural consequences for barium structures. It seems like the B-site ions were trying to compensate for some local oxygen vacancies and defects but to no avail, hence all this disproportionation and not uncommon charge-mixing effect.^{109–117}

As for Ce^{3+} ingredients, intercalated water in cerium (III) carbonate was assessed via TG and powder XRD patterns to show that it has mainly $\times 6\text{H}_2\text{O}$ with $\text{CeOH}(\text{CO}_3)$ & a small amount of $>1\%$ $\text{Ce}(\text{CO}_3)_2$. Decomposition and post-reduction of usual $4+$ ceria gave worse results of CeO_{2-x} complexes (like Ce_7O_{12} or Ce_3O_5), so this original substance (Ce^{3+} donor) seemed to be a more practical solution in producing $\text{Ce}^{3+}/\text{W}^{5+}$ phase (& dopants) in comparison to the other $\text{Ce}^{4+}/\text{W}^{4+}$.

Knowing that oxidizing Pr^{3+} oxides do not usually create pure PrO_2 but, instead, often host a complex mixture of $\text{Pr}_n\text{O}_{2n-2}$ phases (analogical to the reduction of CeO_2 or decomposition of its bicarbonates resulting with, at worst, a mix of 5 RE^{4+} ingredients)¹⁶⁷ we have decided to try with Pr_6O_{11} since other remaining phases ($1.585 < n < 1.833$) rapidly interchange between each other with alternating temperature, moisture, and carbon dioxide content (charge transfer and much more).^{168–171} It is the best, relatively stable Pr oxide having an equilibrium phase close to Pr^{4+} -rich PrO_2 hosting $4\text{PrO}_2 \times \text{Pr}_2\text{O}_3$.¹⁷² With it, we could hope for the best $\text{Pr}^{4+}/\text{W}^{4+}$ ratio with competing $\text{Pr}^{3+}/\text{W}^{5+}$ set within the matrix. On behalf of this and the previous paragraph, one could genuinely deduce that BPW and BCW reactants will behave incongruently compared to SCWO & CCWO analogs, which are built up from much more stable, fully oxidized B-site ions.

To accommodate for the failed attempts of creating a complementary, pure $\text{RE}^{3+}/\text{W}^{5+}$ phase (in comparison to $\text{RE}^{4+}/\text{W}^{4+}$, which resulted in a similar imbalance reaching the aforementioned discrete ratio but from the other side of the BB'-site spectrum), a W^{5+} carrier substance was created – at least as an intermediate solution for SSR. Early mentioned $\text{RE}^{3+}/\text{W}^{5+}$ cationic pair could be made predominant according to previously displayed recipes in *Figure 12*. Conventionally affordable methods on a laboratory scale could be modified for use from such articles.^{173–176} There, $\text{W}_{18}\text{O}_{49}$ has the best $4+/5+$ ratio of ions in all WO_{3-x} family. Also known as $\text{WO}_{2.72}$, made by T. Fita-Chala et al.,¹⁷³ has the best reaction properties recorded due date. It is the closest formula to non-existent, perfect $5+$ $\text{W}_2\text{O}_5/\text{WO}_{2.5}$ oxide that one could wish for but also the most unstable one. Slow oxidation, outside of suspension and protective environment, to W^{6+} is imminent; thus, we recommend using it no longer than 3–4 days after successful synthesis. W_2Cl_{10} alternative is not a good solution since it would contaminate structures with persistent chlorides even if it is truly stable.^{177,178} Because of those problematic issues, and the aforementioned BB' self-balance that still maintains itself despite using different reactants and environments, I refrained from studying this family of samples and simply chose to stick with $\text{RE}^{4+}/\text{W}^{4+}$ reactants. Simply because the one thing that changed was the balance of $\text{Ce}^{3+}/\text{Ce}^{4+}$ to $\text{W}^{5+}/\text{W}^{4+}$ - the properties remained the same. No pure ($\text{Ce}^{4+}/\text{W}^{4+}$ or $\text{Ce}^{3+}/\text{W}^{5+}$) cationic phase was finally achieved during the nontrivial synthesis attempts, so there was no difference in where to start. Furthermore, not only the mixed pairing was unavoidable, but also, using unstable oxides turned out to introduce more contaminants into the samples – they had more unintegrated, distorted oxides, and W^{6+} appeared at the end, further complicating the structure, creating more vacancies, mixed B-site pairs, and highly-unwanted minor phase in the form of A^{2+}WO_4 scintillators. That is why one should stick to $\text{RE}^{4+}/\text{W}^{4+}$ substrates.

2. Physicochemical characterization in ambient conditions

2.1 SEM imaging and morphology

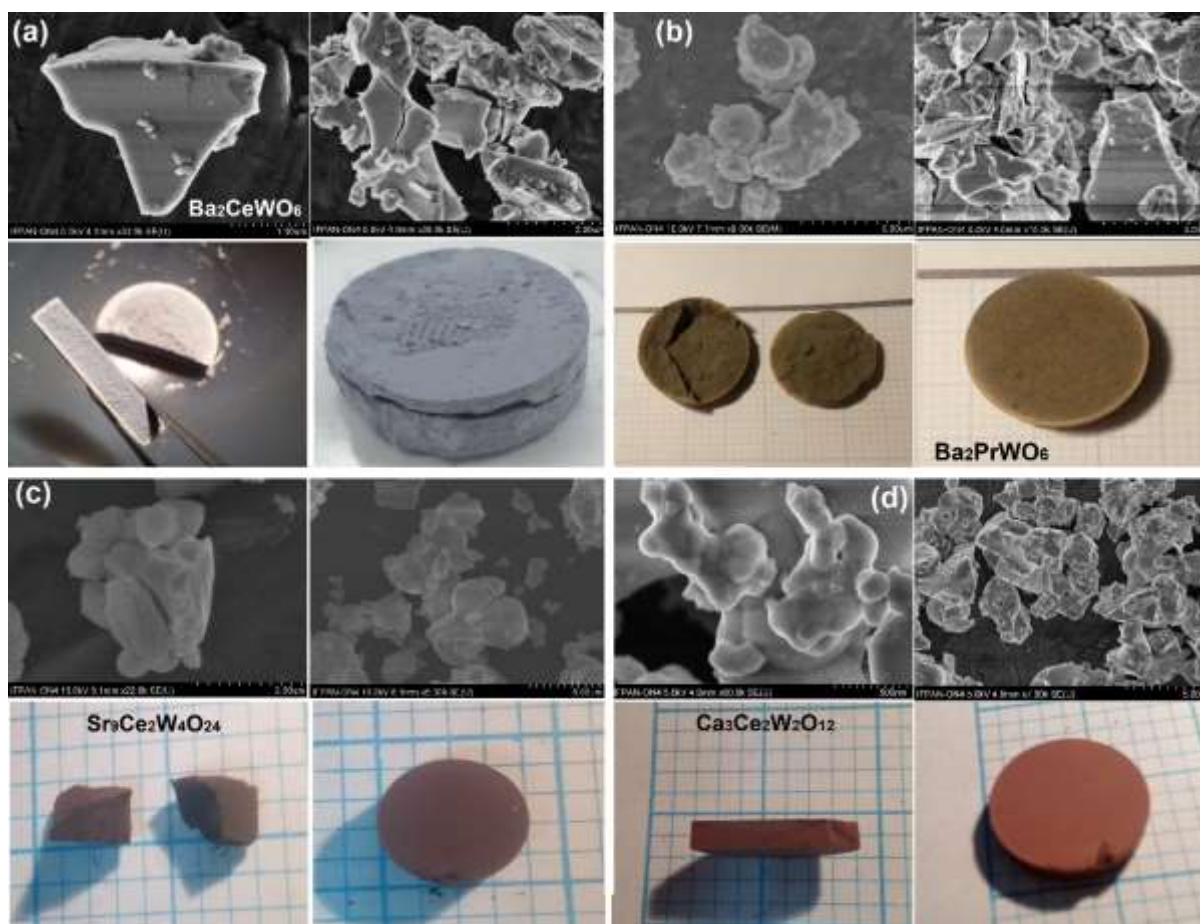


Figure 13 Micro (upper) and Macro (lower) -scopic photos of newly-created materials: dualistic (a) greyish-blue BCW; and (b) greenish-grey BPW double perovskites; with homogeneously brown (c) SCWO quaternary perovskite; and slightly lighter (d) CCWO ilmenite. One can notice that BCW & BPW have more defined microcrystalline structures, but their surface is riddled with defects. BCW's outer shell has a mostly unwanted, white, and brittle layer of (BaO+CeO_{2-x}). BPW, on the other hand, has a lighter crust containing BPWO+Pr₂O₃ intrusions.

Micro- and macro visualizations of all samples' final forms post-solid-state reaction are presented in *Figure 13*, where one can see that (a) BCW, just like its praseodymium counterpart (b) BPW, has more defined, smaller, polycrystalline structures, at least from the inside. Both double perovskites show dualistic nature and some pronounced discoloring on the outside. As it turns out, the brighter exterior shells are more defective and riddled with minor phases. Steel-grey color BCW crystalline polygons (a) are covered with a brittle (BaO+CeO_{2-x})-rich layer that could be easily scraped. It occurs due to slow but inevitable WO₂ evaporation that could be mitigated by adding a slight surplus of that ingredient. BPW case (b) is less severe since the brighter layer contains more Pr³⁺ in the form of mixed BaPr₂WO₇+Pr₂O₃ impurity phases (majority in terms of total interior content is ~Pr⁴⁺/W⁴⁺ pair as shown later in XPS/XAS related paragraph). SCWO (c) quaternary perovskite and (d)

CCWO ilmenite also share few similarities (being visibly different from the previous two materials) and are a lot more homogeneous. However, occasionally those two congruent substances form a darker hue on the surface - a very thin but grainy in touch, which is easy to remove even with a simple touch of a scalpel or tweezers – it is most probably a layer of W/WC ceramics originating from oxidized W and carbonate reactants. That is also why one should not use graphite crucibles as sample holders during HT synthesis. Both perovskite derivatives have brownish color – SCWO being slightly darker. Widely spread micro-to-nano-sized spherical particles are mostly cluttered together as fine dust, which is not easily separated and seems to attract each other statically; thus, each particle’s size is much more difficult to assess. Before divulging into analytics, it is worth mentioning that all SEM experiments were performed on actual grains extracted from the middle of the pellets where the actual phase resided - XRD was also done on samples collected the same way. Outer layers were still mostly scraped off for a separate measurement after realizing the material's dual character shortly after cracking the pellet in half shortly post-synthesis. We also wanted to avoid eventual contamination from interior grains falling into the middle of the pellet while cracking them open.

A statistical overview done in Image J software on randomly-chosen, well-defined 20 grains (d_{AV20}) is summarized in *Table 5*. Debye-Scherrer equation: $d_{AVDS} = \frac{K \times L}{FWHM \times \cos(\theta)}$ was also incorporated to provide some insight into this subject based on provided XRD patterns. According to tables in Langford, Willson, and Scherrer's work,¹⁷⁹ shape constants (K) were derived from 3 chosen, well-distinguished XRD peaks at higher 2Θ angles.¹⁸⁰ Their hkl's and FWHM's, alongside with some visual recognition, were crucial to estimate proper calculating factors; however, one must note that for grains larger than 200 nm, these Debye-Scherrer estimates are much less accurate. This highly depends on the limitations originating from the sheer size of used X-ray wavelength (L) – in our case, equal to Cu $K\alpha$.

Table 5 A summary of estimations and calculations performed respectfully on SEM images and powder XRD patterns to determine grains' actual size and shape parameters for all newly acquired rare-earth tungstate materials.

Material	d_{AV20} [nm]	$\pm\sigma_{AV20}$ [nm]	Shape K_{hkl}	Peak angle [2Θ]	Peak FWHM [Rad] $\times 10^{-3}$	d_{AVDS} [nm]	$\pm\sigma_{AVDS}$ [nm]
BCW	781.3	95.5	0.88÷0.90	134; 92.5; 69.6	3.2; 1.86; 1.4	861.6	48.0
BPW	854.6	90.6	0.88÷0.92	110; 69.8; 85.4	2.6; 1.9; 2.08	913.6	7.0
SCWO	1128.3	104.9	0.83÷0.86	81.3; 72.8; 44.1	1.6; 1.7; 1.5	972.9	42.6
CCWO	1035.8	114.7	0.84÷0.86	77.2; 44.6; 31.9	1.86; 1.5; 1.5	915.9	17.8

Sometimes the shape (K) and (d_{AV}) of the particles were hard to assess because of joined conglomerates in large proportions hardly visible through SEM microscope, like in the case of SCWO and CCWO. This was most probably due to low contrast from low conductivity/charging effects and the aforementioned electrostaticity of the fine dust. However, some trends could be easily established based on just a few lone, widely scattered particles present on the edges of the browsed area, as seen in *Table 5*. Interior BCW and BPW microcrystals (which we should focus on due to lack of contaminants) are generally

slightly smaller, with BCW being the tiniest. Their shape is also more cuboidal (higher K constants) due to their polyhedral, polycrystalline nature. BCW also has more diverse crystal sizes, as can be deduced by average deviation factors (σ_{AV}). The last statement also stands for CCWO; however, alongside with SCWO, their characteristics open up a different group or rather a pair of conjoined similarities: they are substantially larger, as mentioned before, their particles attract each other like in fine dust; and finally, their shape is more spherical as showcased by the K value. SCWO is also the largest of them all. From a statistical standpoint, d & σ parameters established from SEM imaging (AV20) and XRD patterns (AVDS) do not differ that much on the nanometer scale and show the same tendency, where $SCWO > CCWO > BPW > BCW$. They provide different insights into the same phenomena that somehow complement each other. However, from an experimentalist point of view, regarding previous remarks about Debye-Scherrer equation limitations, SEM imaging seems to be a more reliable source, especially because one physically sees the nature of incoming results in a larger specimen group.

2.2 Powder XRD analysis

Ba_2CeWO_6 (BCW) and Ba_2PrWO_6 (BPW) double perovskites are identical in many crystallographic features, as seen in *Figures 14 & 15*, respectively. First, they turned out to be as planned in shape and form. The upcoming analysis is based on similar isostructural conclusions regarding SG assignment, polymorphism, changing axial constraints, and dualistic $RE^{3+/4+}/W^{5+/4+}$ ionic nature. Even in such doubtful subjects like overall appearance, impurities, degradation, and LT phase transitions, everything is connected, so it is just natural to start the discussion with those 2.^{181–183}

According to ICSD crystallographic database, their isostructural predecessors are $Fm-3m$ Ba_2BiYO_6 (record nr 65555) from Lenz et al.¹⁸⁴ work or Ba_2SrWO_6 (246108).¹⁸⁵ Other typical examples of such already-developed structures might be found amongst $Ba_2Ce_{3/4}\square SbO_6$ or $Ba_{1.75}\square_{0.25}BaWO_{5.75}\square_{0.25}$, where \square is a vacancy.^{183,184,186,187} Garrido Barrios event predicted such BPW creation and appearance using just DFT calculations.¹⁸⁸ Rietveld refinement, however, showed poor results regarding peaks' shapes vs. intensities ratio for this specific space group (BCW $R_{wp} \sim 23.5\%$; BPW $\sim 23.9\%$). BCW acquired first,¹⁸³ ideologically has an almost cuboid unit cell – slightly monoclinically tilted, straight β angle by 0.015° (at worst batch 0.04°). Similarly, BPW, but with an even more subtle tilt of $\sim 0.01^\circ$. Best fittings (BCW $R_{wp} \sim 20.9\%$; BPW $\sim 17.5\%$) suggested that we have encountered a problematic case of mostly indistinguishable patterns: of $I2/m$ (Glazer tilt system $a^0b^-b^-$)⁶⁹ shown above and alternative, $R-3$ (a-a-a-) SG stored in supplementary *Figures A4 & A5*. Their refinement data were summarized in *Table 6* and *Table A1*, respectively. It does not mean that other possibilities were not explored - the rest of rejected SG matches and the whole meticulous fitting process are presented as insets of aforementioned *Figure A4(a-d)* (and refer to BPW processing too) but, without additional assets like neutron diffraction, or information about behavior outside of ambient conditions, gathered through, i.e., HP XRD, there was no way to ascertain the final point group matches at that point in time. Especially after all (4) applications for much-needed synchrotron time were rejected, the only possibility was to follow Lufaso¹⁸¹ and Howard⁷¹ et al. research where they saw $I2/m \leftrightarrow R-3$ interchanging into

one another at specific p(T) conditions – they state that both SG could coexist together at room temperature since their phase transition belongs to broad second order type later turning into more recognizable, monoclinic $P2_1/n$. To limit future discussion, let's focus more on the former, $I2/m$, since this SG assignment will be more accurate at the end of the discussion.

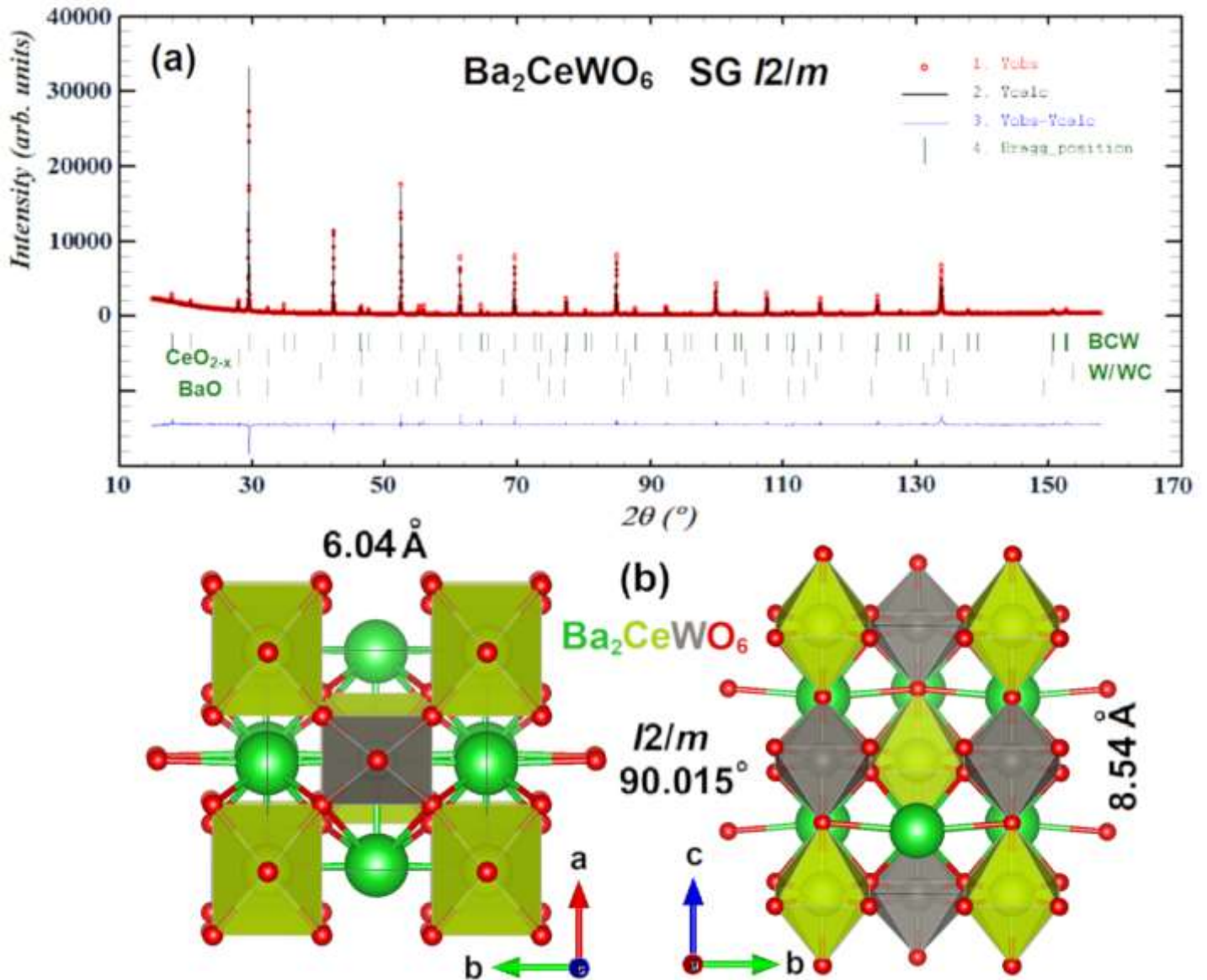


Figure 14 Original XRD pattern (a) acquired after Rietveld fitting and refinement of best, 95% pure $I2/m$ phase containing traces of BaO and highly-distorted CeO_{2-x} (probably Ce_7O_{12}) oxides alongside minor traces of W/WC phase. Red dots are the manifestation of a naturally observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown in partially ball-stick, partially polyhedral convention in figure (b) with some additional structural parameters alongside the ‘c’ (left) and ‘a’ (right) axes. Yellow spheres and polygons are ceria, red is oxygen, and grey is assigned to tungsten. Barium is depicted solely as green spheres to improve overall visibility since its dodecahedra would cover most of the view.

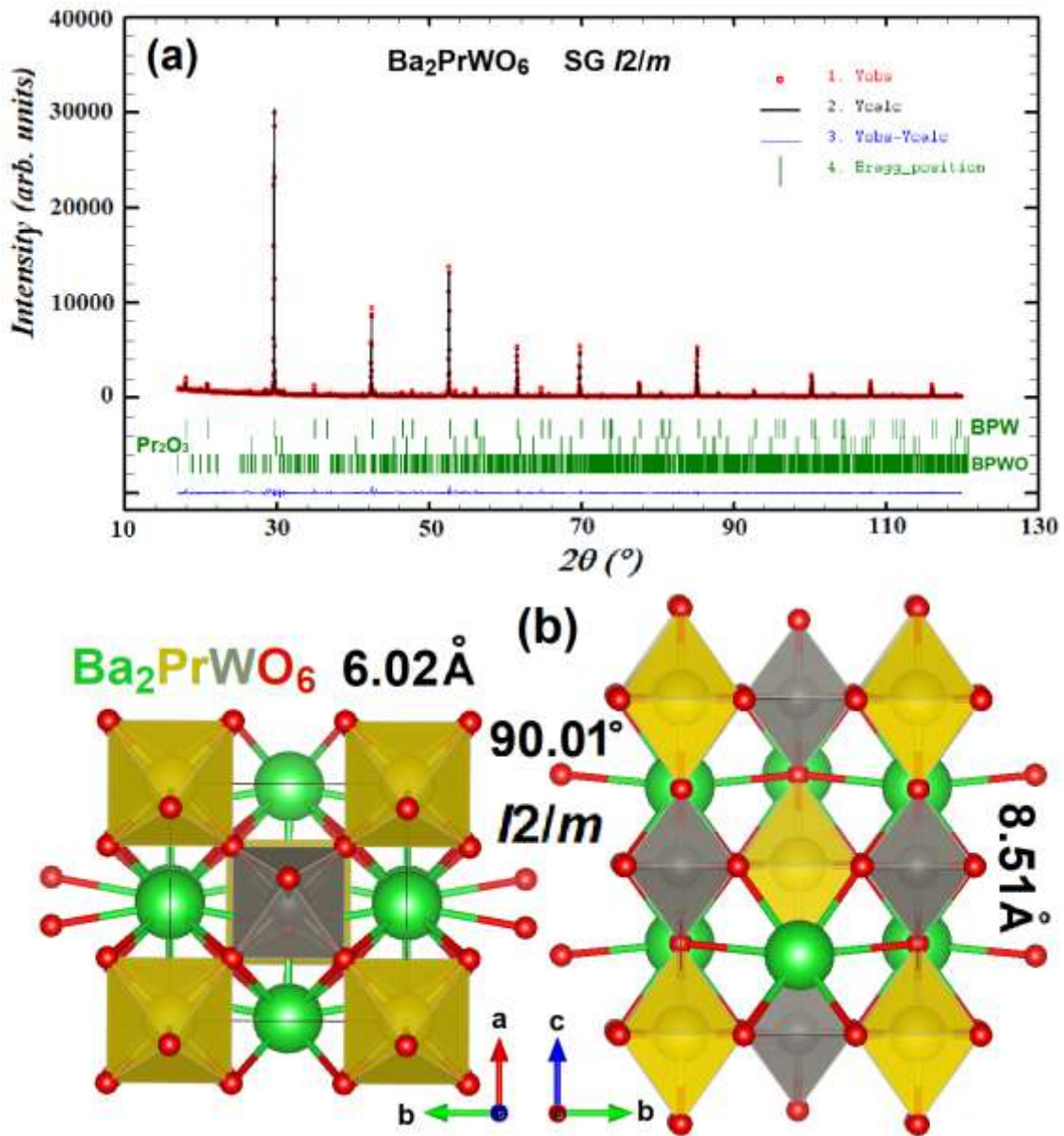


Figure 15 Original XRD pattern (a) acquired after Rietveld fitting and refinement of best, ~96% pure $I2/m$ phase Ba_2PrWO_6 (BPW) containing traces of $BaPr_2WO_7$ (BPWO) and distorted Pr_2O_3 oxide impurities. Red dots are the manifestation of a naturally observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown in partially ball-stick, partially polyhedral convention in figure (b) with some additional structural parameters alongside the ‘c’ (left) and ‘a’ (right) axes. Yellow spheres and polygons are Pr ions. Red is oxygen, and grey is assigned to tungsten. Barium is depicted solely as green spheres to improve overall visibility since its dodecahedra would cover most of the view.

The unit cell of $I2/m$ SG is generally pretty small and compact. The tilt is not even that visible. The same can be said about BPW – the only difference here is that Pr ions localized in central octahedra are slightly smaller due to the bigger electronegativity of the shrinking RE ionic radii and, consequently, the total volume compared to Ce.¹⁸⁹ That ultimately also influences the distortion of β angle, constricting the monoclinic cell more towards its perovskite-like cuboid character. All crystallographic parameters for this and other, mainly chosen SG are summarized

in Table 7; meanwhile, alternatives, like R-3 (which analysis would be analogical despite having larger axial distances and volume), are placed in supplementary Table A2.

Finally, synthesized double perovskites were not exactly pure. Much struggle has been experienced while optimizing the solid state reaction (SSR) via the shake & bake method – to improve batch from 86% to 94-95% of needed content. The most encountered impurities were sweated out RE oxides (BCW = CeO₂, BPW = Pr₂O₃) joined with their alkaline counterparts (BaO), seemingly trying to form some sort of distorted cerate in the case of BCW - Figure A6; or praseodymate – BaPr₂WO₇ regarding BPW (Figure 15).¹⁹⁰⁻¹⁹³ Those minor phases resulted from bad RE ion migration dictated by Tamman’s rule (connected to elevated melting points),¹⁵⁵ surface effects, and environmental exposure of the pellet’s outside layers. Some RE are also known for their mixed valence character (in that case referring mainly to the substrates) like Ce₇O₁₂ (ICSD 4113), Ce₃O₅ (621709)^{194,195} or Pr₆O₁₁ (69109), and Pr₅O₉ (105542).^{196,197} Those exotic compounds like to form due to subtle chemical imbalance, lower enthalpy, and uncontrollable, time-dependent charge transfer partially influenced by atmospheric oxygen. That pronounces their occurrence to be problematic and somehow unavoidable. Though the judgment of XRD HighScore databank redirects us to Pr₂O₃ (61179)¹⁹⁸ in terms of BPW (Figure 15) and possibly Ce₇O₁₂ species for BCW (Figure A7), all remarks mentioned above and in Table 4 would suggest better reference those intermediate species as CeO_{2-x} or PrO_{2-x} group, where x: 0<x<0.5 – mostly 0.167<x<0.286.^{169-171,194,196,199-201}

The other most common impurity is W or WC originating from graphite crucibles or contact between used carbonates with evaporating WO₂. As mentioned before, those ceramics are easily removable, but that instigates usage of a small surplus of WO₂ (not higher than 1÷2%) while creating a synthesis batch. Too much of it or the presence of residual oxygen (i.e., after not evacuating the furnace or not creating an inert atmosphere with slightly lowered pressure that enables the gas to come out of the compressed substates) will definitely result in creating excessive amounts of BaWO₄ – Figure A8. This will not only negatively affect Raman but also emission and excitation spectra creating false signals and covering more valuable, subtle bands of particular interest.^{2,202,203} BCW and BPW could experience such intrusions because they are built up from BaREO₃×BaWO₃ as tungstate-based double perovskites. Already crystallographic inconsistencies in BaO and CeO_{2-x} lattice implied a try-hard fusion to BaCeO_{3-x} species.

Table 6 Conventional Rietveld reliability factors for all investigated compounds and their minor phases present in the best, possibly purest batch. Data refers to diffractograms hosting Bragg contribution patterns corrected with the background to the most plausible, undoped SG. Alternates are presented in the supplement Table A1.

Formula	SG	Z	V [Å ³]	d ^{cal} [g/cm ³]	R _B	R _P	R _{WP}	R _{EXP}	N _σ GoF	χ ²	Fract [%]
Ba ₂ CeWO ₆	<i>I2/m</i>	2	311.342	7.410	13.5						94.92
CeO _{2-x}	<i>Fm-3m</i>	4	164.194	6.963	43.2	23.6	20.9	4.36	1304.809	22.5	2.39
BaO	<i>Fm-3m</i>	4	168.010	6.062	36.7						2.37
W(/WC)	<i>Im-3m</i>	2	31.695	19.264	98.9						0.32
Ba ₂ PrWO ₆	<i>I2/m</i>	2	308.843	7.479	5.75						95.71
BaPr ₂ WO ₇	<i>P112₁/b</i>	4	623.818	7.613	28.5	18.3	17.3	2.66	2674.274	42.2	3.97
Pr ₂ O ₃	<i>P-3m</i>	1	77.676	7.051	57.9						0.32
Sr ₉ Ce ₂ W ₄ O ₂₄	<i>I4₁/a</i>	4	2201.420	6.602	3.82	11.1	13.2	6.24	223.051	4.45	99.41
W(/WC)	<i>Im-3m</i>	2	31.677	19.257	72.4						0.59
Ca ₃ Ce ₂ W ₂ O ₁₂	<i>R-3c</i>	18	4528.414	6.338	4.29	8.31	9.72	2.67	799.386	13.2	99.63
W(/WC)	<i>Im-3m</i>	2	31.683	19.261	67.0						0.37

Table 7 Crystallographic parameters obtained for all acquired structures regarding their mainly discussed, undoped space groups stored in CIF files hosted at the CCDC database. Alternative SG projections are shown in the appendix alongside with similar summary in Table A2.

Formula	Main SG	Unit cell axis length [Å]			Angles [°]			Volume [Å ³]
		a [100]	b [010]	c [001]	α	β	γ	
Ba ₂ CeWO ₆	<i>I</i> 2/ <i>m</i>	6.0376(0)	6.0369(0)	8.5420(0)	90.000	9.015	90.000	311.34(2)
Ba ₂ PrWO ₆	<i>I</i> 2/ <i>m</i>	6.0219(3)	6.0218(3)	8.5167(3)	90.000	9.010	90.000	308.84(3)
Sr ₉ Ce ₂ W ₄ O ₂₄	<i>I</i> 4 ₁ / <i>a</i>	11.6220(9)		16.2980(1)	90.000		2202.42(0)	
Ca ₃ Ce ₂ W ₂ O ₁₂	<i>R</i> -3 <i>c</i>	9,7258(3)	55.2793(0)		90.000	120.000		4528.41(4)

Sr₉Ce₂W₄O₂₄ quaternary perovskite (SCWO) has a larger, densely packed (over 345 atoms) unit cell riddled with many defects from shared Ce/Sr or half-occupied A-atom sites. Its complex nature, alongside with evenly unique XRD pattern, can be seen in Figure 16. Dimensional data is stored in Table 7. Though there is still much debate about the assignment of these freshly-classified compounds, finding a proper isostructural matrix was not yet an easy task among those few articles that reported on them. Smirnov was the first to discover Sr₉La₂W₄O₂₄ as an archetype in the late 80s.²⁰⁴ His work has been, however, buried in time since he did not publish any relevant structural data in the form of CIF. Only after some time (the late 90s) Xie⁷⁹ and Bramnik²⁰⁵ et al. developed their own Sr₉RE₂W₄O₂₄:Eu (RE=Gd, Y) structures (JCPDS 50-0375 and 49-0568 respectively)²⁰⁶ on the basis of well-known Sr₁₁W₄O₂₄ and have noticed his input. The early 2000s brought us Ba₉Ln₂W₄O₂₄ and a wide variety of Sr-related compounds Sr₉Ln₂W₄O₂₄ (Ln = La, Nd, Eu, Gd, Y),^{75,78,204,206-209} but the accurate description and debate came with the work about Sr_{7/4}□_{1/4}SrReO₆ which referred to Ba₁₁Os₄O₂₄ as its ancestor.²¹⁰ The thing is, most of these compounds are cubic (*Fm*-3*m*) in nature and have neatly long-range order of defects stacked throughout enormous unit cells. Our SCWO falls out of that first assignment– it is undoubtedly tetragonal, *I*4₁/*a*, just like BaWO₄ and those last 2 Sr/Ba compounds. Its refinement data can be viewed in Table 6 and confronted with the *Fm*-3*m* alternative in supplementary Table A1. The difference in R_{wp} (Tetragonal SG ~13.2% while cubic ~ 80.2%) will be astronomical, making the final choice rather obvious. Figure A9 & Table A2 also show those remaining doubtful SCWO projections with simplified dimensional parameters.

This compound is relatively pure and homogeneous despite generating much debate amongst scientists (mainly about the current compound assignment within the ilmenite/garnet subgroup, which connections are not clearly agreed upon). The only noticeable traces of impurities that were somehow distinguishable in a meaningful way from the background noise was W/WC ceramics present only at the surface. As mentioned before: very thin and easy to remove, at almost negligible content ~0.59% mass.

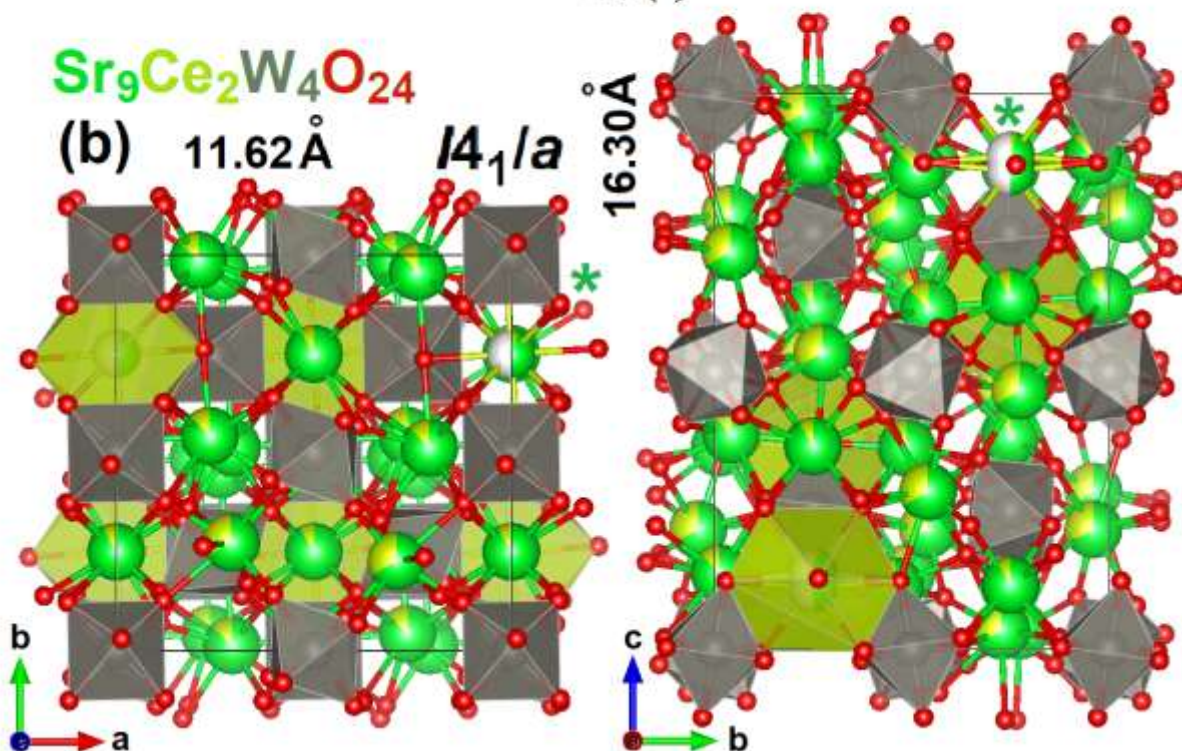
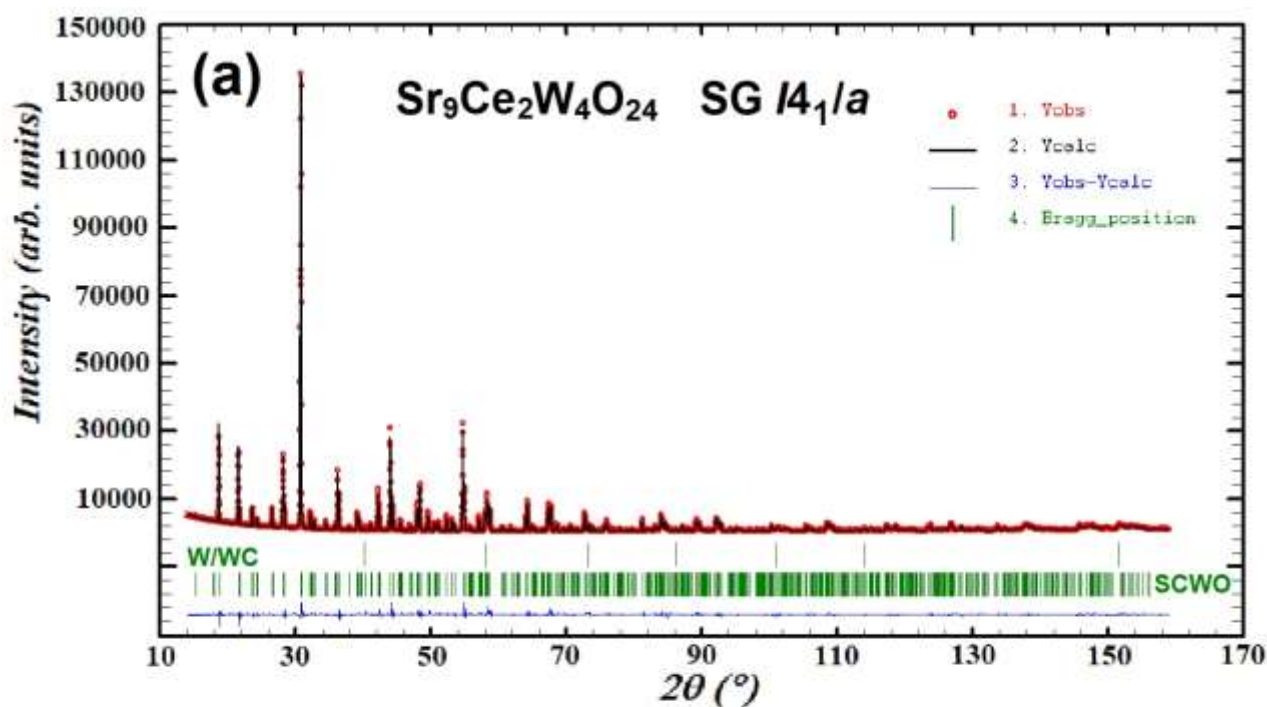


Figure 16 Original XRD pattern (a) acquired after Rietveld fitting and refinement of best, $\sim 99\%$ pure $I4_1/a$ phase $\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$ (SCWO) containing traces of W/WC impurities. Red dots are the manifestation of a naturally observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown in partially ball-stick, partially polyhedral convention in figure (b) with some additional structural parameters alongside the 'c' (left) and 'a' (right) axes. Green spheres are strontium atoms. Yellow is cerium. Red is oxygen, and grey is assigned to tungsten space filling. Some of the Sr and Ce polygon projections were omitted to improve the overall visibility. Green asterisks mark artificially removed polygons to showcase a unique feature of some neighboring, half-occupied Sr/Ce sites (0.497/0.03).

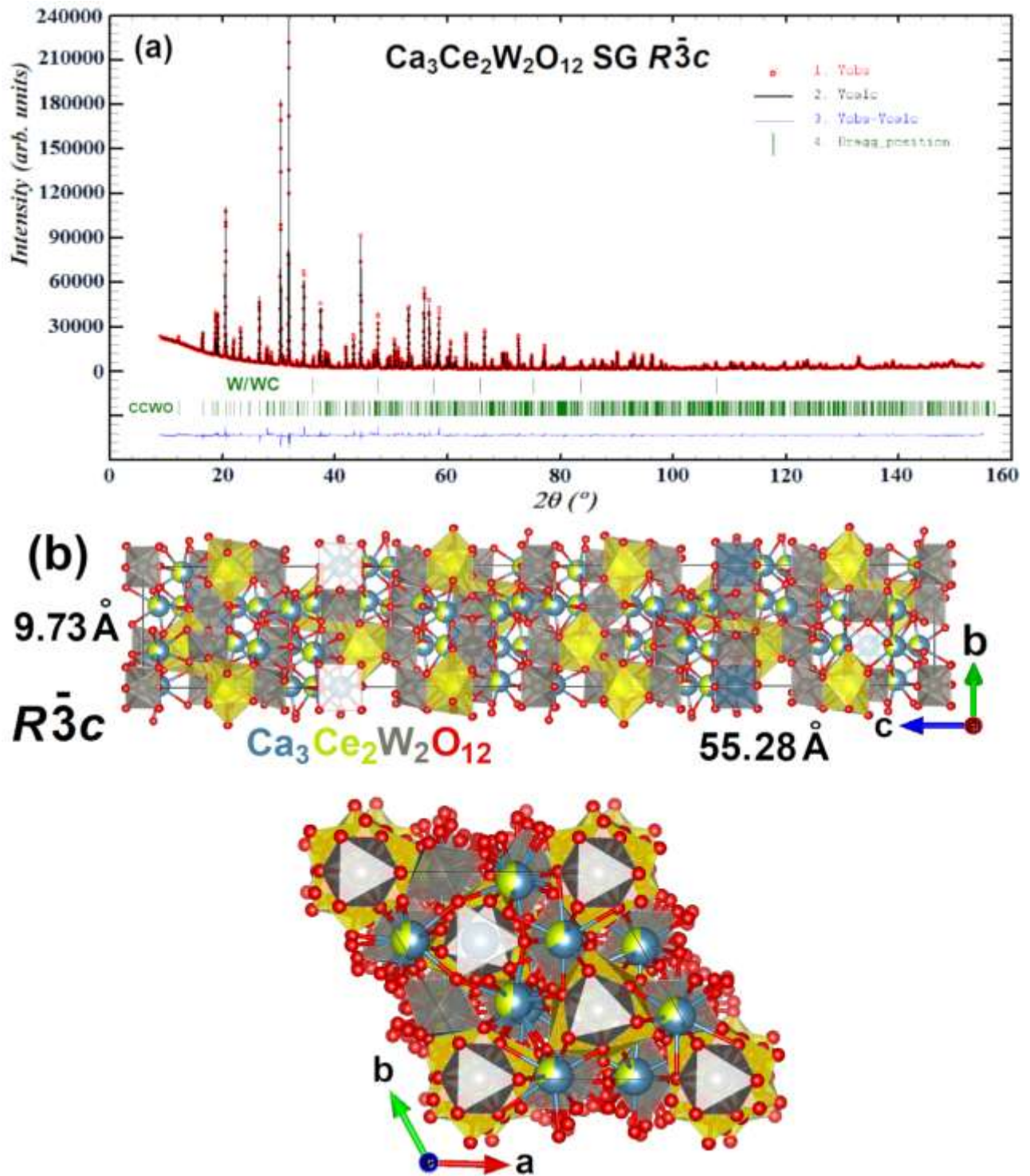


Figure 17 Original XRD pattern (a) acquired after Rietveld fitting and refinement of best, 99.7% pure $R\text{-}3c$ phase containing only minor W/WC impurities. Red dots are the manifestation of a naturally observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown in partially ball-stick, partially polyhedral convention in figure (b) with some additional structural parameters alongside the 'c' (down) and 'a' (up) axes. One might notice that both calcium and cerium atoms partially share some sites, so only 2 of them (with different shapes and most pure occupation) were chosen to be depicted as polyhedral to maintain visibility. The remaining atoms are: yellow – cerium, red – oxygen, and grey – tungsten.

Similar statements are true for $\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$ (CCWO) ilmenite presented in *Figure 17* above. It even looks the same internally (giant unit cell with staggering 342 atoms additionally sharing A^{2+}/RE sites) and externally - it is in a slightly brighter shade of brown, as shown in *Figure 13*. The only impurity would be an identical, minuscule amount of W/WC phase ($\sim 0.37\%$) allocated on the surface. Furthermore, it behaves the same in various conditions and is only one atom short of typical garnets like $\text{Mg}_3\text{Al}_2\text{Si}_3\text{O}_{12}$.²¹¹ So, it truly makes a point to give a question of whether or not CCWO is an ilmenite or a garnet - analogically to SCWO if it is really a quaternary perovskite or something else. Only more profound studies and statistical overview could tell. We definitely saw a lot of common features between them and garnets, so I support that statement.

CCWO is directly related to its isostructural counterpart in the form of already discovered $\text{Ca}_3\text{La}_2\text{W}_2\text{O}_{12}$ with JCPDS pattern no 49-0965.^{163,164,166} This compound, however, was reported to have its own archetype – $\text{Ca}_5\text{RE}_2\text{O}_{12}$ made by Villars et al.²¹² He, unfortunately, did not determine all lattice parameters through structural refinement so we were obliged to do and present them in *Table 8*.

Table 8 Unique structural details of newly acquired $\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$ structure, based on the Rietveld refinement of preferably chosen $R\text{-}3c$ space group. Note that U_{iso} has been set to be the same for all O and W atoms, respectively.

Site	x	y	z	U_{iso} [\AA^2]	Occ. (<1)
Ca1	0.0093(2)	0.35914(19)	0.18694(2)	0.0033(4)	0.678(2)
Ce1	0.0093(2)	0.35914(19)	0.18694(2)	0.0033(4)	0.322(2)
Ca2	0.37275(16)	0.04466(14)	0.04948(2)	0.0017(3)	0.5541(19)
Ce2	0.37275(16)	0.04466(14)	0.04948(2)	0.0017(3)	0.4459(19)
Ca3	0	0	0.10415(3)	0.0007(6)	0.303(6)
Ce3	0	0	0.10415(3)	0.0007(6)	0.697(6)
Ca4	0	0	0.25	0.043(2)	
W1	0.33174(11)	0	0.25	-0.00009(14)	
W2	0	0	0.162812(18)	-0.00009(14)	
W3	0	0	0	-0.00009(14)	
O1	-0.0035(14)	0.1863(10)	0.22340(15)	0.0025(9)	
O2	0.0036(13)	0.1713(11)	0.02094(14)	0.0025(9)	
O3	0.1355(10)	0.1911(11)	0.17969(16)	0.0025(9)	
O4	0.1863(10)	0.0745(11)	0.14022(15)	0.0025(9)	
O5	0.1927(10)	0.3107(12)	0.10760(14)	0.0025(9)	
O6	0.2460(8)	0.1049(9)	0.0836(2)	0.0025(9)	

As seen above, the disputed unit cell is huge and rhombohedral – a straightforward assignment. The only problem is whether SG is centrosymmetric ($R\text{-}3c$) or not ($R3c$). Again, in the case of powder XRD, they are indistinguishable. Even their VESTA/CIF projections do not show any visual differences at first glance, given that those cells are so complicated and densely packed. They exhibit the same hkl parameters hence $R3c$ projection was omitted. The presence of a symmetry center could be confirmed by more sophisticated techniques like high-resolution single-crystal XRD or NPD; yet again, we had not received access to those facilities (rejection of proposals during Covid), nor was crystal form of this substance achievable at Polish institutes. The other possibility was to reach out to the Kharkiv Institute of Single-Crystal, but the war came. However, upcoming sections about Raman spectroscopy

and polymorphism outside ambient conditions slightly turn the balance towards *R-3c* SG. Again, all conventional refinement reliability factors with structural parameters for chosen SG are stored consecutively in *Tables 6 & 7*, whereas alternatives in appendix *Tables A1 & A2*.

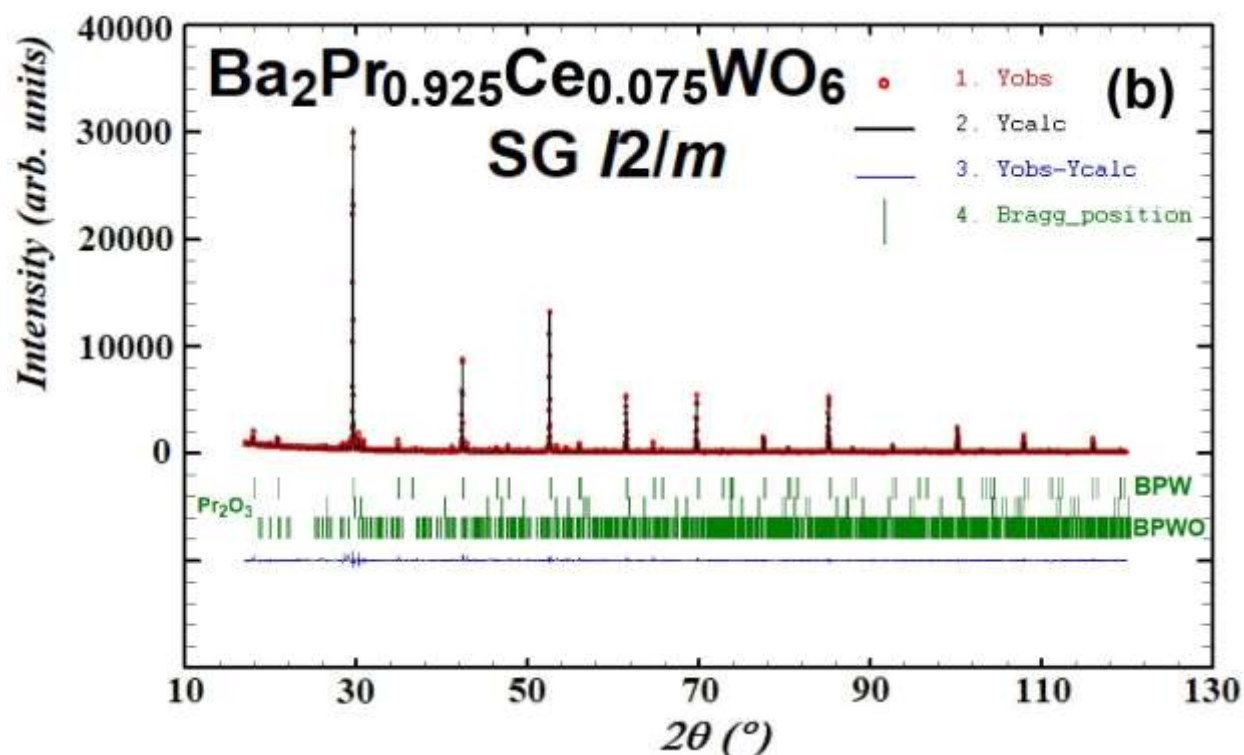
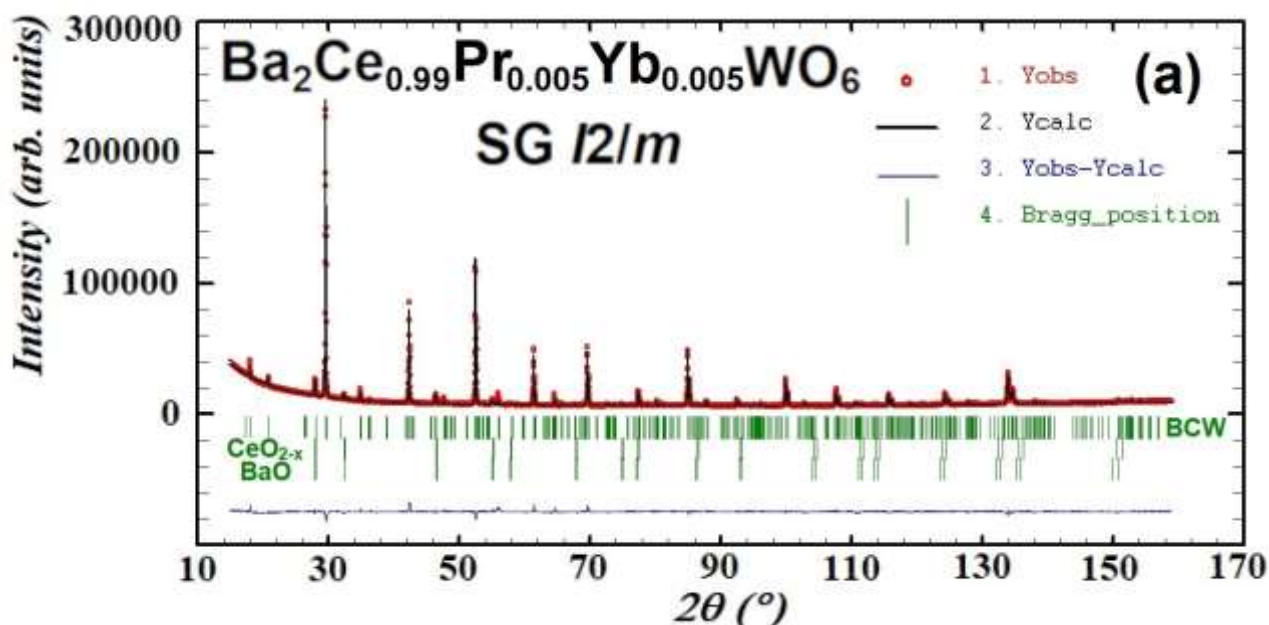
The repeatable occurrence of shared multisites is worth mentioning, but now only between Ca and Ce – similar to shared Ca/La sites in Li, Hirrlinger, and Autenrieth et al. cases.^{163,164,166} No noticeable vacancies and half-occupied sites, in any case. However, the material is still defective, which will be showcased later in the luminescence chapters. That should not be a surprise since the volume of CCWO is approximately two times bigger than SCWO. That inflammation might be responsible for better internal distribution of atoms in a longer perspective. More details about exact occupancies – their sites, position, neighborhood, and percentages can be viewed online via CCDC CIF file repositories. That applies not only to precedencies' of CCWO and SCWO but also to more simplified structures like BCW and BPW.

Doped materials look genuinely the same and will be shown and discussed only, referring to the preferred SG. The main concern were the impurities – will chosen RE dopants fully integrate with the matrix given theoretically chosen substitution places? First of all, cerium-tungstates, so BCW, CCWO, and SCWO, were substituted with Pr³⁺ and/or Yb³⁺ to accommodate planned NUV→VIS→NIR energy transfer. The amounts of chosen dopants to (ideologically) swap with Ce³⁺ were 0.5, 2.5, and 5%, respectively. These values refer not to the total amount but to the amount of each RE ion if stacked together. That way Ba₂Ce_{0.99}Pr_{0.005}Yb_{0.005}O₆, Ca₃Ce_{1.9}Pr_{0.05}Yb_{0.05}W₂O₁₂, and Sr₉Ce_{1.8}Pr_{0.1}Yb_{0.1}W₄O₂₄ were acquired. Of course, since Ca and Sr partially share sites with Ce, the intentional substitutional effect is also mixed – resembling much Ba²⁺/Ce exchange in BaWO₄:Ce mentioned at the beginning of introduction² - Yb/Pr could be lodged there as well. Regarding sole praseodymium tungstate BPW instead of Pr constituent, a 7.5% amount of Ce was introduced.

The other thing were the minor phases that can occur outside of already reported ones in pure matrices. So in all cases ubiquitous W/WC; in BCW – something else than BaO, CeO_{2-x} (samples with BaWO₄ were automatically discarded due to substantial PL interference & misleading Raman response); Ba₂Pr_{0.925}Ce_{0.075}WO₆ would have BPWO and PrO_{2-x} intrusions. The presence of those REO_{2-x} oxides was most unfortunate from the beginning, however, not entirely avoidable. There PL interference was to be expected, yet, after 1.5 years of optimization, experiments must still have been tried out - the only solution was to be sure that Yb₂O₃ would not fall out of the matrices just like Ce^{3/4+} oxides in BPW or Pr^{3/4+} in BCW, SCWO or CCWO. Thus, it did not, as seen in *Figure 18* and related to it *Table 9*.

Table 9 Conventional Rietveld reliability factors for all investigated compounds and their minor phases present in the best possible batch. Data refers to diffractograms hosting Bragg contribution patterns corrected with the background to the most plausible, doped SG.

Formula	SG	Z	V [\AA^3]	d_{cal} [g/cm^3]	R_B	R_P	R_{WP}	R_{EXP}	N_o GoF	χ^2	Fract [%]
BCW:Pr,Yb	<i>I2/m</i>	2	311.342	7.423	13.4						94.49
CeO _{2-x}	<i>Fm-3m</i>	4	168.103	6.802	25.9	28.4	19.4	3.95	1488.539	24.0	2.89
BaO	<i>Fm-3m</i>	4	168.846	6.033	29.9						2.62
BPW:Ce	<i>I2/m</i>	2	308.918	7.477	6.41						95.01
BaPr ₂ WO ₇	<i>P112₁/b</i>	4	624.984	7.599	41.1	26.4	25.2	10.82	245.779	5.44	4.47
Pr ₂ O ₃	<i>P-3m</i>	1	77.893	7.032	65.7						0.52
SCWO:Pr,Yb	<i>I4₁/a</i>	4	2202.314	6.600	5.39	12.1	14.9	3.69	992.835	16.2	99.29
W(WC)	<i>Im-3m</i>	2	31.698	19.211	64.6						0.71
CCWO:Pr,Yb	<i>R-3c</i>	18	4522.411	6.357	6.66	12.3	13.2	4.49	510.730	8.69	99.58
W(WC)	<i>Im-3m</i>	2	31.669	19.211	64.9						0.42



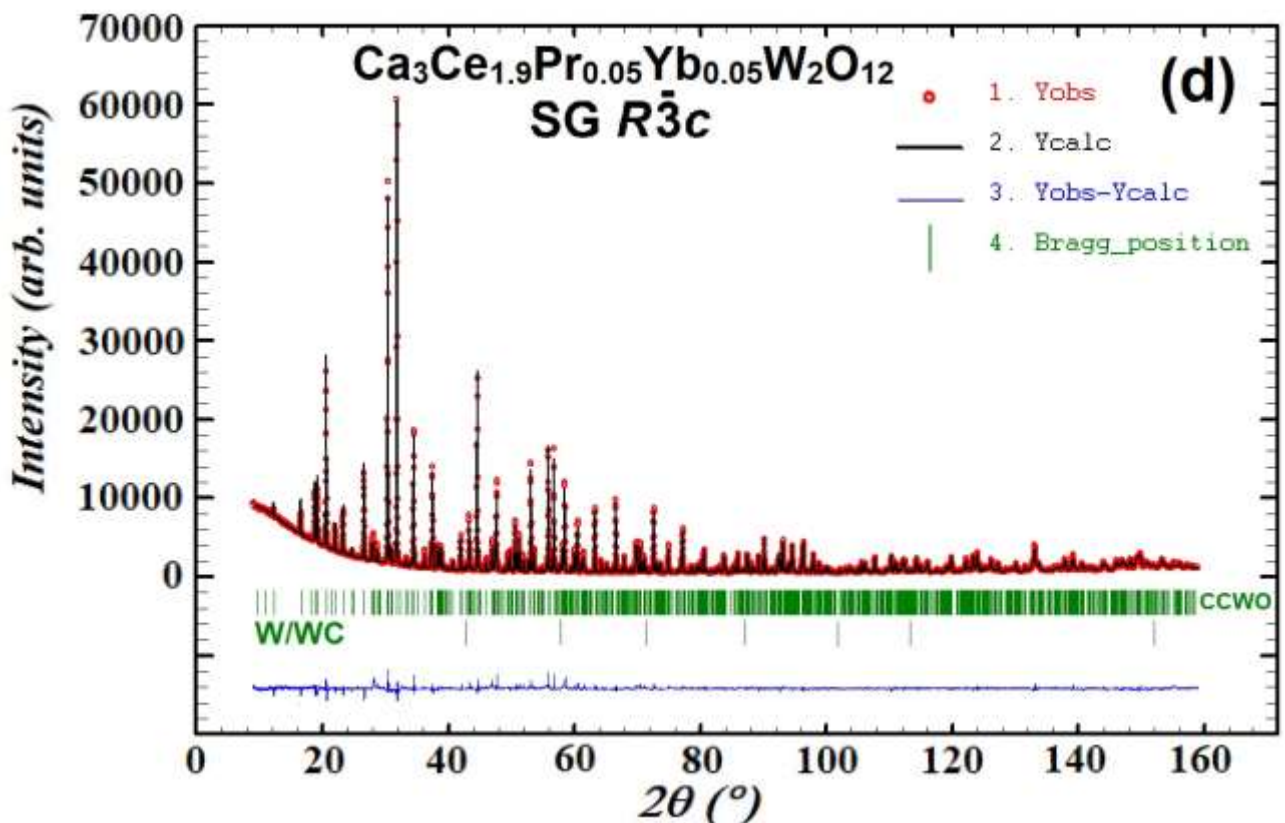
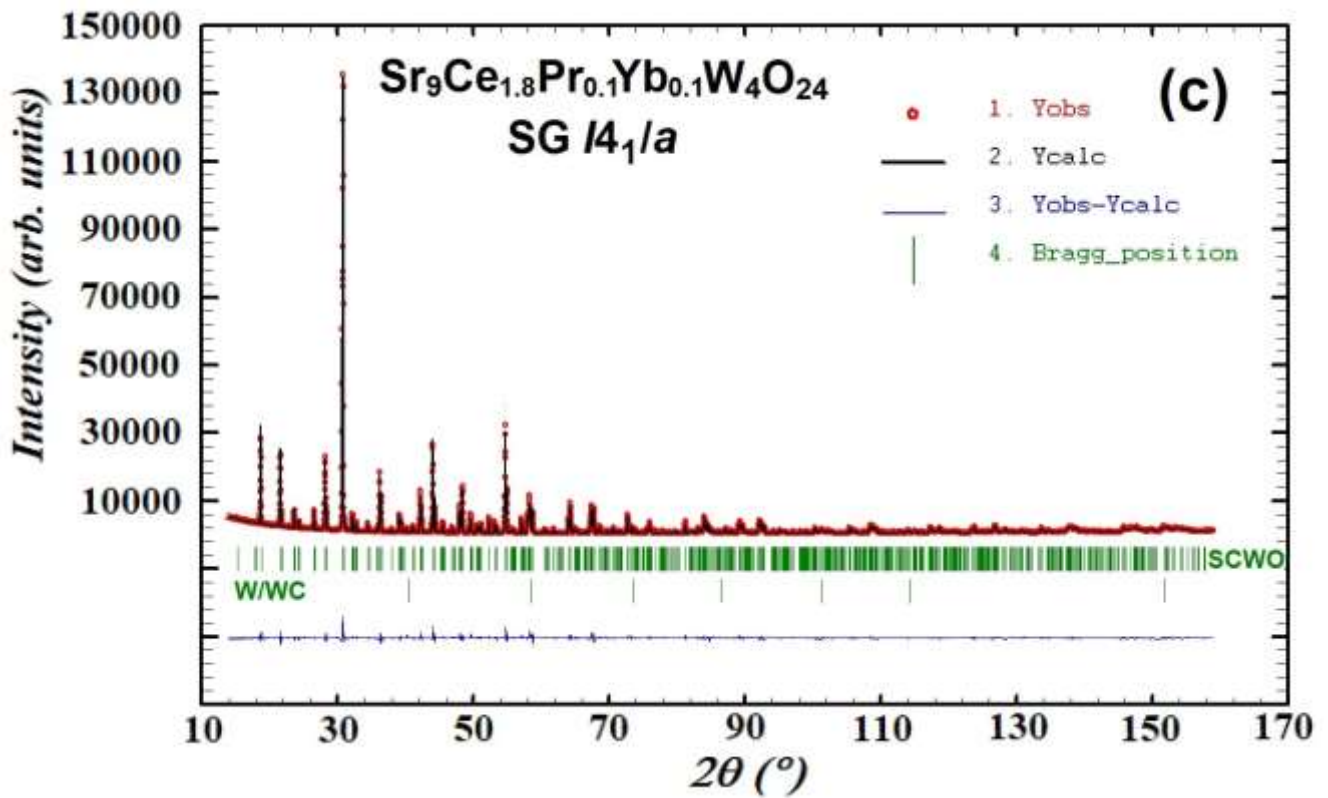


Figure 18 Refined XRD patterns of doped: (a) BCW:Pr,Yb (0.5%); (b) BPW:Ce (7.5%); (c) SCWO:Pr,Yb (5%); and (d) CCWO:Pr,Yb (2.5%). Red dots are the manifestation of a naturally observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases.

2.3 X-ray spectral characterization (XPS & XAS)

The application of other complementary X-ray analytic techniques was necessary to estimate the actual content of our powderish materials. This means assessing the cation diversity for carefully balanced structural or stoichiometric purposes and regarding optical properties.²¹³ It is a well-known fact that only a few ions having specific electron configurations in unique bonding states (i.e., non-empty valence shells, paired mesomeric systems in short or long-range conjunction with good excitations sources) possess proper radiative response if built in a suitable matrix of sufficient crystallographic field strength.^{214,215} Furthermore, BB'-site substrates used here have from 2 (RE) to 3 (W) different charges available that might change into one another during the synthesis (such is their unstable nature), so their ratio and appearance must be examined. Shortly speaking, optically active Ce/Pr³⁺ could easily change to naked-shell 4+, rendering them partially useless.^{167,195,196} The same stands for W 4 or 5+, which should be semi-metallic and provide a small band gap compared to W⁶⁺, which is more insulating.^{216,217} These things are crucial to monitor when developing novel, PL-active matrices while searching for an eventual explanation of synthesis/property failure, incoherent behavior outside ambient conditions, or simple defect probation. Characteristics of other applied ions could also help in optimizing various processes, establishing hypotheses, confirm or debunk many questions taking into account that this study is also fundamental in its origin, so oxygen and alkali ions (such as Ba, Sr, and Ca)²⁺ will also be briefly discussed further. For now, let's just shift our focus to the most important RE cations presented in *Figure 19*, where XAS (deep penetrating technique) & XPS (more surface-related) measurements seem to show some coherent postprocessing results.

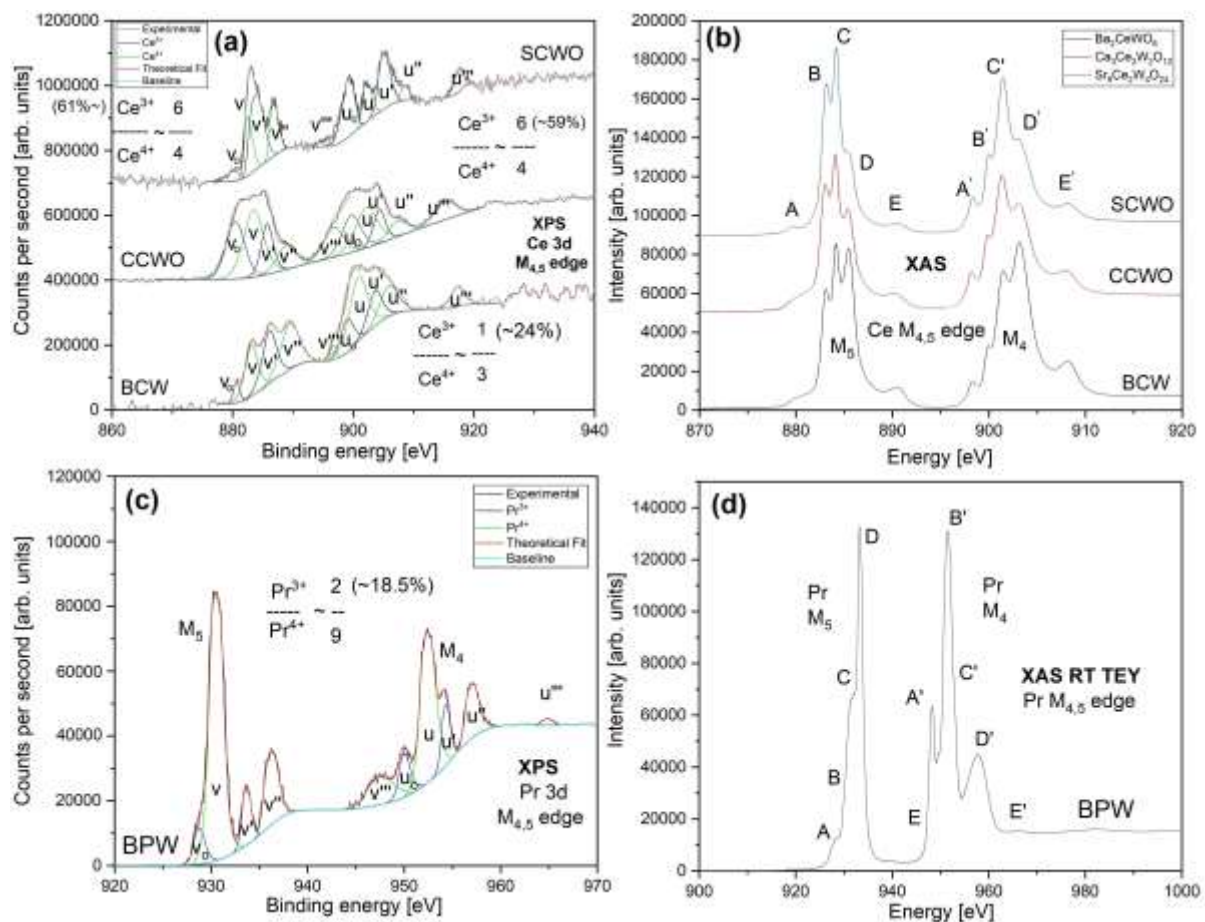


Figure 19 Crucial X-Ray Spectra collected at room temperature and ultra-high vacuum for B-site rare-earth 3d atom shells such as Ce for BCW, CCWO, and SCWO using (a) XPS and (b) complementary XAS techniques. Figures (c) and (d) feature the same methods in respective order but for Pr in BPW.

Cerium and Praseodymium 3d spin-orbit shell and $M_{4,5}$ absorption edges show complex, multiplet band structures at their respective energies (880-915 and 925-965 eV, respectively). The versatility of ionic forms could be assessed by careful and simple deconvolution, knowing need-to-know basics from the literature and reference spectra collected pure substrates at own leisure, as shown in appendix *Figure A10*. Two sets of five XPS peaks ($3d_{5/2}$ v, v_0 , v' , v'' , v''') and $3d_{3/2}$ (u, u_0 , u' , u'' , u''') could be found at the beginning ($3d_{5/2}$) or end ($3d_{3/2}$) of each range mentioned previously.²¹⁸⁻²³¹ Despite having slightly different principles, intensities, and ratios, XAS presents its content the same way - shallower M_5 edge being the first set (A-E) while deeper M_4 edge (A'-E') is shown later, at higher energies. The spectrum of pure Ce^{4+} (i.e., CeO_2) should consist of three multiplets related to $4f^n$ ($n = 0,1,2$) electron interactions denoted as v, v'' and v''' components in $d_{5/2}$ shell (marked as green)^{227,231-234} similarly to B, D, E peaks in XAS. Separated by an 18.5 eV gap and typical 1.5 v/u intensity ratio should be mirrored by a $d_{3/2}$ set (u, u'' , and u''') withholding analogical meaning – in XAS: B', D' and E'.²³⁵⁻²³⁷ The rest of the peaks are denoted to Ce^{3+} in the same manner (and marked as blue), so $4f^n$ ($n = 1,2$) XPS (XAS) modes related to $d_{5/2}$ component (M_5) are marked as v_0 and v' (A,C) while $d_{3/2}$ (M_4) as u_0 and u' (A',C').^{219,225,228,229,234,238,239} Knowing this, establishing proper positioning, peak area, and FWHM could be used for calculating the actual Ce^{3+}/Ce^{4+} ratio from totally added content. $Pr^{3/4+}$ signals sets are interpreted somehow analogically but at higher energy ranges.^{220-223,240} Apart from the information provided in the insets of *Figure 19*, one can view them collectively in summary provided in *Table 10*, but for now, the true aim is to establish whether or not B and B'-site ions (Ce/Pr and W) can balance each other out in the core – could they be coexisting nearby in subtle, charge-balanced octahedral sites showing some explicit dependencies, and, if so, will they explain eventual discrepancies in future optical or polymorphism studies? Or possibly justify the existence of other perovskite derivatives?

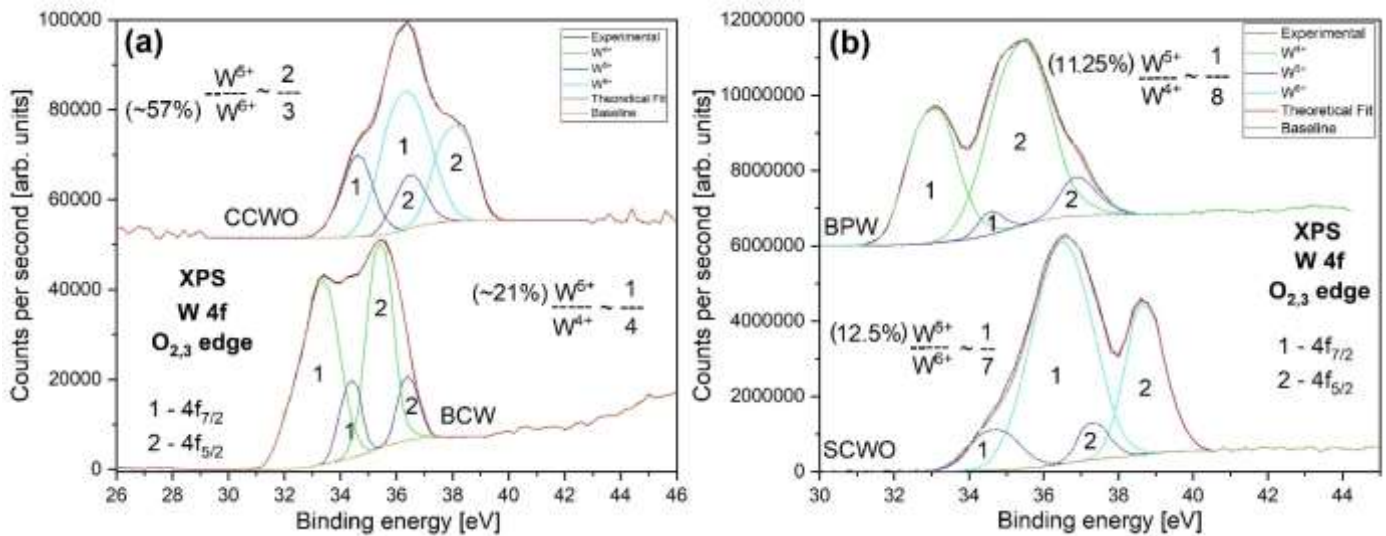


Figure 20 XPS 4f-shell spectra taken at ambient conditions for all four investigated materials regarding the other B'-site ion, tungsten. Figure (a) refers to BCW and CCWO, respectively, hosting fully balanced 8+ or 9+ composition to their Ce counterparts, while (b) presents uneven BPW and SCWO.

Hence, the joint quantitative analysis with W 4f (O_{2,3} edge) presented in *Figure 20*. Such grouping is not coincidental. Both Ba_2CeWO_6 (BCW) and $Ca_3Ce_2W_2O_{12}$ (CCWO) in *Figure 20a* have very close, complementary content to their RE counterparts showcasing structural balance in pronounced ionic duality within the matrix. Double perovskites seem to prefer 8+ BB'-site core balance compared to derivatives, such as ilmenites, where this number is higher by one, 9+. That fact is satisfied by: the Ce^{3+}/Ce^{4+} ratio in BCW, equal to roughly 1/3 (24/76%), while W^{5+}/W^{6+} does not fall far behind with its 21/79% values meaning, consequently, that Ce^{3+} should accommodate for its neighbor W^{5+}

maintaining aforementioned 8+ balance. Thus, creating a phase sustaining $\sim 1/5$ to $1/4$ of total content. The rest, $\sim 75\div 80\%$, would create $\text{Ce}^{4+}/\text{W}^{4+}$ ‘artificial’ pairs. Artificial because this scenario seems somewhat idealistic – there is no way of telling whether or not this type of segregation is genuinely taking place in the nearest neighborhood - rather not. The ions are mostly spread apart, not ideally – distanced, compensating the charge only throughout the whole volume of the sample. In powders, this phenomenon would create a lot of local defects and vacancies via not perfectly balanced multisites, which are undoubtedly present. Future discussion will tell us more on that matter, but what is most important for now is that ‘pairings’ and ‘couples’ will be used loosely to abbreviate such conditions as stated above to shorten an already elaborate discussion.

So, CCWO follows close the same patterns as Barium DPs but $2/3$ of $\text{Ce}^{4+}/\text{Ce}^{3+}$ (41/59%) ions are joined by $\text{W}^{5+}/\text{W}^{6+}$ ‘pairs’ (43/57%) featuring the same $2/3$ combination of 9+ phases. This would analogically form $\sim 60\%$ of the $\text{Ce}^{3+}/\text{W}^{6+}$ matrix while the rest would be scattered as $\text{Ce}^{4+}/\text{W}^{5+}$ environment. Again, not essentially well mixed or agglomerated, as PL behavior would also suggest (in future *chapter 4*). The situation complicates for the other pair in *Figure 20b* and requires some extensive commentary because of the prolific lack of W^{5+} in comparison to $\text{Ce}^{3+}/\text{Ce}^{4+}$ content. That suggests many possible problems in the form of gaps, vacancies, or other distortions forming within the unit cell and creates an opportunity to elaborate more on the nature of XPS 4f shell build. In each picture, two sets of overlapping $4f_{7/2}$ (1 - lower energy) and $4f_{5/2}$ (2 - higher energy) spin-orbit doublets are visible – between 30 and 40 eV per sample. The lowest pair (green) is denoted to W^{4+} and usually mismatches the WO_2 reference not exceeding ± 0.5 eV.²⁴¹ Another set (blue) was first thought to be W^{6+} (later colored as cyan), however, the WO_3 reference has shown that its peaks are localized much higher. After much research, the discrepancy matched other WO_{3-x} species containing some exotic W^{5+} ions.^{242–245} That was a pleasant surprise considering their high instability (no pure W_2O_{10} in existence); however, compared to the other XPS-related articles showcasing materials like W_2Cl_{10} as a reference, there was at least some proof to justify the existence of such cations with some additional, future evidence. The position mismatch was also small – not bigger than $\pm 0.2\div 0.3$ eV.^{246,247} Next, all cyan W^{6+} signals were firm enough that their characteristics and positioning did not cause any interpretation issues. Fitting WO_3 reference perfectly, the only concern was big FWHM which suggested a lot of scattered multisites and interference from ionic mismatches scattered throughout a complicated, powderish matrix.^{244,248–250} Again, the analyzed parameters are summarized in *Table 10*. Pure, metallic W (impurity detected by XRD) would also be visible here in between 31–33 eV, just below W^{4+} doublets; however, possibly due to low ($<1\%$) content, it was not detectable, especially if it was primarily present as scrapable WC ceramics localized on the surface of originally acquired pellets.^{251–254}

Focusing solely now on Ba_2PrWO_6 (BPW) and $\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$ (SCWO) content, Pr/Ce vs. W ratios vaguely suggest the same conclusions as for BCW and CCWO regarding 8 and 9+ charge balance. However, the concentration of exotic W^{5+} ions seems to be upset – the proportions are $1/8$ and $1/7$, respectively, compared to their predominant W^{4+} and W^{6+} counterparts. This pronounced mismatch is so significant that it is not safe to say whether or not (18.5%) $\text{Pr}^{3+}/\text{W}^{5+}$ (11.25%), (81.5%) $\text{Pr}^{4+}/\text{W}^{4+}$ (88.5%) in BPW; or (39%) $\text{Ce}^{4+}/\text{W}^{5+}$ (12.5%), with (61%) $\text{Ce}^{3+}/\text{W}^{6+}$ (87.5%) phases in SCWO actually form. There is a distinct probability in BPW, though. Both materials would try to compensate for such distortions in some shape or form. Since, according to powder XRD, SCWO is riddled with a lot of A^{2+} half-occupied vacancies, this W^{5+} deficiency would be highly responsible for all eventual V_o holes or gaps. To some extent, this also stands true for BPW, which hosts a considerable amount of interstitial oxygen (O_i) but more about that in future EPR and optical chapters regarding charge-transfer phenomena.

Lastly, it is worth mentioning that XAS spectra of W were unavailable to provide any support since the capabilities of BL04 beamline's equipment exceeded noteworthy edges (just outside of provided energy range or due to detector limitations).

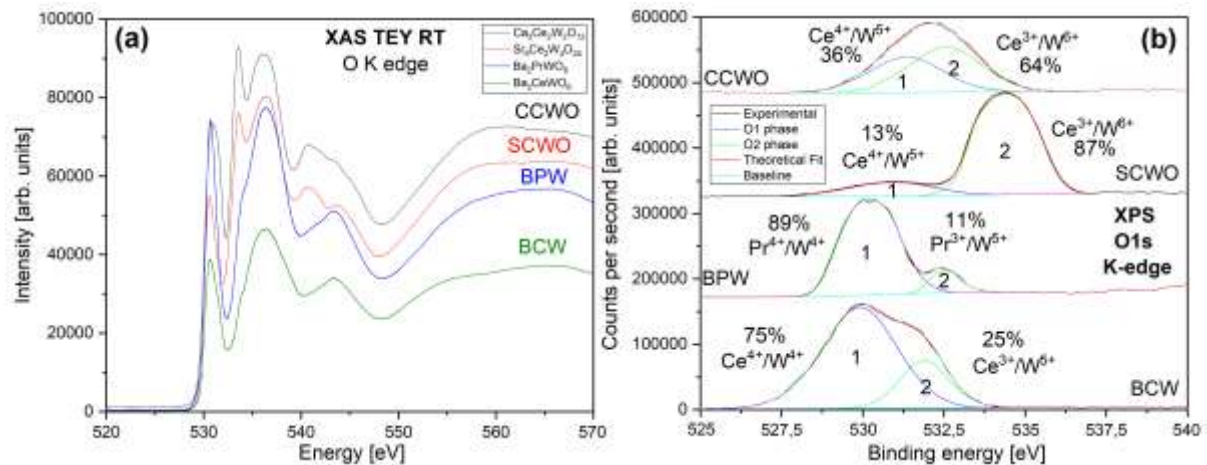


Figure 21 Convolved oxygen K-edge XAS (a), and simple XPS O1s spectra (b) taken at ambient conditions for all tungstates to depict their actual characteristics and compare various interatomic bondings and recurring matrix interactions.

Figure 21 showcases data about numerous oxygen σ - π bondings present within investigated cerates/tungstates/praseodymates and oxygen interactions with the local chemical environment. Since all materials were made using all kinds of oxides, this element's shell or edge could tell us some crude but valuable information about overall interactions, polarity, vacancies, and interjections within these systems. It serves as a unique fingerprint to justify eventual material group assignment; however, many features will also naturally but inevitably repeat themselves.^{255,256}

XAS K absorption edges in Figure 21a are more complex and, after deconvolution, tell more about the materials, so it is good to start with them. Several bands are grouped in well-defined regions and will be interpreted and compared collectively according to the literature²⁵⁷⁻²⁶⁰ since precise DFT calculations without information from additional NPD measurements would be just pure guessing. All edges start with an abrupt onset at around 530 eV assigned to the antibonding π^* bond. The rest peaks onward are just different σ^* bonds of various origins. Second, weakly developed shoulder might appear in derivatives, more visible in SCWO and CCWO, since hybridized antibonding O2p-B-site d-molecular orbitals might differentiate from the other, second B'-site 'pair' like RE or W^{6/5+} (in CCWO ~6:4 ratio) and hypothetically W^{5/4+} ions (weaker in BCW ~2:8). More critical is the 533 eV mode typical for ilmenites (significantly weaker in BCW or BPW structures). Its intensity/ratio is not so prominent even for SCWO, but, for sure, this material is more like CCWO in terms of internal structure - truly an intermittent (in between perovskite-ilmenite like) material riddled with a lot of multisites, vacancies, and defects affecting the structure via distorted A²⁺/RE-O-W bonds via dislocated dangling (O⁻) or interstitially trapped (O_i) oxygen. In derivatives, it is undoubtedly related to alkaline atom partially-substituted or half-occupied site trying to break off the current symmetry. Moderately distinguishable features at 534-538 eV seem to be simultaneously visible in all the structures. They describe interactions between A²⁺ ions, oxygen, and chosen d-block type states, i.e., 5d W-O-Sr 4s / Ba 6s or 5d/4f Ce-O-Sr 4s / Ba 6s. However, they show little to no changes in terms of structural displacement after deconvolution. The literature does not state anything out of the ordinary in this region, but at least the lack of sharper peaks still supports partial substitution influencing local amorphization of the structure. Hybridization difference is mainly driven here by the bonding angle of the central ions controlling the bandwidth. That is why the region near 540-545 eV in core structures

of higher symmetry would show some prolific differences. According to it, alongside the peak around 533 eV, dissociating derivatives from double perovskites is more plausible. Not only does the outside appearance support this claim, but also later spectral appearance and SG changes (HT, LT) and assignment through Howard et al. research.^{71,72} Shared, substituted A, and B-sites with prominent vacancies contribute to signals closer to 540 eV. Many different interactions at B-sites make those bands more distinguished; however, corner-sharing arrangements in perovskite cells also tend to have two reinforced oxygen bands visible in that region – although at slightly higher energies (at the end of the energy range mentioned above ~545 eV). It is due to highly-induced, short-range coordination associated with the formation of triply coordinated oxygen and self-achieved long-range periodicity of joined double triclusters ($2 \times \text{ABO}_3$). After this, the edge creates an extensive loss feature.^{261,262}

Complementary XPS spectra are depicted in *Figure 21b*. Although less detailed, all of them show two unevenly split bands telling us still enough about two types of oxygen bridges currently available - between various B-site W-O-Ce and W-O-Pr ion species in both derivatives (with predominant $\text{Ce}^{3+}/\text{W}^{6+}$ against $\text{Ce}^{4+}/\text{W}^{5+}$ 9+ balanced ‘couples’) and double perovskites (containing mainly 8+ $\text{RE}^{4+}/\text{W}^{4+}$ vs. $\text{RE}^{3+}/\text{W}^{5+}$ ‘pairs’, $\text{RE} = \text{Pr}/\text{Ce}$).^{263–266} In BPW and BCW, the first strong band, localized at ~530 eV, is surely related to the main $\text{RE}^{4+}/\text{W}^{4+}$ phase having ionic difference $\Delta Z=0$, hence the lowest polarization and binding energy. Formation of the second, broader band at higher positions fits more +8 ionic $\text{Re}^{3+}/\text{W}^{5+}$ ‘couple’ with $\Delta Z=2$. Such disfigured shape could accommodate for the amount of proportionally stored interstitially oxygen that would not only distort the structure but also introduce more oxidized tungsten ions – mainly responsible for W^{5+} appearance, which actually also limits (the visibility of) this exotic phase by sheer deficiency of that ion. That holds especially true for BPW considering future data presented at EPR vs. PL chapters. Overall, the same relations could be told for both derivatives.^{267,268} Both modes relate to present RE-O-W bridges with charges balanced around 9+ BB’ site total balance. Signals are localized at slightly higher energies (532–535 eV) since ΔZ also grew from 0÷2 to ($\text{Ce}^{4+}/\text{W}^{5+}$) 1÷3 ($\text{Ce}^{3+}/\text{W}^{6+}$). Major XPS signals related to W^{6+} phases probably also host oxygen related to trapped or symmetry-breaking oxygen defects (V_o) - see FWHM’s, precise positions and final stoichiometric formulae in *Table 10* with insets of *Figure 21b* for reference.^{269–272}

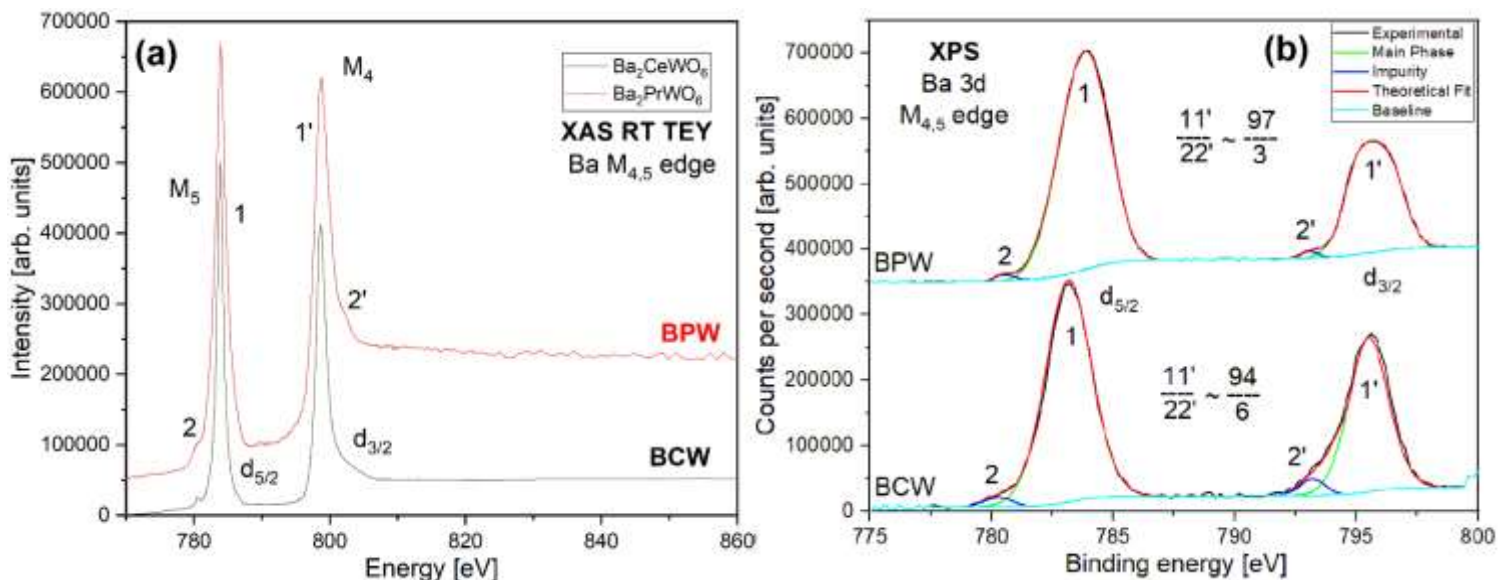


Figure 22 Barium (a) XAS M-edge, and (b) XPS 3d spectra collected at ambient conditions showcasing main (1) phase and secondary (2) phases related most likely to impurities such as BaO (in BCW) or BPWO (in BPW) in double perovskites. The strange, non-symmetric occurrence of both shoulder bands in XAS vs. XPS spectra is worth noting for yet unknown reasons.

BCW and BPW barium absorption edges with 3d shell spectra - low energy 5/2 (M_5) & 3/2 (M_4) localized higher - are presented in *Figure 22*. Both are quite straightforward in interpretation and possess one main peak, related to the primary phase, accompanied by a small shoulder designated to impurities – to BaO and BaPr₂WO₇ (BPWO) in BCW and BPW, respectively.^{273–275} While considering their suitable match with the content of XRD diffractograms, a strange, non-symmetric occurrence of both 22' shoulder bands should be noted in XAS – *Figure 22a*. Their spectra are mirrored in contrast to typical, repeatable XPS configuration for yet unknown reasons - just like in our previous work.¹⁸³ BPW peaks are separated by subsequently ~ 3.3 and 2.3 eV gaps from the main peaks.^{276,277} In BCW first value slightly decreased to 2.8 eV.¹⁸³ This might advocate the statement that the distorted Ba⁽⁺¹⁾O suboxides trying to fully oxidize (+2) & merge into different species like BPWO or BaCeO_{3-x} due to REO_{2-x} separation during synthesis. However, the main peaks do not fall much behind their referral positions by ± 1.5 eV.^{278,279} All read-out and calculated parameters are gathered with the rest of the XPS data in *Table 10*.

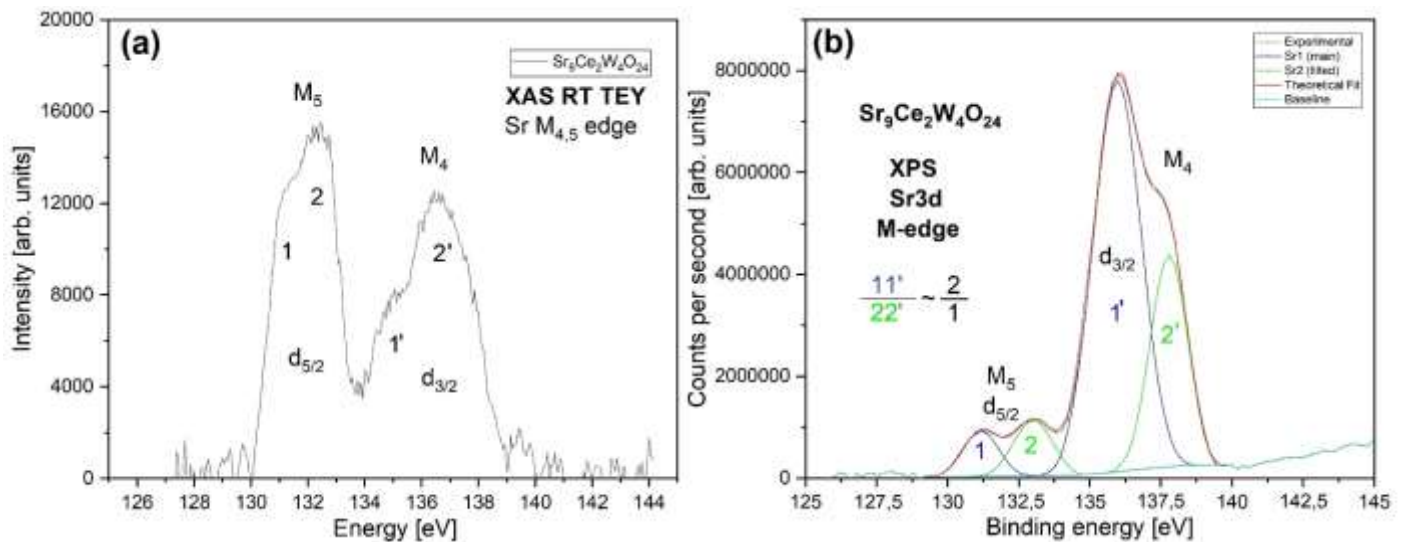


Figure 23 Strontium (a) XAS M-edge, and (b) XPS 3d-shell spectra collected at ambient conditions from SCWO quaternary perovskite featuring the main phase (1) with secondary peak (2) related most likely to symmetry breaking strontium ions escaping their regular sites towards RE positions due to prominent tilting.

Strontium M_{4,5}-edges in *Figure 23a* were tricky to collect on XAS due to the detector's bottom energy limit (~ 100 eV) and high noise-to-signal ratio. However, a quick comparison was made to 3d XPS spectra in *Figure 23b*. After applying a proper baseline and filtering off all other low-energy edges, the same set of signals was separated from the background. This means that XPS would also be more reliable than XAS, but still, based on variable intensity and noise, a strong correlation can be found between each 3d set of M₅ edges. 3d_{5/2} 131.2 & 133.0 eV XPS peaks are positioned analogically to M₄ (3d_{3/2} 136 & 137.6 eV) bands. Compared to the referral positions, the discrepancy between registered signals is around ± 1.0 eV – slightly smaller than for barium.^{278,280–282} Signals' duality is also surely unrelated to impurities since SCWO had no Sr-related minor phases (only W/WC). The other possibility, supported by literature, would be defective Sr⁽⁺¹⁾-O bonding present due to current tilt and unit-cell packing defects caused by symmetry breakage featuring partial substitutions and prominent vacancies.^{256,282,283} They are definitely noticeable quantities to report since the 11'/22' area ratio is close to 2:1. Yet, again, some strange phenomenon has been observed between XAS and XPS spectra – peaks 1,2 and 1',2' have been flipped around regarding the area and intensity ratio for a yet unknown reason -

just like in barium.^{274,278,280,284} The interpretation stays the same, however. Worth adding is a comment about symmetry breakage - here, concern is the small disproportions in the overall stoichiometric formula of this quaternary perovskite which is close to $4\times$ the standard formula of $A_2BB'O_6$.^{270-272,285} Strictly speaking, $Sr_9Ce_2W_4O_{24}$ has some distinct resemblance to $Sr_8Ce_4W_4O_{24}$ which after division (by 4) would give Sr_2CeWO_6 double perovskite. The same statement concerns $Ca_3Ce_2W_2O_{12}$ below when divided by two, would give Ca_2CeWO_6 if not for the unfortunate lack of squeezed-out 1 Ca atom.¹⁸³

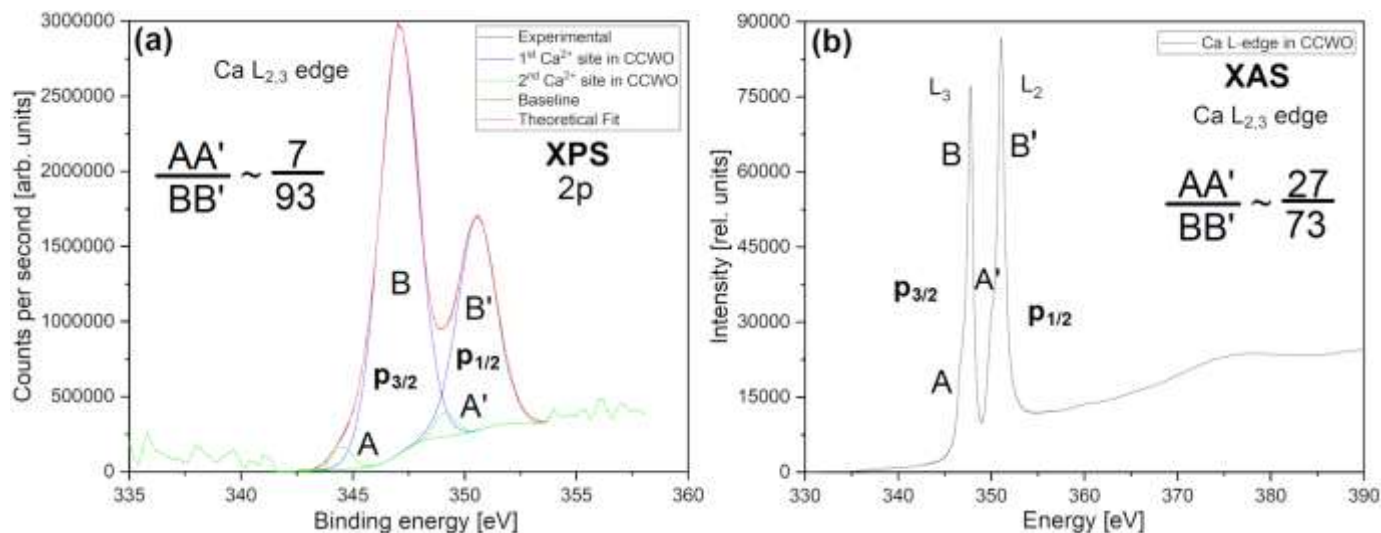


Figure 24 (a) XPS spectra of Ca $2p_{3/2}$ and $2p_{1/2}$ shell within CCWO ilmenite matrix alongside (b) XAS Ca L-edge measurements showing main symmetry-breaking interactions at room conditions and ultra-high vacuum.

Calcium $2p_{3/2}$ and $2p_{1/2}$ levels seem to correspond well to symmetrical $L_{2,3}$ edges in *Figure 24*. The only issue is the intensity/ratio discrepancy between the satellites – in XAS, the area covered by them is much bigger (27:73%) than on XPS (only 7 to 93%). Although the origin of those bands is not a mystery, their disproportionate values can cause confusion, especially knowing that none match the aforementioned Ce^{3+}/Ce^{4+} (59/41%) content – XAS is much more reliable due to deep penetration. Nevertheless, both sets should be interpreted the same way as SCWO: AA' shoulders as tilted, broken symmetry sites with $Ca^{(+1)}$ -O suboxide interactions while dominating BB' bands correspond to the primary phase and its bridging RE-O-Ca-O-W bonds.²⁸⁶⁻²⁸⁹ The issue here is surely not related to impurities like in barium double perovskites since the purity of CCWO is $\sim 99\%$ & less than 1% is related to W/WC. Worth noting is the displacement of the main XAS and XPS bands against the referral values by ± 1.5 eV – this might be an issue related to calibration, but the separation between peaks and their satellites is an issue purely related to chemical interactions and their origin.^{290,291} Here, shoulders are spread apart from their bigger neighbors by approximately 2 eV, which further supports $Ca^{(+1)}$ -O interpretation, but the shape of the peaks (their FWHM) and their individual AA' gap reaching 4.3 eV implies high disorder and a lot of scattered multisites involved in their contribution. Generally, the gaps are more significant than in barium double perovskite edges (which are more ordered structures), so this conclusion is reasonable.^{273,292}

Table 10 Summary of all crucial core level information experimentally investigated on XPS: their content, characteristics, and stoichiometric formulae based on total atomic ratios for both BCW & BPW double perovskites (normalized to W); and the other two SCWO & CCWO derivatives (normalized to Ce).

Material	Core level	s-o split	Position [eV]	FWHM [eV]	Total Atom [%]	Fraction
BCW	W4f	7/2	33.4 ; 34.4	1.65 ; 0.86	8.96	1
		5/2	35.4 ; 36.4	1.13 ; 0.88		
	O1s	---	529.9 ; 531.8	1.78 ; 2.68	61.01	6.81
	Ba3d	5/2	780.4 ; 783.2	2.21 ; 1.47	19.46	2.17
		3/2	793.2 ; 795.5	1.16 ; 1.88		
	Ce3d	5/2	880.6;883.1	1.08;2.20	10.57	1.18
			886.2;889.6;896.0	2.56;3.64;1.03		
		3/2	898.6;900.8 904.0;906.6;917.3	1.91;2.63 2.31;3.09;2.51		
BPW	W4f	7/2	33.1 ; 34.6	1.76 ; 0.62	8.69	1
		5/2	35.4 ; 36.9	1.89 ; 1.13		
	O1s	---	530.2 ; 532.6	2.47 ; 1.05	62.95	7.24
	Ba3d	5/2	780.5 ; 783.9	0.89 ; 2.41	18.15	2.09
		3/2	793.4 ; 795.7	0.79 ; 3.04		
	Pr3d	5/2	928.8;930.5;933.6	1.32;1.73;1.04	10.21	1.17
			936.2;947.5	1.61;2.91		
		3/2	949.9;952.4;954.4 957.0; 964.8	1.26;1.78;0.98 1.52;1.42		
SCWO	W4f	7/2	34.7 ; 36.5	1.47 ; 1.89	9.49	3.82
		5/2	37.3 ; 38.5	0.90 ; 1.25		
	Sr3d	5/2	131.2 ; 133.0	1.45 ; 1.60	22.91	9.22
		3/2	136.0 ; 137.6	1.98 ; 1.55		
	O1s	---	531.0 ; 534.4	3.03 ; 3.01	62.63	25.20
	Ce3d	5/2	880.4;882.7;885.1	2.28;1.39;2.52	4.97	2
			887.8;896.4	1.68;1.80		
		3/2	899.5;901.6;904.6 907.8;917.2	2.15;1.72;2.18 1.62;1.92		
CCWO	W4f	7/2	34.6 ; 36.4	1.27 ; 1.96	9.56	1.92
		5/2	36.5 ; 38.1	1.39 ; 1.76		
	Ca2p	3/2	345.8 ; 347.2	1.23 ; 2.00	15.56	3.12
		1/2	349.2 ; 350.7	1.03 ; 1.94		
	O1s	---	531.4 ; 532.6	1.59 ; 2.60	64.92	13.04
	Ce3d	5/2	880.5;883.4	3.53;3.66	9.96	2
			885.8;889.0;896.8	2.26;2.85;3.17		
		3/2	899.8;901.8;904.8 907.6;916.4	3.22;3.72;3.02 3.07;2.84		

After extensive data gathering and qualitative analysis of provided atomic shells and edges, one can pursue quantitative calculations based on *Table 10* to estimate the actual stoichiometric proportions of all our compounds and confront them with references from the XRD database. It is a well-known fact that there are always some discrepancies between ideal, isostructural formula, and realistic composition, but the proper ratios need to be

established in order to: push the overall work onward; evaluate the product quality, arrangement, and mineral assignment; check for eventual mishaps and plan optimizing procedures for future reference. So in the case of barium double perovskites, both rare-earth tungstates were normalized to their respective d-block tungsten. The idea is to always choose a central, most crucial ion or an atom closest to the least common multiple. In that case, both criteria fit W, so the total atomic ratios calculated from XPS spectra turned out to be $\text{Ba}_{2.17}\text{Ce}_{1.18}\text{WO}_{6.81}$ and $\text{Ba}_{2.09}\text{Pr}_{1.17}\text{WO}_{7.24}$ in comparison to idealistic Ba_2CeWO_6 (BCW) and Ba_2PrWO_6 (BPW). These values are not far from the goal considering slow WO_2 evaporation (amended by $\sim 2\%$ surplus of the substrate). There is undoubtedly more Ba due to BaO and BPWO impurities involved in messing around the perfect deconvolution alongside REO_{2-x} species. If gathered together, they could support the case of creating distorted (single) perovskites like $\text{Ba}_{0.17}\text{Ce}_{0.09}\text{O}_{0.51}$, but that is just a long shot. The remaining CeO_{2-x} and Pr_2O_3 phases would still exist. More critical is the surplus of oxygen, which signifies a possibility of interstitial oxygen traps and vacancies that could be responsible for slight oxidation experienced with $\text{W}^{4+} \rightarrow \text{W}^{5/6+}$ during synthesis despite using inert Ar gas (with a slight addition of 1-5% of H_2). This will influence band gaps, future PL measurements, and energy transfer properties if not attended to.

Derivatives also have more oxygen - not only ideal but in absolute configuration. Since they have more W in their configuration (SCWO ~ 4), the least common multiple for jointly present ions would be 2. for Cerium, so that was the chosen basis of multiplication and division. As can be deduced from the lower half of *Table 10*, $\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$ (SCWO) turned out to be $\text{Sr}_{9.22}\text{Ce}_2\text{W}_{3.82}\text{O}_{25.2}$ and $\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$ (CCWO) $\sim \text{Ca}_{3.12}\text{Ce}_2\text{W}_{1.92}\text{O}_{13.04}$. Tungsten deficiency has been noted (for the reasons already stated before). The only difference would be Ca and Sr, which surplus could not be accommodated by any impurities. Their abundance would presumably influence symmetry-breaking interactions and the eventual occurrence of vacancies/half-occupied sites by tilting those giant, densely-packed unit cells.^{256,269-272,285}

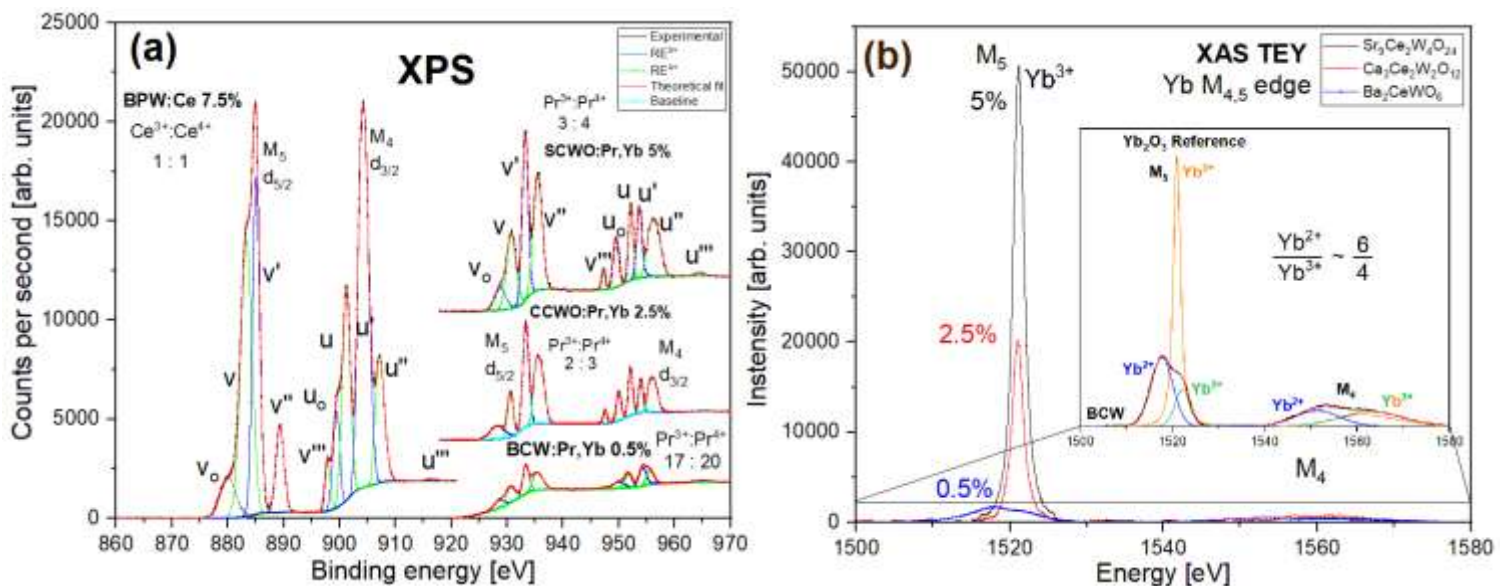


Figure 25 Room temperature X-ray spectra for all doped materials with selectively chosen RE ions: (a) XPS focuses on Ce/Pr 3d shell, while (b) XAS on Yb. Insets provide additional data about content features such as deconvoluted peaks or eventual ratios compared to some RE-oxide references.

A brief comment should be made on the dopants since, till now, all discussion focused solely on pure matrices. Firstly, they were doped evenly, so every stated percentage amounts concern all cations separately but equally. Secondly, both graphs in *Figure 25* are crucial in determining the factual content of PL-active dopant concentration after introducing them to the matrices. Chosen amounts of RE³⁺ cations substituted other RE oxides in various degrees since they possess unstable nature and usually are mixed with some other, differently charged atoms.^{169,171,195,196,199,200,292} That is why the true values of excitable centers are always lower than expected - it is good to account for that loss or just accept the benefit of lower PL quenching factors. Thirdly, the quantitative assessment of such phenomena might help in predicting future emission and (Ce,Pr NUV sensitizers → NIR Yb activators) energy transfer properties, not to mention damage control (eventual influence from defects).²⁹³⁻²⁹⁸ Such research data from neighboring edges of: (a) energy emitters like Pr doped cerates or Ce-hosting praseodymates;^{220-231,240,299} and (b) energy-accepting Yb ions,^{300,301} were gathered in *Table 11* below. Fourthly, the interpretation is consistent and follows the analysis & conclusions presented previously for constituent Pr/Ce ions.

Table 11 XPS deconvolution data sampling actual content of PL-active RE dopants stored within investigated phosphors previously mentioned in *Figure 25*.

Abbreviation	Ideally planned stoichiometric formula	Total percentage of each introduced dopant [%]	Sensitizer ratio (Ce/Pr) ³⁺ / (Ce/Pr) ⁴⁺	Real sensitizer RE ³⁺ content [%]	Activator Yb ²⁺ / Yb ³⁺ ratio	Real activator Yb ³⁺ content [%]
BCW:Pr,Yb	Ba ₂ Ce _{0.99} Pr _{0.005} Yb _{0.005} O ₆	0.5	17:20	0.23	6:4	0.20
CCWO:Pr,Yb	Ca ₃ Ce _{1.9} Pr _{0.05} Yb _{0.05} W ₂ O ₁₂	2.5	2:3	1.00	0:1	5.00
SCWO:Pr,Yb	Sr ₉ Ce _{1.8} Pr _{0.1} Yb _{0.1} W ₄ O ₂₄	5.0	3:4	2.14	0:1	2.50
BPW:Ce	Ba ₂ Pr _{0.925} Ce _{0.075} WO ₆	7.5	1:1	3.75	---	---

An unusual disproportion was noticed for Yb. Just like W⁵⁺, a less common ion in the form of Yb²⁺ was found in BCW. Unintentionally and unbeknownst to researchers who just added Yb₂O₃ (presented in the inset of *Figure 25b*), this discovery was surprisingly intriguing. The origin is probably the same as the appearance of W⁵⁺, but here some unintentional charge transfer must have occurred between B-site or maybe partially from evacuated O²⁻ ions during SSR creating residual, interstitial oxygen while kept in Ar:H₂ atmosphere. It gives an unprecedented opportunity to look into this nontrivially charged element; however, Yb³⁺ is an ion with which we will be most invested since the goal was to create an efficient downconverter emitting white light or NIR radiation. Furthermore, unfortunately, Yb²⁺ has [XE] 4f¹⁴ 6s² electron configuration, which is optically inactive.³⁰²⁻³⁰⁵ This means that we will eventually try to switch it back to Yb³⁺ (through HT annealing or light-induced reverse charge transfer) to improve the odds of emission by increasing the receiver's concentration, thus improving the photon's absorption cross-section.^{3,163,306,307}

2.4 Raman and FTIR spectroscopy

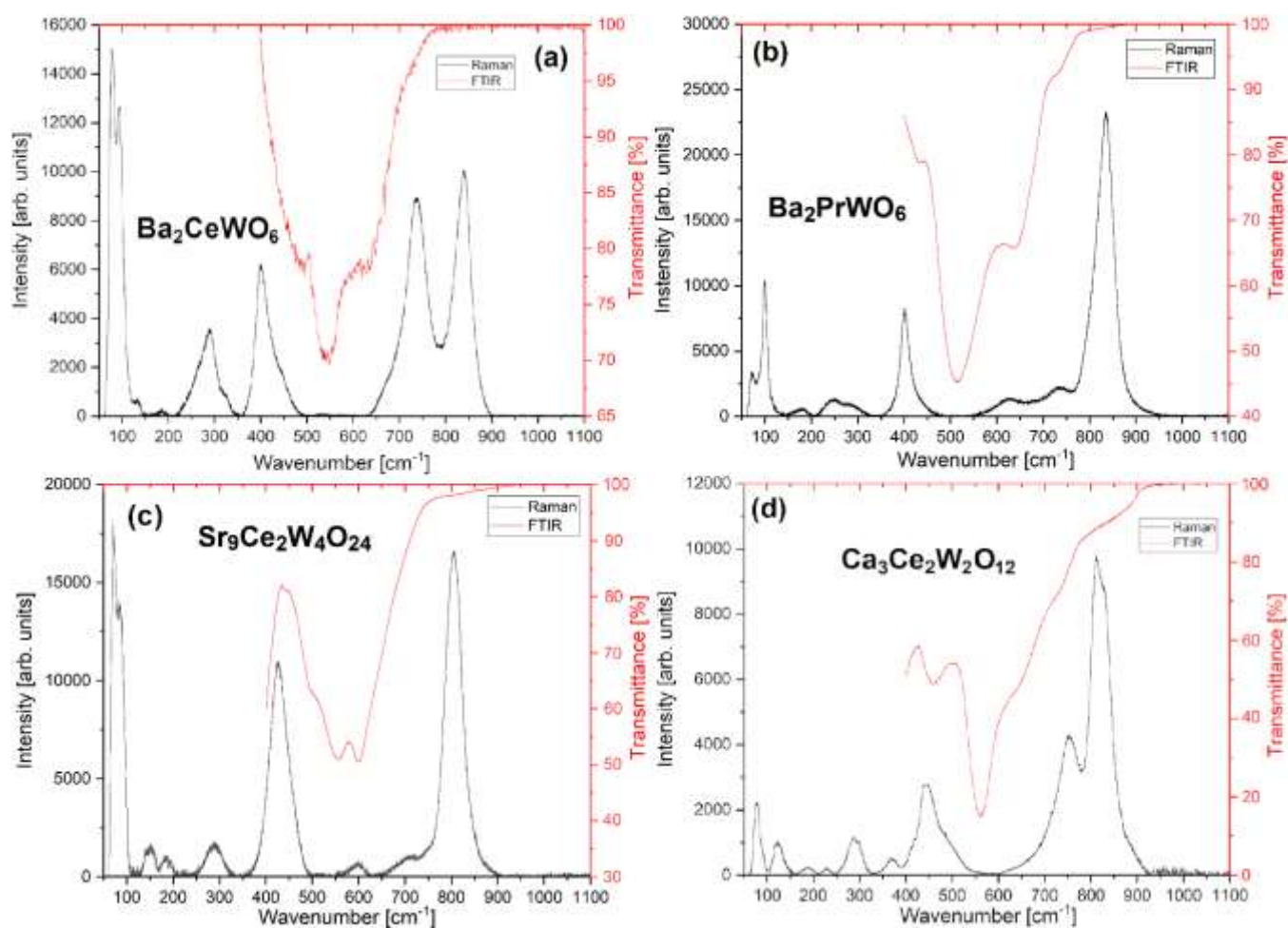


Figure 26 Joined Raman (down-black) and FTIR (up-red) spectra of undoped (a) BCW, (b) BPW double perovskites, and two other (c) SCWO, (d) CCWO derivatives. Data were collected at ambient room conditions in the region of interest - no additional signals were detected between 1100 cm^{-1} and 4000 cm^{-1} .

Techniques such as FTIR and (more importantly) Raman spectroscopies were involved in clearing up any doubts originating from XRD SG selection. Knowing their complementary factor in analysis through group theory calculations, one could support crystallographical findings by matching the number of currently visible modes to assigned phases in various conditions (ambient vs. HP, HT, LT changes). However, a fundamental base must be established to evaluate a well-described fingerprint region for upcoming polymorphism studies. Sometimes this is enough to clear any initial discrepancies (unfortunately, not here).

Based solely on XRD data, a preemptive selection of SG has been made in *section 2.2*. More ordered BCW and BPW double perovskites were assigned to slightly tilted cuboid (monoclinic β angle) $I2/m$. Derivatives like SCWO quaternary perovskite and CCWO ilmenite to $I4_1/a$ and $R-3c$, respectively. The results of group theory calculations for these choices are presented in *Table 12*. Factors for alternatives like $R-3$ (BCW, BPW), $Fm-3m$ (SCWO), and $R3c$ (CCWO) were gathered in supplementary *Table A3*. Together, all the data can be confronted with the experimental results shown in *Figure 26*.

Table 12 Factor group analysis for all mainly chosen space groups of discussed tungstates.

Material (SG)	Ion	Wyckoff Positions	Point Symmetry	Red. Representation
BCW	A ²⁺ (Ba)	4i	C _s	2A _{1g} +A _u +B _g +2B _u
BPW	B ^{3+/4+} (Ce/Pr)	2a	C _{2h}	A _u +2B _u
	B ^{4+/5+} (W)	2d	C _{2h}	A _u +2B _u
I2/m (C_{2h})	O ₁ ²⁻	4i	C _s	2A _{1g} +A _u +B _g +2B _u
	O ₂ ²⁻	8j	C ₁	3A _{1g} +3A _u +3B _g +3B _u
$\Gamma_{\text{TOTAL}} = 7A_{1g}+7A_u+5B_g+11B_u$; $\Gamma_{\text{ACUSTIC}} = A_u+2B_u$; $\Gamma_{\text{IR}} = 6A_u+9B_u$; $\Gamma_{\text{Raman}} = 7A_{1g}+5B_g$				
SCWO	A ²⁺ (Sr)	8e, 16f	C ₂ , C ₁	4A _g +4A _u +4B _g +5 ¹ E _g +5 ¹ E _u +5 ² E _g +5 ² E _u
	B ^{3+/4+} (Ce)	8e, 16f	C ₂ , C ₁	4A _g +4A _u +4B _g +5 ¹ E _g +5 ¹ E _u +5 ² E _g +5 ² E _u
I4₁/a (C_{4h})	B ^{6/5+} (W)	8c, 8d	C _i	6A _u +6 ¹ E _u +6 ² E _u
	O ²⁻	16f	C ₁	3A _g +3A _u +3B _g +3 ¹ E _g +3 ¹ E _u +3 ² E _g +3 ² E _u
$\Gamma_{\text{TOTAL}} = 11A_g+17A_u+11B_g+13^1E_g+19^1E_u+13^2E_g+19^2E_u$; $\Gamma_{\text{ACUSTIC}} = A_u+^1E_u+^2E_u$; $\Gamma_{\text{IR}} = 16A_u+18^1E_u+18^2E_u$; $\Gamma_{\text{Raman}} = 11A_g+11B_g+13^1E_g+13^2E_g$				
CCWO	A ²⁺ (Ca)	2a, 4c, 12f	C ₁ , C ₃ , D ₃	4A _{1g} +5A _{2u} +9E _u +9E _g
	B ^{3+/4+} (Ce)	4c, 12f	C ₁ , C ₃	4A _{1g} +4A _{2u} +8E _u +8E _g
R-3c (D_{3d})	B ^{6/5+} (W)	6e, 4c, 2b	C ₂ , C ₃ , C _{3i} ,	2A _{1g} +4A _{2u} +7E _u +5E _g
	O ²⁻	12f	C ₁	3A _{1g} +3A _{2u} +6E _u +6E _g
$\Gamma_{\text{TOTAL}} = 13A_{1g}+28E_g+16A_{2u}+30E_u$; $\Gamma_{\text{ACUSTIC}} = A_{2u}+E_u$; $\Gamma_{\text{IR}} = 15A_{2u}+29E_u$; $\Gamma_{\text{Raman}} = 13A_{1g}+28E_g$				

Unfortunately, both barium double perovskites have space groups hosting coincidentally the same number of modes for Raman – 12 (*I2/m* 7A_{1g}+5B_g or 4A_g+4¹E_g+4²E_g for *R-3*). Furthermore, even if a befitting number of peaks was an issue (usually +2÷4 more counting the defects and probable impurities), all spectra seem eerily identical considering shape and form. Even comparing with derivatives bearing many different stoichiometric formulae. FWHM is considerably large, ~36÷50 cm⁻¹ wide. Those problematic similarities could be attributed to various disordered (Ce,Pr)^{3+/4+}/W^{4/5/6+} ion ‘pairings’ manifesting themselves in indigenous multisites – different charges, V_o defects, scattered throughout the matrix contributing to a specific stretch or bending. All materials harbor some kind of BB’-site ionic balance – BPW & BCW 8+, SCWO, and CCWO 9+. Keeping this order in mind, careful theoretical interpretation denoted with unique modal characteristics were gathered in *Tables 13 & 14* with support from the literature.^{308–313}

Table 13 Conjoined experimental Raman data collected at ambient conditions with literature-verified phonon assignments emphasizing observable differences and similarities for both BCW and BPW double perovskites.^{308–313} Asterisks mark modes that could also be associated with REO_{2-x} impurities according to previous XRD results & some theoretical energy-scattering calculations.^{321,324}

Wavenumber ω_0 [cm ⁻¹]		Signal assignment & Characteristics	
Ba ₂ CeWO ₆	Ba ₂ PrWO ₆		
78.6	72.2	BaO ₆ multisite complex lattice deformations at A-sites of mixed B _g and ¹ E _g / ² E _g origin signifying disorder & divergence from a perfect cuboid structure.	BCW s, sh BPW mw, sh, left shoulder
96.5	99.5		BCW s, sh BPW m, sh
134.1	163.1		BCW w, sh, right shoulder BPW w, br
182.8	179.8	In-plane σ_p & out-of-plane σ_τ bending in REO ₆ and WO ₆ clusters of mixed B _g and ¹ E _g / ² E _g origin between shared and not shared B-sites of various 3/4/5+ charges.	BCW w, br BPW w, br
231.2	247.4		BCW, w, sh, left shoulder BPW mw, br
267.0*	---		mw, sh, left shoulder
288.3	280.0	*possible REO_{2-x} O²⁻ vacancy (V_o) interference	BCW m, sh BPW mw, br
320.3	315.7	In-plane σ_{sc} & out-of-plane σ_ω bending in REO ₆ and WO ₆ clusters of mixed B _g And ¹ E _g origin between shared and not shared B-site cations of various charges.	BCW mw, sh, right shoulder BPW w, br
---	367.1		BPW w, br, left shoulder
400.1	401.2		BCW m, sh BPW m, sh
441.9	434.8	Distorted impurity of v_{sym} REO₂, (F_{2g})	BCW mw, sh, right shoulder BPW mw, sh, left shoulder
459.9	458.2		mw, sh, right shoulder
532.5*	564.4*		w, br
576*	625.8*	Transverse REO ₆ and WO ₆ motion of mixed A _{1g} , and A _g character between shared and not shared B-sites of cations having different charge / *REO_{2-x} complex lattice V_o distortions (RE³⁺:RE⁴⁺)O₇.	BCW w, br BPW mw, br
667.2	678.7	v _{asym} & v _{sym} stretches of REO ₆ & WO ₆ clusters of A _{1g} and A _g for various Wyckoff sites not only partially occupied but also hosting B-site RE/W ions bearing different charges.	BCW mw, sh, left shoulder BPW mw, br, left shoulder
733.5	734.5		BCW s, sh BPW mw, br
814.5	791.3		BCW m, sh, left shoulder BPW m, sh, left shoulder
838.9	835.1	---	BCW s, sh BPW s, sh
---	851.6		BPW s, sh, right shoulder

w-weak; m-medium; s-strong; sh-sharp; br-broad; v_{sym} – symmetric stretching; v_{asym} – asymmetric stretching; σ_{sc} – scissoring deformation; σ_ω – wagging deformation; σ_τ – twisting deformation; σ_p – rocking deformation; V_o – oxygen vacancy; RE – neighboring rare-earth (Pr/Ce) oxide clusters

As for derivatives, they both have way too few peaks to be simply assigned to either phase: SCWO tetragonal group has 48 (A_g+B_g+¹E_g+²E_g) modes in total while cubic 25 (A_{1g}+E_g+T_{2g}). The difference is less severe for the latter group (by 8 than 31), yet still, it does not explicitly clarify anything about the other SG. Accounting for all the strontium-exclusive sites being half-occupied or others riddled with vacancies after partial substitution with Ce^{3/4+}, that maybe should not surprise. This statement is also true for CCWO, where more suitable noncentrosymmetric SG hosting 21 (A₁+E) peaks is pointed against the chosen R-3c with 41 (A_{1g}+E_g) modes. Without NPD measurements, only polymorphism measurements outside of ambient conditions could help.^{71,72}

Table 14 Conjoined experimental Raman data collected at ambient conditions with literature-verified phonon assignments emphasizing both observable differences and similarities in SCWO and CCWO perovskite derivatives.^{308–313}

Wavenumber ω_0 [cm^{-1}]		Signal assignment & Characteristics	
$\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$	$\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$		
71.1	81.3	SCWO s, sh CCWO m,sh	
85.0	92.4	CaO ₆ & SrO ₆ multisite complex lattice deformations at A-site of mixed B _g /E _g origin signifying disorder & divergence from a perfect cuboid structure.	
141.5	123.5		SCWO w, sh, left shoulder CCWO mw, sh
150.9	---		SCWO mw, br
186.5	186.3	SCWO w, br CCWO w, br	
---	228.3	w, br	
---	261.7	In-plane σ_p & out-of-plane σ_τ bending in CeO ₆ and WO ₆ clusters of mixed B _g /E _g origin between shared and not shared B-sites of various 3/4/5/6+ charges.	
289.3	288.6		w, br, right shoulder SCWO mw, br CCWO mw, sh
---	297.7		mw, sh
391.3	370.8	SCWO mw, sh, left shoulder CCWO mw, sh	
411.4	414.5	In-plane σ_{sc} & out-of-plane σ_ω bending in CeO ₆ and WO ₆ clusters of mixed B _g /E _g origin between shared and not shared B-site cations of various charges.	
426.0	445		SCWO m, sh, left shoulder CCWO mw, sh, left shoulder
---	499.3		SCWO m, sh CCWO m, sh
557.6	---	mw, sh, right shoulder	
598.7	626.7	Transverse CeO ₆ and WO ₆ motion of mixed A _{1g} /E _g character between shared and not shared B-sites of cations having different charges.	
710.9	710.8		w, br, left shoulder SCWO w, br CCWO w, br, right shoulder
749.2	753.2		SCWO mw, br CCWO mw, sh, left shoulder
784.7	812.8	v _{asym} & v _{sym} stretches of CeO ₆ & WO ₆ clusters of A ₁ and A _{1g} for various Wyckoff sites not only partially occupied but also hosting B-site RE/W ions bearing different charges.	
805.6	828.3		SCWO mw, sh, left shoulder CCWO m, sh
844.2	888.1		SCWO m, sh, left shoulder CCWO s, sh
		SCWO s, sh CCWO s, sh	
		SCWO mw, br, right shoulder CCWO mw, sh, right shoulder	

w-weak; m-medium; s-strong; sh-sharp; br-broad; v_{sym} – symmetric stretching; v_{asym} – asymmetric stretching; σ_{sc} – scissoring deformation; σ_ω – wagging deformation; σ_τ – twisting deformation; σ_p – rocking deformation;

FTIR did not provide much support either, given it showed not only that the signals were much broader (+64 cm^{-1}) FWHM but also that many of the B_u/E_u modes were cut off below 400 cm^{-1} limit (considering filters and air interference). Data in a similarly ordered fashion is presented in *Table 15*, giving interpretation support from the literature.^{313–319}

Table 15 Conjoined experimental data and literature phonon assignment collected at ambient conditions using confocal, reflective FTIR spectroscopy for all investigated compounds to emphasize observable differences and similarities.^{313–319}

BCW	Wavenumber ω_0 [cm^{-1}]			Characteristics	Assignment
	BPW	SCWO	CCWO		
	435	445	459	m, sh, left shoulder	In- and out-of-plane bending WO_6 bands and partially substituted REO_6 clusters of $E_u/{}^2E_u$ origin down to 100 cm^{-1}
476		497		m, sh, left shoulder	
554	514.5	554	560	s, sh	ν_{sym} stretch of both WO_6 , and partially substituted REO_6 clusters of A_u/A_{2u} origin.
		599		s, sh	
632	639	648.5		m, sh, right shoulder	
	731.5	797	722	mw, sh, right shoulder	ν_{asym} stretch of both WO_6 , and partially substituted REO_6 clusters of A_u/A_{2u} origin.
	834	874	855	w, br	

w-weak; m-medium; s-strong; sh-sharp; br-broad; ν_{sym} – symmetric stretching; ν_{asym} – asymmetric stretching; RE – neighboring rare-earth (Pr/Ce) oxide clusters

Worth mentioning are a few peaks of debatable origin or modes clearly related to impurities, such as not-negligible shoulder present at $\sim 460 \text{ cm}^{-1}$ in most double perovskites. That band is often denoted as stretching mode F_{2g} of fully oxidized CeO_2 .^{320,321} Assuming PrO_2 would be slightly heavier, appearance should follow shortly after at lower wavenumbers. Another sight to monitor is a weak set of broad peaks occasionally appearing near $550 - 600 \text{ cm}^{-1}$ – visible as hashtags in zoomed insets alongside other remarked peaks in *Figure A11*. Some articles refer collectively to them as deficient CeO_{2-x} V_o modes where $0 < x < 0.5$ could cause additional lattice deformations around $230 \div 250 \text{ cm}^{-1}$.^{321–324} Here, most likely PrO_{2-x} analogs of Ce_7O_{12} or Ce_3O_5 like Pr_3O_5 or Pr_6O_{11} could also occur.^{325–327} This supports the statement that Raman spectroscopy is truly a surface-based analytic tool (together with XRD diffractograms and manual examination of pellets confirm REO_{2-x} influence), but the contribution from the actual structure could not be fully discarded here. The appearance of those modes could account for the same differently-charged Pr/Ce-O-W/Ba bridges forming inside the rich-in-defects lattice.¹⁸³ At this point, only further HP & LT Raman studies could tell. Finally, one should critically avoid BaWO_4 easily emerging from the matrix if an excessive amount of oxygen (atmospheric or interstitial) is present (marked as asterisks in *Figure A11a*). Not only does this tungstate have very strong scattering signals, but it also has prominent PL as a scintillator promoting by its quantum efficiency false results covering the true, more subtle ones from the originally developed matrices.^{2,202,203,328}

3. Pressure and temperature studies of polymorphism and stability

3.1 Raman pressure vs. temperature dependencies

In search of possible (indirect) SG confirmation via the eventual track of phase-transitions at various conditions, as documented in the vast perovskite literature,^{37,71,72} an approach was taken to investigate polymorphism using all on-sight available assets. Given the limitations of accessible techniques (XRD, XPS), more detailed results could be obtained with the use of European synchrotron beamlines (like HP, LT XRD & NPD - they would help to assess not only the nature of eventual changes but also establish a set of bulk moduli with respectable Grüneisen parameters)³²⁹⁻³³² the only resort left was to use the technique that our scientific team specialized in – Raman spectroscopy incorporating cryostats and DAC. Sparked out of curiosity, this first small research out of RT boundaries led to an extensive investigation elaborating on the unusual behaviorism of the samples conjuring more non-trivial questions with time (about stability and sensory capabilities featuring charge transfer, oxidization, and photobleaching phenomena) which will be answered at the conclusions after all the evidence presented in this (*section 3*) in upcoming (*section 4*) chapters.

Starting from HP, all samples were submerged in quasi hydrostatic Ar PTM, creating soft, butter-like jelly when compressed under nitrogen-level temperatures up to ~15 GPa.^{333,334} Together with all surrounding air and moisture, a mild interference in the form of weak but sharp signals could be seen in the range up to 140 cm⁻¹ – see supplemental *Figure A12* for reference. Those modes do not shift or change their intensity during the experiment and are usually easily distinguished and filtered out by the software if visible.^{335,336} That is because they usually are swallowed by good scattering responses from the investigated materials. The initial goal was to reach at least 20 GPa – at that point, 0.45 mm diamonds would withstand the pressure with inert Ar inside without risk of breaking, and the samples would probably not be exploited above that threshold. Other, more compressible, and hydrostatic PTMs like oils and alcohol mixtures affect double perovskite structures to some degree (intercalary amorphization), so this option was omitted until absolutely necessary.^{337,338}

Barium double perovskites were the first to be investigated. Their compression and decompression cascades are presented in *Figure 27*. As arrows guide the eye, nothing remarkable happens – typical slow red-shift of the bands (previously registered and categorized at RT in *Figure 26ab* & *Table 13*) is observed at various rates with gradual amorphization, but no phase transition occurs.^{339,340} Some of them are so weak they fade and reappear spontaneously in the spectra – that concerns, especially those debatable ones related to CeO_{2-x} defects (featured in *Figure A11*).³²³⁻³²⁶ Wavenumber vs. pressure dependencies and all accountable coefficients are presented for BCW & BPW in *Figure 28* and *Table 16*, respectively. *Fm-3m* SG would transform several times by now (up to 20 GPa) compared to *I2/m* or *R-3*, the monoclinic tilt would be unquestionably apparent.^{37,71,72,341-343} Higher transition thresholds could not be disqualified for the other two; however,³⁴⁴⁻³⁴⁶ both materials seem quite stable in these conditions as hysteresis during release is not higher than 2 GPa.¹⁸³ Nevertheless, the experiments must have been halted at some point since, as predicted, the Ni750 gasket would not withstand such elevated pressures – average deformations from the original indentation state can be viewed in appendix *Figure A13*.

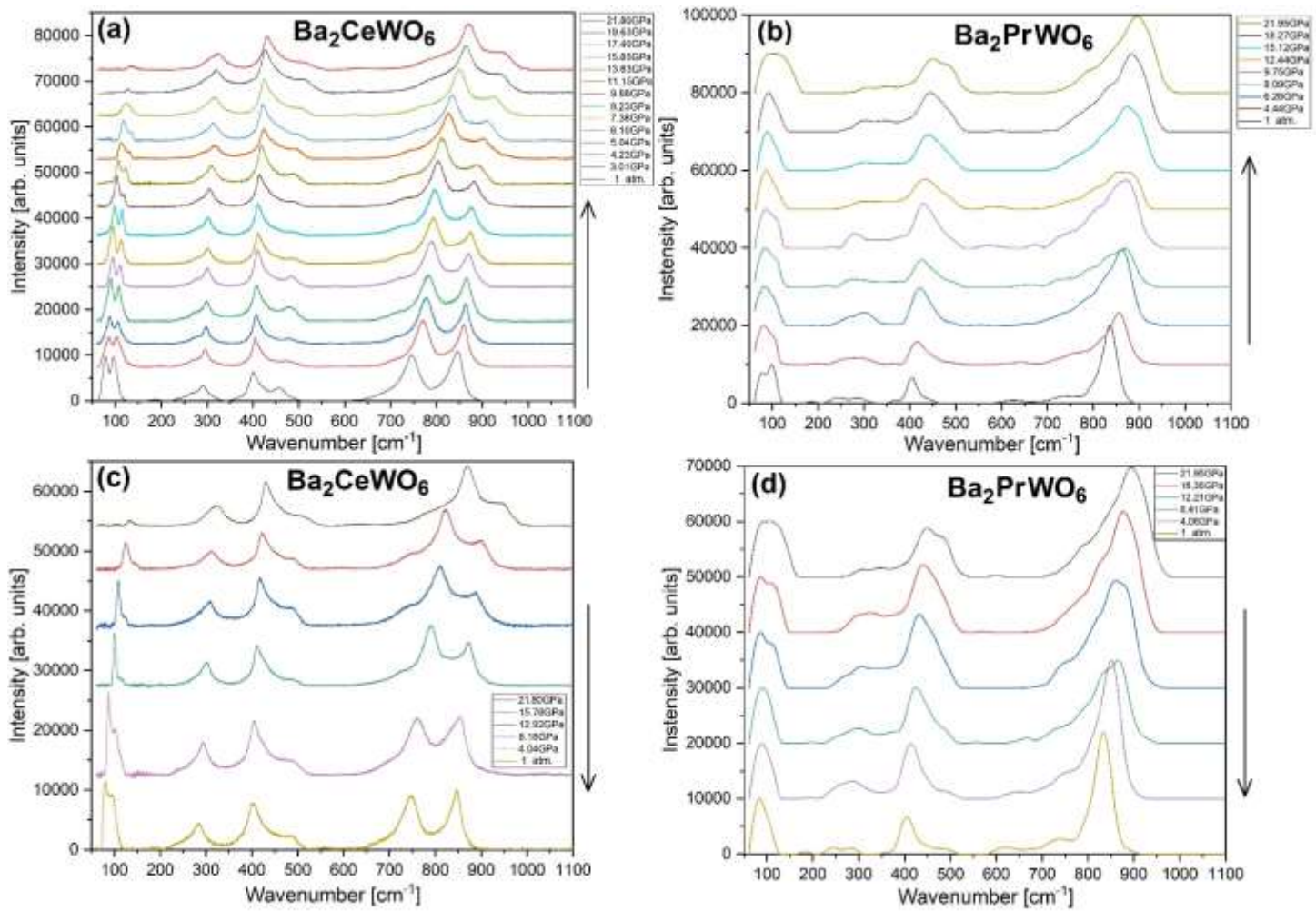


Figure 27 Raman spectra collected under hydrostatic argon PTM while using consecutive compression (top) – decompression (bottom) cycles for barium double perovskites BCW (left), and BPW (right). Arrows guide the eye. Despite apparent, slow amorphization accompanied by red-shifting peaks no phase transitions were denoted.

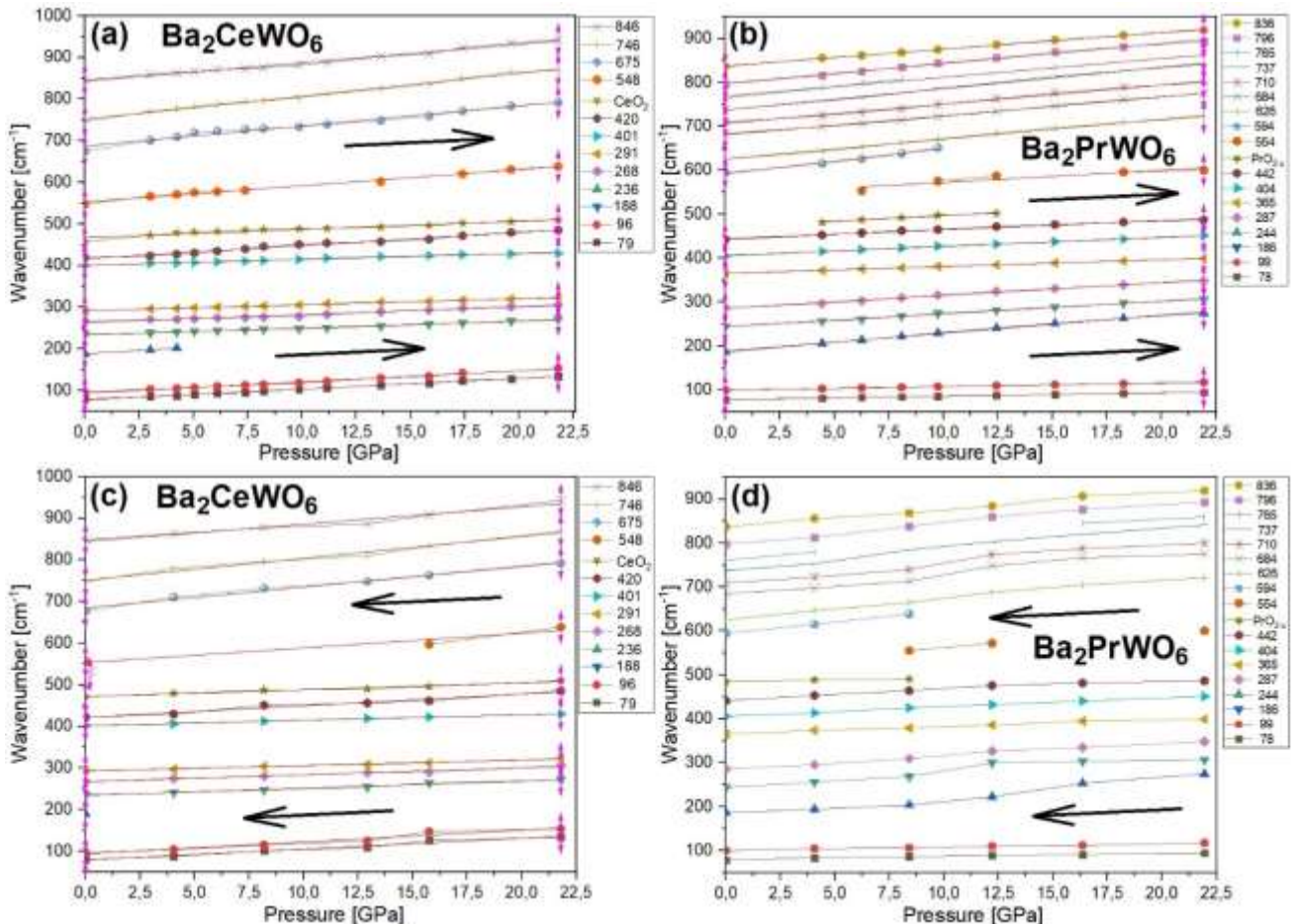


Figure 28 Raman peak position vs. pressure relations for double perovskite modes showcased in previous *Figure 27*. Presented similarly, compression is shown on the top, while decompression dependencies are on the bottom. BCW to the left, BPW to the right. Inserted right and left arrows guide the eye. Pink anchors mark the areas of linear fits. Error bars are too small to be legible.

Table 16 Room temperature wavenumbers vs. pressure coefficients of Raman lines presented previously in *Figure 28* compression cycles for both BCW and BPW double perovskites.

Ba₂CeWO₆			Ba₂PrWO₆		
ω_0 [cm ⁻¹]	$d\omega/dP$ [cm ⁻¹ /GPa]	R ²	ω_0 [cm ⁻¹]	$d\omega/dP$ [cm ⁻¹ /GPa]	R ²
79	2.57 ± 0.04	0.99	78	0.76 ± 0.01	0.99
96	2.66 ± 0.07	0.99	99	0.82 ± 0.01	0.99
188	3.27 ± 0.17	0.99	186	4.08 ± 0.09	0.99
236	1.65 ± 0.05	0.98	244	2.89 ± 0.06	0.99
268	1.87 ± 0.10	0.96	287	2.91 ± 0.09	0.99
291	1.45 ± 0.03	0.99	365	1.55 ± 0.02	0.99
401	1.40 ± 0.05	0.98	404	2.07 ± 0.04	0.99
420	3.15 ± 0.11	0.98	442	2.06 ± 0.08	0.98
548	3.90 ± 0.13	0.99	554	2.77 ± 0.61	0.83
			594	5.74 ± 0.40	0.98
			626	4.52 ± 0.13	0.99
675	4.82 ± 0.22	0.97	684	4.30 ± 0.09	0.99
			710	4.35 ± 0.12	0.99
746	5.52 ± 0.14	0.99	737	4.86 ± 0.10	0.99
			765	4.26 ± 0.21	0.98
			796	4.54 ± 0.12	0.99
846	4.45 ± 0.16	0.98	836	3.80 ± 0.06	0.99

On the other hand, both tungstate derivatives (SCWO & CCWO) show more noticeable changes at HP, as seen in *Figure 29*. A set of new peaks appear in the range of 500-700 cm⁻¹ with some exceptions at the 250 or 1050 cm⁻¹ (in SCWO), showcasing emerging defects or worsening symmetry. Those modes alongside originally registered peaks swing in place, losing their linearity before adjusting to new $\omega(p)$ trends. For CCWO, the process starts at 11 GPa, and for SCWO, as far as 15 GPa – those are also the limits till which the linear coefficients were calculated in *Table 17*. Those changes could be entirely dismissed as decomposition since they are completely reversible,³⁴⁷⁻³⁴⁹ however, past this point, a noticeable hysteresis could be noted - reaching ± 3÷3.5 GPa in comparison to BCW & BPW (0÷0.5 GPa). Below them, at 9 (CCWO) & ~12 GPa (SCWO), no delayed response was noted. It is not likely to fully and truly ascertain the nature of those transitions without access to other, more sophisticated HP techniques (XRD), but with the literature mentioned above and some analogy to neighboring structures that previously had their DFT calculations done. When followed thoroughly, peak vs. pressure dependencies in *Figure 30* feature a slow, intermittent zone of fluctuations that could be distinguished while samples transition to a new form suggesting a (Landau-Ginzburg) second-order phase transition as disputed by Howard et al.^{71,72} Those two derivatives surely lower their symmetry given common tilt systems, wildly spread defects, and partially substituted A²⁺/B(RE^{3+/4+}) sites – materials usually transition to a monoclinic C2/c (CCWO & SCWO) or C2/m (SCWO) SG – yet still, it is to be confirmed in the future given lack of HP XRD data and denoted slight hysteresis (typical for first-order).³⁵⁰⁻³⁵⁴

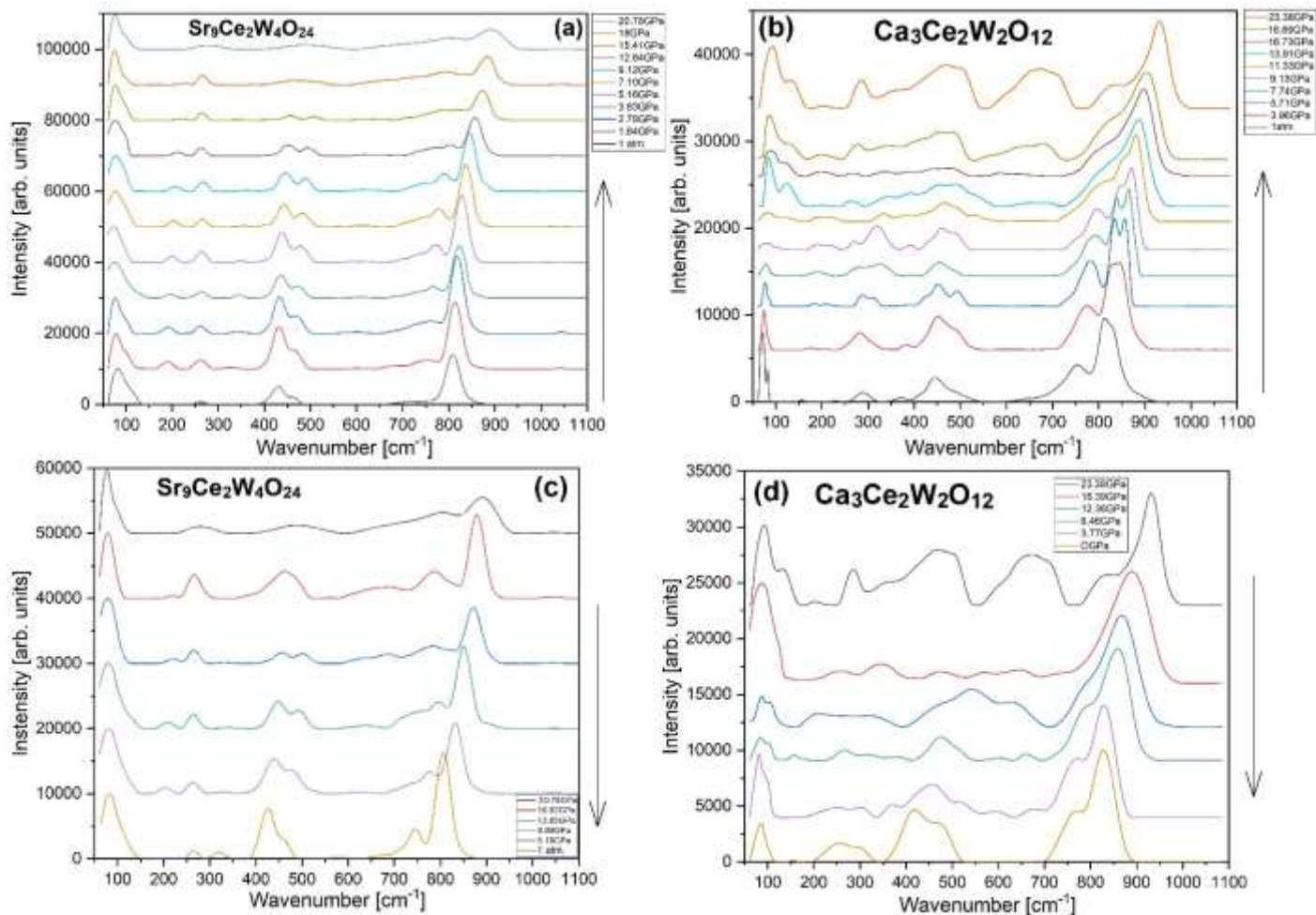


Figure 29 Raman spectra collected under hydrostatic argon PTM while using consecutive compression (top) – decompression (bottom) cycles for both SCWO (left) and CCWO (right) derivatives. Arrows guide the eye. Red vertical lines cut the pressure range into sections where theoretically reported second-order phase transition transpires.

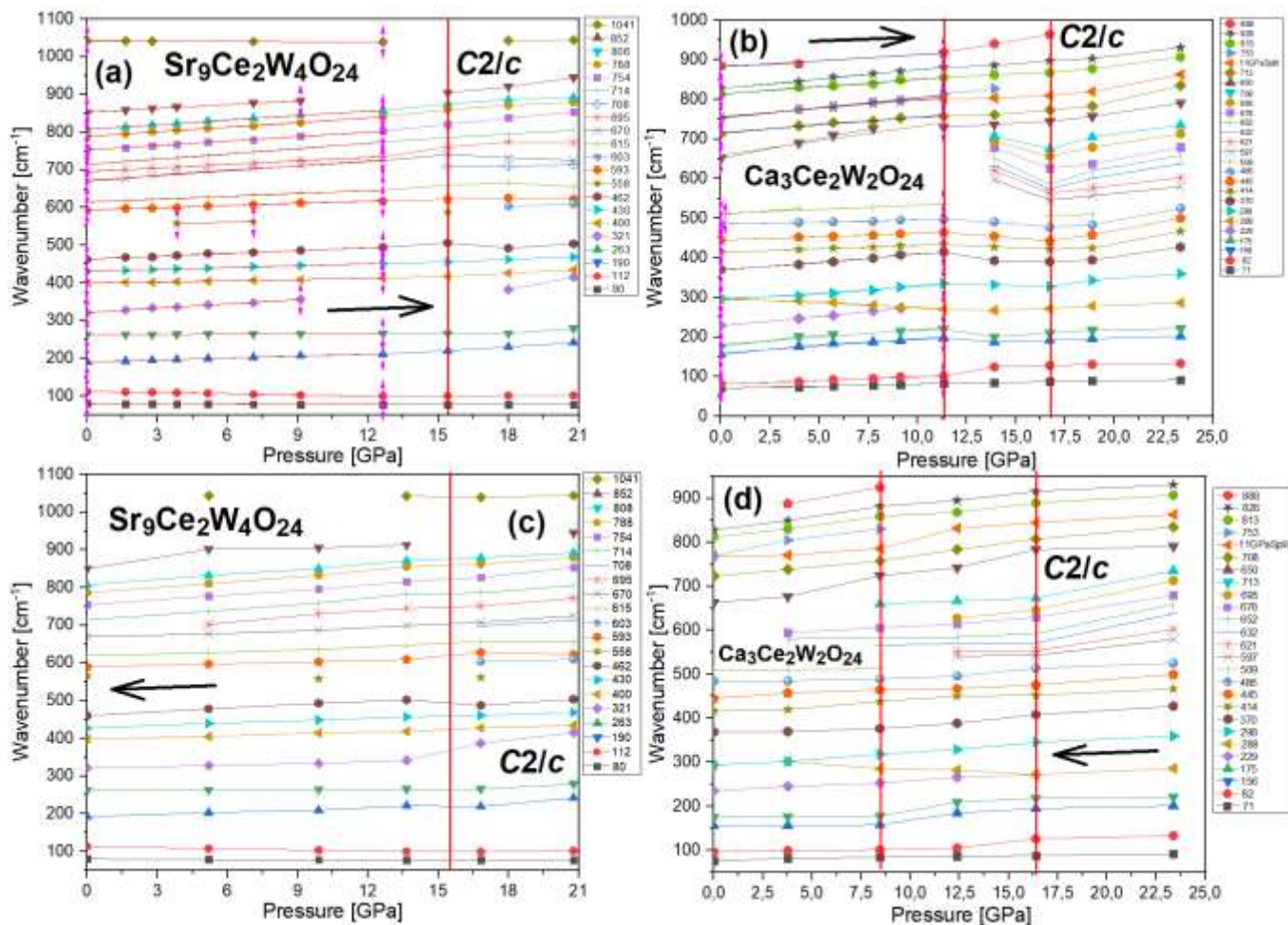


Figure 30 Raman peak position vs. pressure relations for modes featured in previous *Figure 29* regarding acquired derivatives. Presented similarly, compression is shown on the top, while decompression dependencies are on the bottom. SCWO to the left, CCWO to the right. Inserted right and left arrows guide the eye. Red vertical lines also cut the pressure range into sections where theoretically-reported second-order phase transition starts. Pink anchors mark the areas of linear fits. Error bars are too small to be visually noticed.

Table 17 Room temperature wavenumbers vs. pressure coefficients of Raman lines presented previously in *Figure 30* compression cycles for both SCWO and CCWO perovskite derivatives before any noticeable changes occurred.

Sr ₉ Ce ₂ W ₄ O ₂₄ (up to 15 GPa)			Ca ₃ Ce ₂ W ₂ O ₁₂ (up to 11 GPa)		
ω_0 [cm ⁻¹]	d ω /dP [cm ⁻¹ /GPa]	R ²	ω_0 [cm ⁻¹]	d ω /dP [cm ⁻¹ /GPa]	R ²
80	- 0.25 ± 0.01	0.97	71	1.09 ± 0.08	0.98
112	- 1.06 ± 0.04	0.98	82	1.85 ± 0.12	0.98
190	1.70 ± 0.05	0.99	156	3.61 ± 0.37	0.95
263	0.19 ± 0.03	0.82	175	3.78 ± 0.58	0.91
321	3.79 ± 0.11	0.99	229	4.81 ± 0.15	0.99
400	1.01 ± 0.05	0.98	288	- 2.29 ± 0.31	0.91
430	1.67 ± 0.04	0.99	298	3.35 ± 0.38	0.94
462	2.53 ± 0.09	0.99	370	4.02 ± 0.19	0.99
558*	1.47	1*	414	1.75 ± 0.06	0.99
593	1.95 ± 0.10	0.98	445	1.68 ± 0.05	0.99
615	2.51 ± 0.08	0.99	486	1.03 ± 0.06	0.98
670	4.31 ± 0.18	0.98	509	2.07 ± 0.41	0.89
695	3.12 ± 0.05	0.99	650	7.27 ± 1.37	0.90
714	4.70 ± 0.09	0.99	713	3.91 ± 0.33	0.97
754	3.88 ± 0.07	0.99	713 SPLIT	4.33 ± 0.36	0.97
788	4.07 ± 0.09	0.99	753	5.13 ± 0.22	0.99
808	4.00 ± 0.07	0.99	813	3.70 ± 0.23	0.98
852	3.37 ± 0.11	0.99	828	4.70 ± 0.13	0.99
1041*	- 0.18 ± 0.07	0.55*	888	3.09 ± 0.71	0.95

* - Peaks having doubtful factors (2-point or low R-squared)

At low temperatures, tungstate derivatives seem to be stable. Nothing out of the ordinary transpires in *Figure 31*. No new peaks or position swings – just slow-but-gradual, linear $\omega(T)$ progress, although in mixed directions, as seen in *Table 18*. The blue/red shifts are so minuscule that one can practically say that after confronting Raman setup, set with apparent error bars for the modes, they are particularly standing in place.

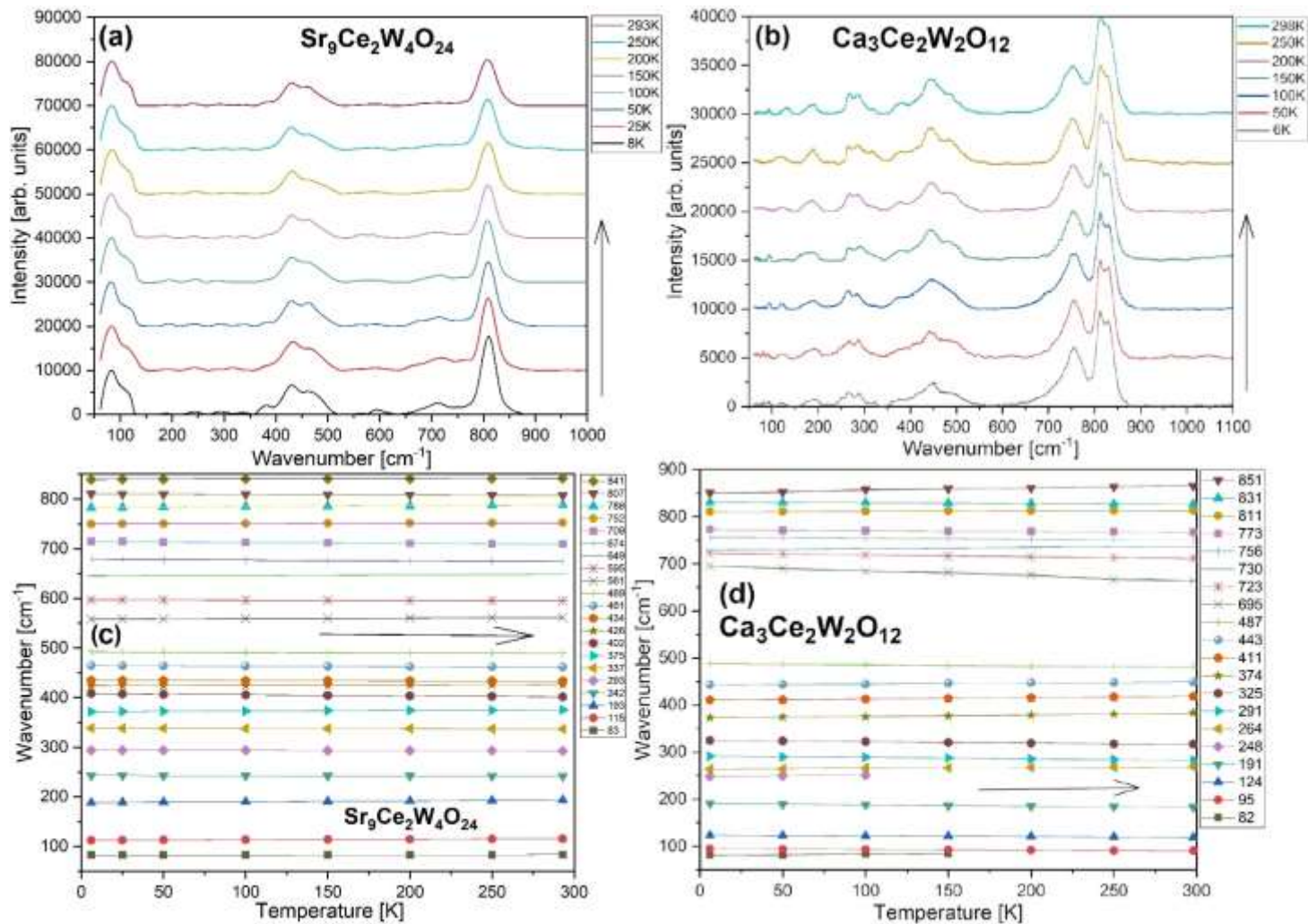


Figure 31 Low-temperature Raman spectra of perovskite derivatives presented in the function of heating from cryogenic to RT while their peak position vs. temperature dependencies are shown below. SCWO quaternary perovskite data is presented to the left, while CCWO ilmenite is to the right. Arrows guide the eye. Error bars are too small to be legible.

Table 18 Linear fit parameters obtained for temperature-dependent data presented in *Figure 31* for both SCWO & CCWO perovskite derivatives after cooling at ambient pressure.

Sr ₉ Ce ₂ W ₄ O ₂₄			Ca ₃ Ce ₂ W ₂ O ₁₂		
ω_0 [cm ⁻¹]	$d\omega/dT$ [cm ⁻¹ /K]	R ²	ω_0 [cm ⁻¹]	$d\omega/dT$ [cm ⁻¹ /K]	R ²
83	0.0031 ± 0.0002	0.98	82	0.020 ± 0.006	0.83
115	0.0113 ± 0.0005	0.98	95	-0.015 ± 0.001	0.98
193	0.0174 ± 0.0005	0.99	124	-0.014 ± 0.002	0.94
242	-0.0041 ± 0.0001	0.99	191	-0.025 ± 0.002	0.98
293	-0.0027 ± 0.0006	0.70	248	0.029 ± 0.003	0.99
337	-0.0047 ± 0.0001	0.99	291	-0.028 ± 0.002	0.98
375	0.0121 ± 0.0005	0.98	374	0.032 ± 0.003	0.95
402	-0.0218 ± 0.0007	0.99	411	0.029 ± 0.001	0.99
426	0.0064 ± 0.0001	0.99	443	0.022 ± 0.002	0.97
434	-0.0041 ± 0.0001	0.99	487	-0.021 ± 0.002	0.96
461	-0.0099 ± 0.0004	0.98			
489	-0.0100 ± 0.0008	0.95			
561	0.0120 ± 0.0003	0.99			

595	-0.0066 ± 0.0002	0.99			
649	0.0116 ± 0.0003	0.99			
674	-0.0156 ± 0.0004	0.99	695	-0.108 ± 0.008	0.97
709	-0.0189 ± 0.0002	0.99	723	-0.038 ± 0.002	0.98
			730	0.025 ± 0.001	0.98
752	0.0090 ± 0.0005	0.98	756	-0.019 ± 0.001	0.97
788	0.0172 ± 0.0005	0.99			
807	-0.0100 ± 0.0003	0.99	811	0.008 ± 0.001	0.92
			831	-0.013 ± 0.001	0.97
841	0.0067 ± 0.0002	0.99			
			851	0.051 ± 0.005	0.95

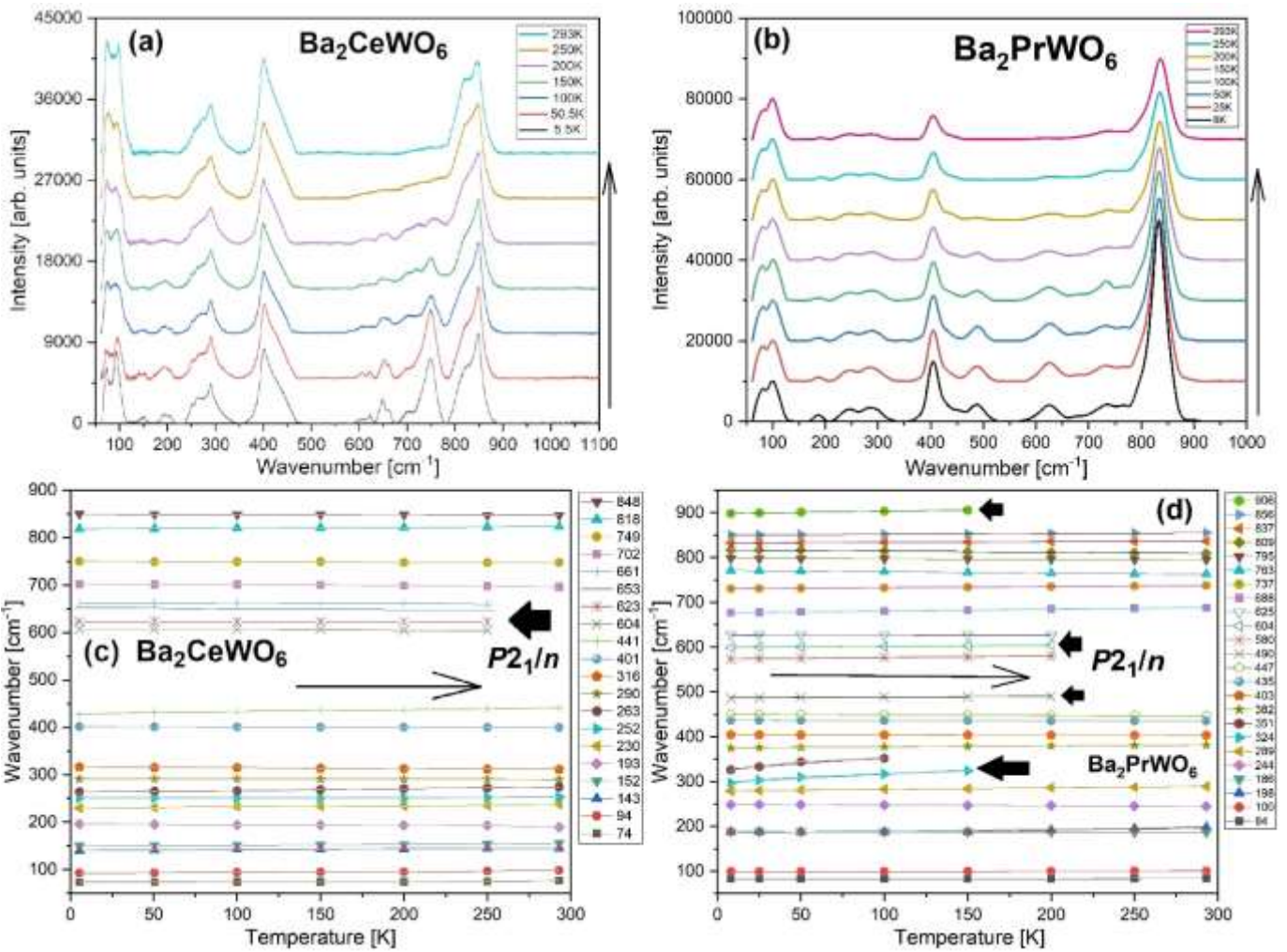


Figure 32 Low-temperature Raman spectra of barium double perovskites presented in the function of heating from cryogenic to RT, while their peak position vs temperature dependencies are shown at the bottom. BCW data is presented to the left, while BPW is to the right. Arrows guide the eye and pinpoint eventual changes related to the second-order phase transition. Error bars are too small to be visually noticed.

Again, the situation changes to the opposite in the case of barium DPs – they seem to undergo some sort of transition just below 200K. As shown in *Figure 32*, new ones appear near the set of typical peaks and sharpen down to He-level temperatures. Mostly between $550\text{-}650\text{ cm}^{-1}$ but in BPW, there are slightly more of them (just above 1000 and around 300 cm^{-1}) - all of them are marked as red in *Table 19* to distinguish them from the others. They appear probably because the material hosts largely unmatched RE/W B-site ‘pairs’, resulting in many oxygen defects and RE

vacancies distorting the overall structures (angle tilt) during temperature shrinkage.^{323–325,355} At that point, one could argue that this change could be decomposition, but that would not be reversible in heating. A more likely scenario is a stretched, continuous (second-order) phase transition, as suggested by Howard et al.^{71,72} or many others in the literature, given the prominent dissonance between *R*-3 & *I*2/*m* SG assessment.^{356–360} As it turns out, both structures could coexist with each other at RT (*R*-3 often transits to *I*2/*m* at ~50÷125°C, but some exceptions may apply). Usually, XRD would not be enough to assess the proper form of the powders at that point if one does not possess an ordered, more sophisticated single crystal or access to NPD.³⁶¹ The exact number of Raman active modes between each group also would not answer their elevated count after cooling. However, if those modes were just not accounted for (not visible) due to weak intensity or reforming RE-O bonds/strengthening defects (*Figure A11*), this would indicate a slow, gradual transition to a lower symmetry SG. Progressing angle tilt often causes a monoclinic SG, like *P*2₁/*n*, to emerge from any tilted cubic structure.^{356,360,362} This still, however, would require some confirmation using techniques like LT XRD, but any temperature dependence could help. HT XRD could tell us even more about angular dependencies concerning a wider range of temperatures if no transitions transpire – discussed later in *section 3.5*. Considering that only a few modes tend to change at LT – some just slightly faster than others, especially at the lowest 5K÷25K - most barely move and should be considered stationary - just like in SCWO & CCWO case. Their $\omega(T)$ coefficients are often so small that the given error bars could be roughly rounded up to zero, given the current resolution of the equipment. So, presented subtle blue/red shifts are just crude, statistical approximations that should be approached or rather judged carefully with some skepticism.

All aforementioned phase transitions and changes in this chapter were briefly summarized in graphs presented in supplementary *Figure A14*. One can view Howard's DFT-calculated DP transitions and counter them against experimental results.^{71,72}

Table 19 Linear fit parameters obtained for temperature-dependent data presented in *Figure 32* for both investigated barium double perovskites after cooling at ambient pressure. New LT peaks are marked as **red** and linearized in the whole range of existence.

Ba₂CeWO₆			Ba₂PrWO₆		
ω_0 [cm ⁻¹]	$d\omega/dT$ [cm ⁻¹ /K]	R ²	ω_0 [cm ⁻¹]	$d\omega/dT$ [cm ⁻¹ /K]	R ²
74	0.009 ± 0.002	0.69	84	0.0044 ± 0.0002	0.99
94	0.018 ± 0.002	0.96	100	0.0041 ± 0.0001	0.99
143	0.016 ± 0.002	0.94			
152	0.017 ± 0.002	0.91	186	-0.0058 ± 0.0003	0.98
193	-0.019 ± 0.004	0.77	198 Split	0.076 ± 0.003	0.99
203	0.023 ± 0.002	0.95			
252	0.006 ± 0.001	0.78	244	-0.0149 ± 0.0003	0.99
263	0.035 ± 0.002	0.98			
290	-0.0035 ± 0.0009	0.71	289	0.034 ± 0.002	0.98
316	-0.0179 ± 0.0007	0.99	324	0.183 ± 0.019	0.96
			351	0.273 ± 0.052	0.90
			382	0.023 ± 0.001	0.98
401	-0.0051 ± 0.0008	0.87	403	-0.0041 ± 0.0002	0.98

441	0.043 ± 0.003	0.96	435	-0.0054 ± 0.0001	0.99
			447	-0.0118 ± 0.0003	0.99
			490	0.020 ± 0.003	0.88
			580	0.035 ± 0.001	0.99
604	-0.014 ± 0.002	0.84	604	0.024 ± 0.002	0.96
623	-0.0018 ± 0.0006	0.62	625	-0.006 ± 0.001	0.78
653	-0.020 ± 0.003	0.86	688	0.040 ± 0.001	0.98
661	-0.019 ± 0.005	0.73	737	0.0241 ± 0.0003	0.99
702	-0.020 ± 0.002	0.92	763	-0.0305 ± 0.0004	0.99
749	-0.009 ± 0.002	0.83	795	-0.0107 ± 0.0004	0.99
818	0.016 ± 0.009	0.98	809	-0.0107 ± 0.0004	0.99
848	-0.0095 ± 0.0009	0.94	837	0.0260 ± 0.0006	0.99
			856	0.0180 ± 0.0003	0.99
			906	0.049 ± 0.003	0.98

3.2 XAS spectra at cryogenic temperatures

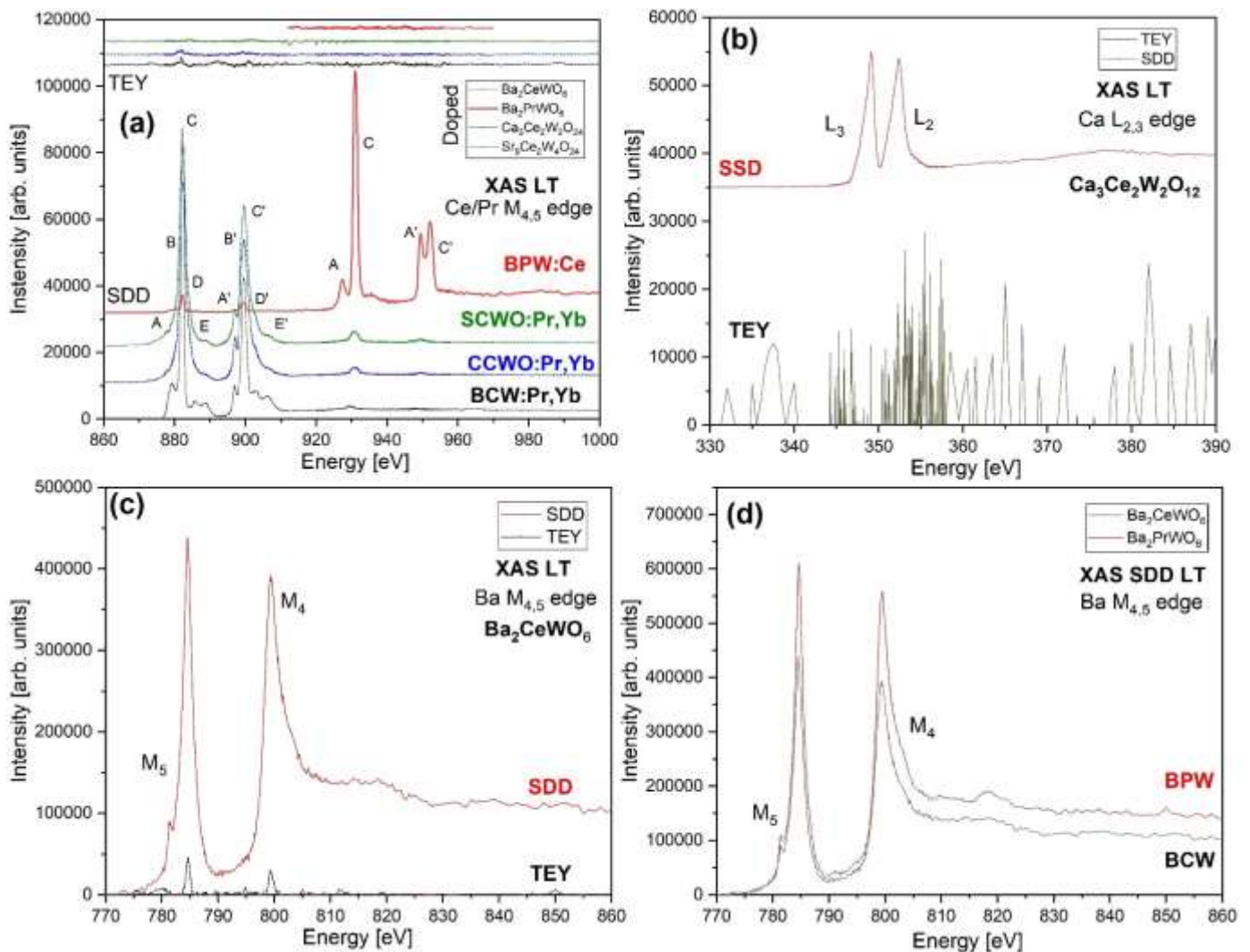


Figure 33 Low-Temperature XAS spectra yielded in dual acquisition mode showcasing prominent quality degradation of registered signals due to a significant drop in conductivity/rising insulating properties at various absorption edges: (a) Ce/Pr $M_{4,5}$ for all doped materials, as well as undoped (b) Ca $L_{2,3}$ edge CCWO, (c) Ba $M_{4,5}$ shells in BCW compared to (d) BPW using the only useful SSD mode.

Low-temperature X-ray absorption studies were done at some point, given access to other synchrotron facilities available on site. Knowing (from Raman spectra) that some of our materials (hosting barium) could have changed below 200K in cryogenic conditions, it is necessary to check using other techniques (like XAS) whether or not any of these claims were true. Here, the charge of given ions (mainly concerning RE/W) could also change during unit cell evolution (shrinkage/tilt) or SG transitions (to lower symmetry),^{363–365} especially concerning future PL energy transfer experiments (in *chapter 4*). So some of the most prominent results adhering to that prospect were collected and shown in *Figure 33*.

The spectra were not outstanding at first glance. SDD acquisition mode only gave some noticeable signals but solely in terms of constituents. Still, they were much weaker and quite noisy than their RT counterparts. RE³⁺ dopants were barely visible, and their complex features were not very distinguishable (especially considering low 1–5% concentrations of Pr). One could fit a lot (of possible peak sets) into those small, noisy spaces, and the error bars would be so high that a scientist could easily find any suitable match in the region of interest – compared to RT results. Befitting those conclusions - nothing majorly could be said about dopants, but two things stood out of the ordinary at LT. Firstly, the TEY acquisition mode barely registered anything at all. This method is a leading technique of signal registration, closely related to the sample's contact and conductivity. It is superior in all means to SDD if the conductivity of the sample is just right. Given all the adjustments: visual confirmation about the physical contact, appearance via camera, no problem with the status of the synchrotron beamline (feed), & the pins holding all the pellets to their molybdenum holders (in 4 separate cases) were good; the only conclusion that could accommodate for such a sudden drop of intensity would be a drastic decrease of conductivity towards insulating values (so that the detector could not receive any low-energy electrons showered off the pellets' surfaces (down the Mo drain)).^{366,367} That prompted several problematic issues concerning upcoming cryogenic PL & EPR experiments for all RE paramagnetic dopants and constituents that could be influenced (potentially frozen e⁻ carriers or magnetic transitions) during that time.^{368–370} What is interesting is that conductivity changes are related to both Ba samples' phase transition phenomena and the whole batch of derivatives (CCWO & SCWO too). The issue might stem from previously mentioned 200–150K (*section 3.1*) changes and some entrapping vacancies or W (mainly exotic & unstable 5+ ion changing quickly to 6+). SCWO & CCWO RE ions are seemingly unaffected, so tungsten might be the actual band gap manipulation culprit; however, it is uninvestigable since all its prominent shell edges lay outside the beamline's experimental range. *Section 2.3* XPS reports on slightly elevated stoichiometric O count, just like in *Table 10*, so that phenomenon might also be partially related to O_i/V_o presence (however severely limited). Expanding on that last issue, some spontaneous chemical reactions (resulting in stoichiometry changes or partial decomposition) might also take place at that time but only for both double perovskites (BPW more than BCW) due to much higher O_i content - these changes would be irreversible.

Barium samples seem to possess grossly elevated Ce³⁺ and Pr³⁺ content at LT, sacrificing much of their 4+ ions, but Sr and Ca derivatives do not. Since XAS edges do not look exactly the same as XPS, a need to compare through an analogy of M₅ A-E & M₄ A'-E' peaks vs. 3d shell d_{5/2} v_o-v'''' + d_{3/2} u_o-u'''' sets is necessary – original RT XAS was presented in *section's 2.3 Figure 19* while reference spectra are placed in appendix *Figure A10* if needed. Since the tungsten mentioned above could not be measured efficiently as an eventual donor of electrons at LT, to capture some perspective on apparent changes, the only way to counter that issue would be

indirect LT PL $W^{6+}O_4$ octahedra measurements - more about it later, in *section 4*. With the assistance of DFT band projections, one could confront models with actual reflectance spectra and discuss the W ionic charge state influence with actual matrix evolution. A lasting impression, for now, would be that some e^- donor, like $W^{4+/5+}$, could influence RE ions to change and upset the internal B-site balance by such mechanism: $2RE^{4+} + W^{4+} \rightarrow 2RE^{3+} + W^{6+}$ or $RE^{4+} + W^{5+} \rightarrow RE^{3+} + W^{6+}$. That charge transfer would surely influence any phenomena regarding energy transfer and overall stoichiometry of Ba_2REWO_6 to something resembling $Ba_{2+x}RE^{3+}\square W^{6+}O_{6+x}$ species (possibly similar to tilted $Ba_3Pr_2W_2O_{12}$ species hosting 9+ balance at LT. If O_i would be introduced into the matrix, that would further complicate the transition/recovery. A hint of possible application as a PL temperature sensor was made at that point; however, as a one-way-ticket if photobleaching (RE^{4+}) effect would take place.^{371,372}

3.3 Low-temperature EPR studies

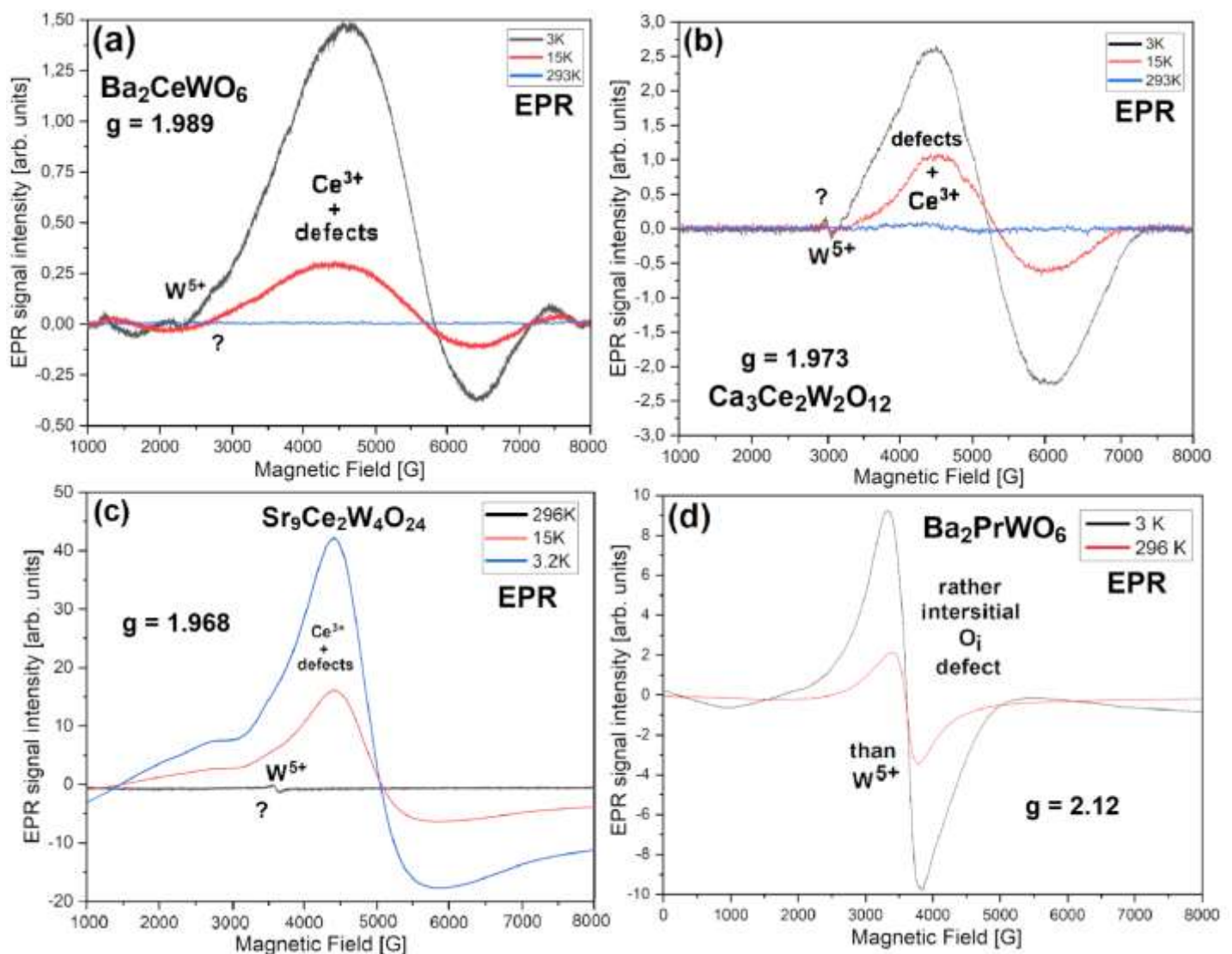


Figure 34 Cryogenically cooled EPR spectra collected for all investigated, undoped tungstate powders placed in perpendicular magnetic fields before any experiments or further modifications.

EPR spectroscopy was performed to assess the true form of ions in discovered, undoped materials knowing that most of RE reveal their paramagnetic nature abruptly at helium-level

temperatures while other ions like d-block W should be visible from the get-go. This technique would also allow the detection of other, specific magnetically active sites (O_i defects) or differentiate between subspecies sensitive to **perpendicular** & horizontal magnetic fields.^{373,374} Since our scientific group has only access to the setup with the **former** orientation of the magnetic field, a careful approach to see and compare the changes in W, Ce, or Pr magnetically active content before and after the eventual phase transition (below 150K) in barium double perovskites. To compare them with the other two tungstate derivatives concerning W^{5+} , Ce^{3+} & Pr^{4+} ions. Complementary verification of (dropping) conductivity through microwaves would also be an asset.^{375,376}

As can be observed in *Figure 34*, all spectra are crude and robust except the last one (d) regarding BPW. They all have an unfortunate effect associated with powders that possess defective, anisotropic structures hosting chaotically scattered paramagnetic domains highly diluted across the matrices. That makes subtle multisites not distinguishable to the naked eye.^{328,377} Moreover, the signal from Ce^{3+} in the first three graphs (a - BCW, b - CCWO, and c - SCWO) envelopes features of something else on its back (lower Gauss range). Given acquired g-factors, which are slightly below 2 (1.989, 1.973, 1.968, respectively), the electro-donor species could be associated only qualitatively via their positions (maximum value of curves' derivative) versus what the literature truly says without procuring any prolific details. A quite similar type of signal was achieved for the precursor, $BaWO_4:Ce$.^{328,378-380} As Ce^{3+} species seem to fit quite well into the Gaussian range (~4.8 kG bracket) showcasing "below 20K" activation mechanisms with $x < 2$ g-factors, these lumps present at the slopes in the lower magnetic fields seem definitely different in origin. - there might be something else there. Possible defects or exotic W^{5+} . The problem with the last one is its low content & highly unstable nature. Furthermore, as a d-block ion, it should be visible already at RT if not influenced by any external charge-transfer sources at LT. It does not possess such an activation mechanism as RE ions & should be visible from the start with the g-factor lowered to 1.6-1.7.^{381,382} The range of this eaten-up signal (~3.4 kG), however, fits that dogma. It is just too weak and misshaped to be sure of its origin. Yes, it is contained as a minor phase, but should not be so weak and disfigured. The answer possibly lies within the BPW spectrum. There, the signal is much clearer, well-defined at RT, and with the given g-factor (above 2 ~ 2.12), it would debunk its presence as interstitial oxygen (O_i).³⁸³⁻³⁸⁵ That would not fit the scenario of g-factors at previous spectra. This, however, might be just an outcome of both species coinciding with each other in one matrix resulting in these final g-factor's values.^{386,387} What is more troubling is the absence of Pr^{4+} , which should be there at LT (visible just like Ce^{3+} given its e^- configuration), around 1.9 kG (g_z) joined by O_i at 3.5 ($g_{x,y}$) as seen in REO_{2-x} references³⁸⁸⁻³⁹¹ in the supplementary *Figure A15*.^{388,392,393} According to XPS & XAS, this ion is abundant in the matrix at RT. The spectrum (d) should look at least similar to the other 3 (a,b,c). Yet still, nothing. That could vouch for either change of Pr^{4+} to Pr^{3+} at LT via photoionization,³⁹⁴⁻³⁹⁶ or antiferromagnetic phase transition encountered previously at similar Ba_2PrRuO_6 family of compounds featuring $T_{Néel}$ in between 90-117K.³⁹⁷ That would render Pr^{4+} & W^{5+} ions completely undetectable at 20-3K.

Later (*chapter 4*), this EPR investigation, alongside XAS & PL, will lay firm fundamentals to support time-resolved, charge-transfer phenomena (between $RE \leftrightarrow W$) research shedding some light on gradual O_i evolution in barium DPs during RT highly-energetic illumination. The nature of possible photoionization at LT would be approved in between 90 & 150K where possible $T_{Néel}$ & Raman phase transition lie.³⁹⁸⁻⁴⁰¹

3.4 DSC & TG stability research

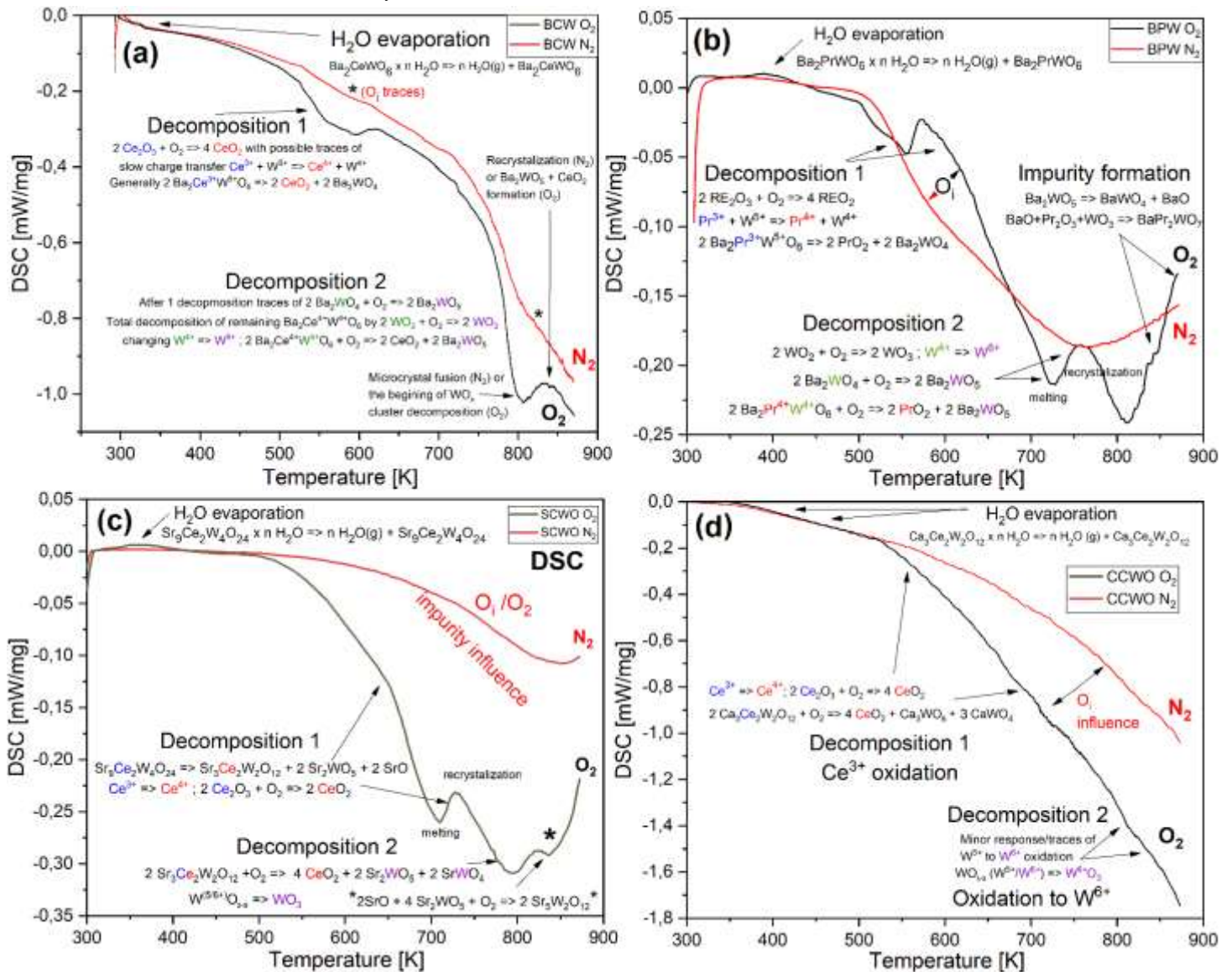


Figure 35 Differential Calorimetry Scans of all investigated, undoped materials in referral, inert N₂ (red) and air (black) atmospheres showcasing the regions of stability and decomposition done in the perspective of future usage at ambient conditions: for both (a) Ba₂CeWO₆, (b) Ba₂PrWO₆ double perovskites, together with (c) Sr₉Ce₂W₄O₂₄, (d) Ca₃Ce₂W₂O₁₂ tungstate derivatives.

Differential Scanning Calorimetry (DSC) and Thermogravimetry (TG) were included as pre-emptive reconnaissance methods to investigate the first signs of materials' instability since perovskites, as a whole mineral group, have a lot of unstable variants – at least when exposed to some specific stimuli in the long run, like heat or strong, monochromatic light in the air. Confronting this data with the literature,^{402–404} some estimates could be made regarding chemical reactions and decay (oxidation) to specific subspecies supported shortly thereafter by HT XRD and Raman spectroscopy studies. That way, some genuine phase transitions could be noticed in protected inert conditions unrelated to decomposition. If not, the patterns and spectra will still reference those changes seen in the air.^{405,406} Also, in the areas of more noticeable changes (air), the techniques mentioned above could efficiently probe the structures in a more detailed, small-step manner without losing any time and resources.

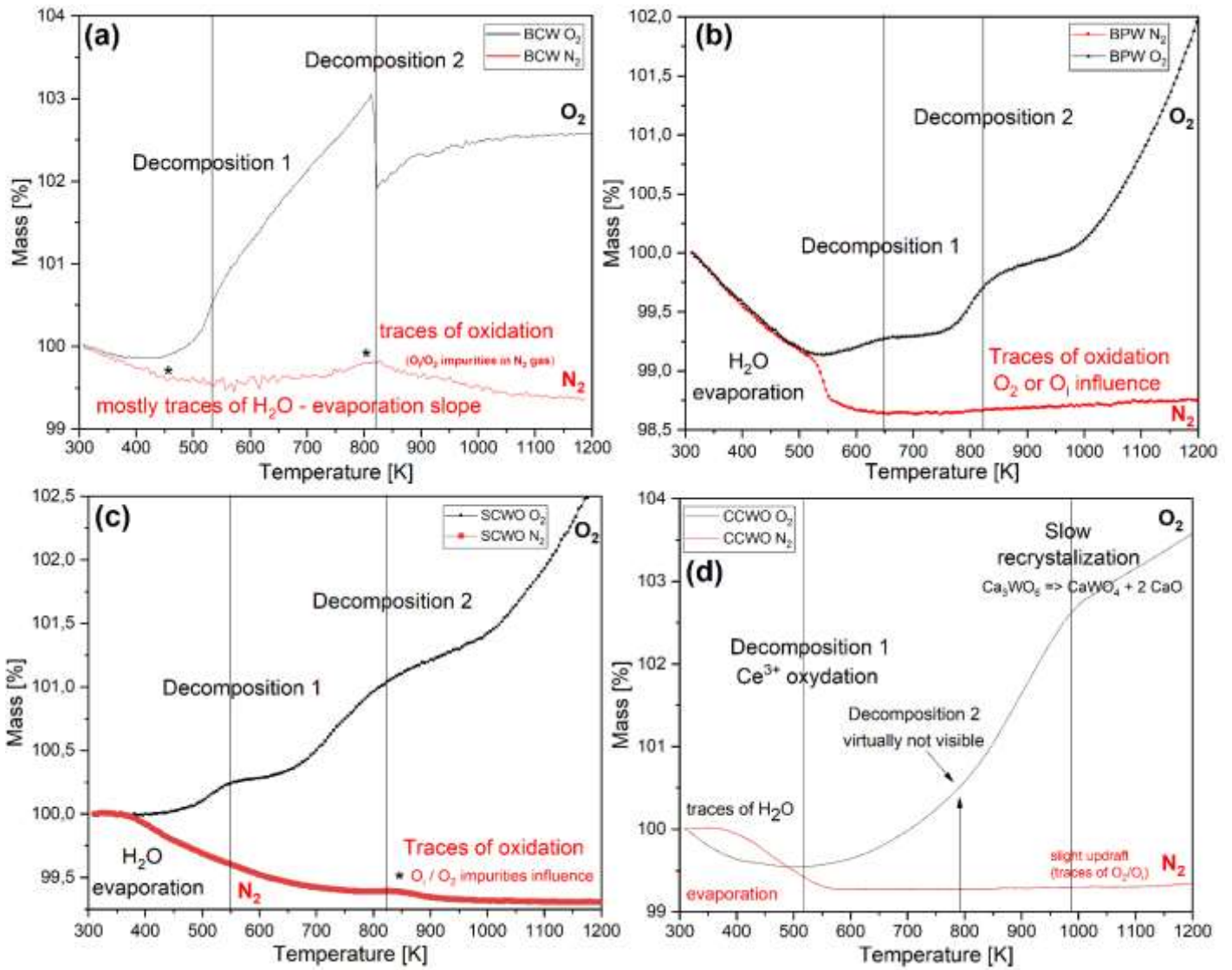


Figure 36 Thermogravimetry diagrams of all investigated, undoped materials in referral, inert N_2 (red) and air (black) atmospheres showcasing the regions of stability and decomposition done in the perspective of future usage at ambient conditions: for both (a) Ba_2CeWO_6 , (b) Ba_2PrWO_6 double perovskites, together with (c) $Sr_9Ce_2W_4O_{24}$, (d) $Ca_3Ce_2W_2O_{12}$ tungstate derivatives.

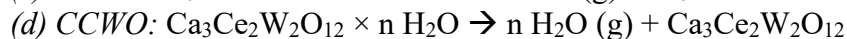
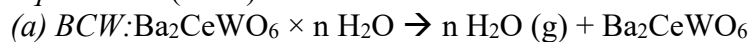
Before divulging specifics about apparent changes deduced from foreseeable features in all graphs presented in that chapter, air-related (**black**) curves show a more complex nature than N_2 -related (**red**) ones. Although much could be told in comparison between each one of them, one could impute that they should be completely flat (if no decay, decomposition, or phase transition occurs); however, despite countless efforts, some influence from ingrained impurities, like O_i , could not be avoided for samples submerged in N_2 .^{407,408} (Intercalated water could be by preheating all materials in the first, inert cycle but this effect is counted more as a genuine, storing feature than a problem). This bends all lines towards each other, creating an exo-energetic effect but with much weaker, less prevalent consequences. Let it be a baseline since O_i could not be expelled from the materials efficiently (they are oxide, after all) without affecting their matrix as a whole (by incoherent reduction, $-OH$ creation, or expansion of intercalated H_2O content). It will still behave like air to some extent but with a limited (trace) supply and higher activation energy.^{387,409,410}

Following DSC measurements in *Figure 35* with complementary TG measurements in *Figure 36*, one can distinguish some common features not only between all materials (double perovskites lines progress similarly just like Sr and Ca derivatives between each other) but also air curves show prominent lumps and dents related to the same type of changes in closely-related temperature frames. N₂ is less rich & prevalent, but it is to be expected. It is no surprise since all of them contain mostly the same ions, just arranged differently – they should be as sensitive and react the same way as presented in the insets. However, the actual question is whether or not they (materials & their curves) relate to the same processes for sure – that is why they should be collectively analyzed not only with literature but alongside some crystallographic (HT XRD) or spectroscopic (HT Raman) techniques to be completely sure. Those measurements were done later on the basis of this data to focus on the main features like “Decomposition 1 or 2” meaning that in the next chapter, they will confirm given reactions with higher precision by a tighter step-by-step method. General processes commence as follows:

First, the evaporation of the water stored in not pre-heated samples is a typical process. Even in a desiccator, the moisturization could not be entirely stopped – it just slowed down to some decent degree. In TG, evaporation is shown as an initial drop of weight while approaching the boiling point of water at ambient pressure. In DSC, the curve should either show an endothermic (+ value) uprise or balance itself out discretely due to some opposite competing processes (as observed here).^{411,412} There, the curves are temporarily flat or show less steep descent until oxidation (more pronounced by O₂ from the **air**) or more subtle by the influence of O_i (confined within the material in limited amount while N₂ is surrounding the material) overtakes the balance as a dominating exothermic (-) process above ~400/500K.

The aforementioned processes look more-or-less like the reactions presented below (or in the insets of *Figure 35*) and will not show much on XRD or Raman spectra if only some traces of water are left. Few patterns might just slightly improve in intensity by lowering the general amorphization caused by water.

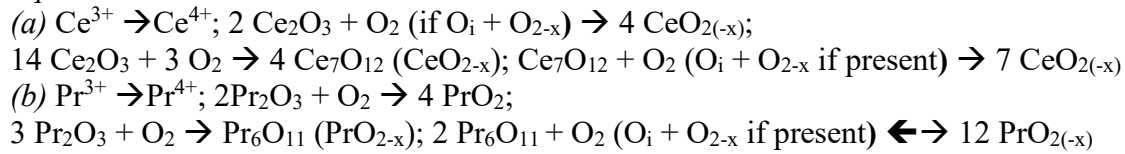
Equation 1: (n ~ 6)



The ‘n’ number is often associated with the ligand number depending on the coordination and strength of the crystallographic field surrounding the main structure. Steric hindrance and the main ions' size or rather charge state dictate it to be 2, 4, 6, or 8. Regarding water dipoles, it is often six as each hydrogen bonding has the angular freedom to bend towards oxygen at each neighboring xyz (+/-) axis. Therefore, water encapsulates such structures in a neat bubble.^{413–415}

The first significant decomposition process is related to RE³⁺ → RE⁴⁺ oxidation. Mainly Ce or Pr, which starts slightly faster for Ce but is more abrupt for Pr. The first definitive signs in decay processes could be distinguished above 475K, but at 550K, the culmination point peaks.^{416–420} This is also the terminus (endpoint) of stability and usage in the air, as later, one will end up with a totally different blend of products at their disposal. The overall reactions dictated by the literature progress at this stage as follows:

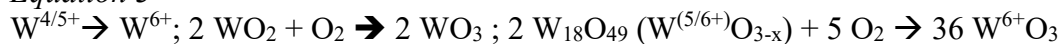
Equation 2:



However, These sets are not customized to actual products, as seen in the insets. More about them in HT XRD/Raman 3.5 chapter below. For now, let us elaborate on the actual mixed valence state of REO_{2-x} oxides present in the equations above. They are all justified due to the availability of various B-site ions previously reported by XAS/XPS. Pure PrO₂ is neither stable nor achievable in its purest form (only temporary) – it reverts to & is referred to as PrO_{2-x} form (Pr₇O₁₂, Pr₅O₉, Pr₆O₁₁, Pr₁₁O₂₀).¹⁹⁶ This last fact might cause a lot of disorder, abrupt breakage, and insert more O_i into the matrix creating V_o, like in BPW.

More pronounced DSC curve features reported (by XPS) (Ce,Pr)^{3/4+/W^{5/4+}} disproportions by smaller peak area for the former RE ‘pair’, which brings us to the following (later) ‘Decomposition 2’ process associated with intense W⁴⁺ oxidation to the W⁶⁺ state. Any remaining residual W⁵⁺ does the same but slowly and subtly.⁴²¹⁻⁴²⁴ The process actually starts above 725-750K but reaches its culmination point at +800K:

Equation 3



The reaction does not stop there. More subtle, slowly progressing exothermic processes could be traced as curves progress vertically up or down (melting or recrystallization since only horizontal flat lines symbolize discrete & brief temperature stop/reaction balance).⁴⁰²⁻⁴⁰⁴ Processes gradually penetrate the material deeper and faster with each minute in the air (in N₂ negligible influence of O_i) while main products dismantle further to salts (tungstates) or rather, more importantly, (alkaline) oxides and carbides (that last group concerns W with graphite/residual carbonates). The last group is only temporarily stable, solely above 1000K (as showcased more prominently by TG & post-processing XRD since DSC could reach that high due to apparatus limit).^{421,425,426} Later, the former alkaline oxides will diminish (on the surface) and react with each other stuff, like moisture or CO₂ from the atmosphere. Traces of this phenomenon will be visible after milling the pellets and seeing their insides –all minor XRD phases are shown in section 2.2. It also seems that AO oxides try to react with REO_{2-x} giving something in the shape of BaREO_{3-x}.

In the range of ambient environmental stability (below any signs of decomposition stage 1), an average heat capacity (C_p) can be calculated against proper references (like sapphire) from data gathered in Figure 37. This extensive physical propriety is crucial in designing materials used in civil engineering, like sensors and solar panels. Working in different temperatures, energy is needed: to control their energy usage, storage, and expansion; to avoid damage and eventual resource loss, so their estimation would be definitely useful.^{427,428} As shown in Table 20, their value strictly depends on the mass and expands with growing stoichiometry. By comparing them to other double perovskites, at least barium compounds fit gently into declared ranges.^{401,429,430} Derivatives do not have much reference to compare with other than close-call garnets vs. ilmenites – matches are good too.⁴³¹

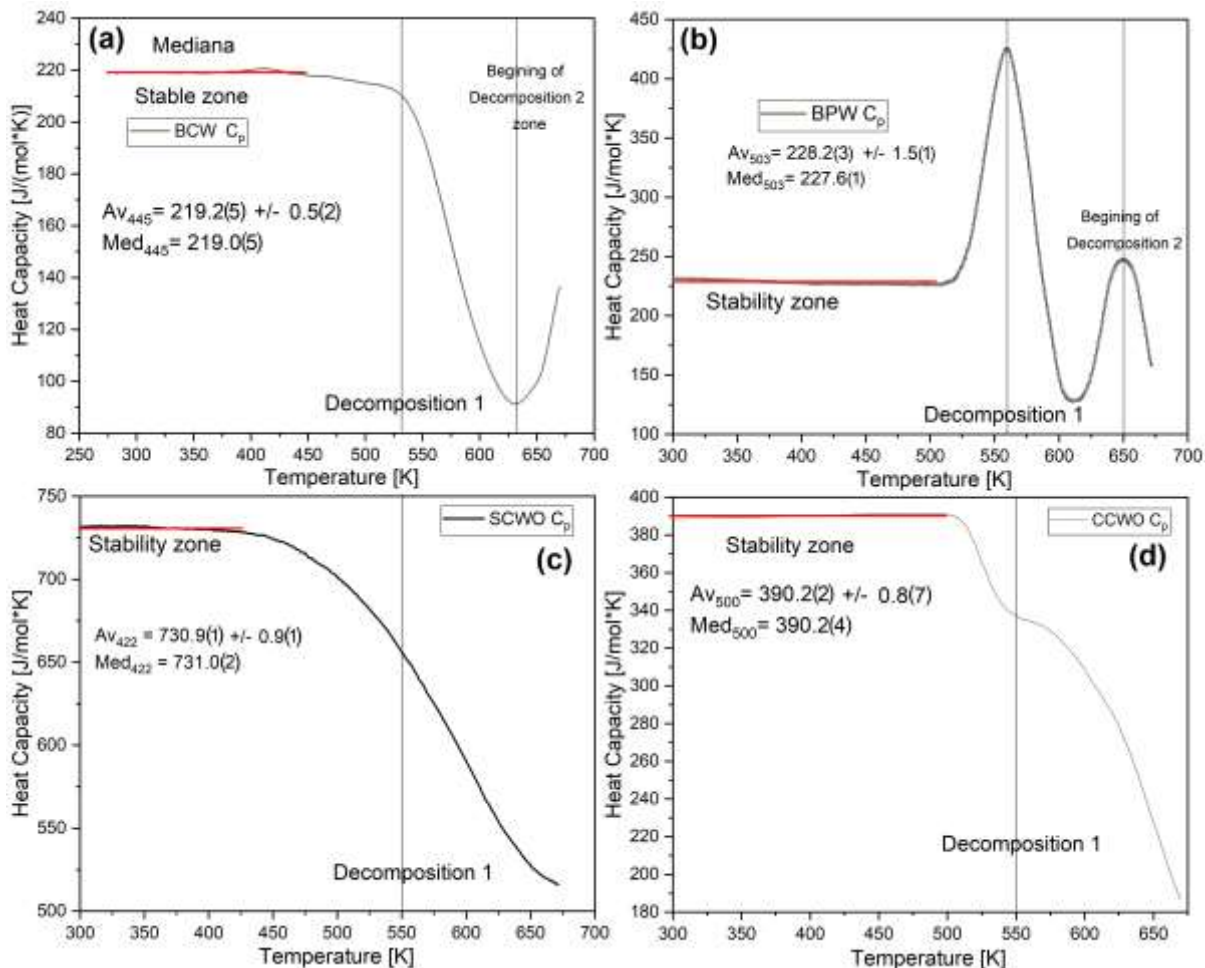


Figure 37 Molar heat capacity measured in the air for all investigated compounds in the function of heating against a sapphire reference. The mean values given in the text refer to stable zones marked by red lines.

Table 20 Standardized experimental molar C_p values calculated for all 4 investigated compounds against sapphire reference in the air during previous DSC measurements in *Figure 37*.

Material	Average [$\text{J} \times \text{mol}^{-1} \times \text{K}^{-1}$]	Average deviation \pm [$\text{J} \times \text{mol}^{-1} \times \text{K}^{-1}$]	Median [$\text{J} \times \text{mol}^{-1} \times \text{K}^{-1}$]	Maximum stable temperature of calculations [K]
Ba_2CeWO_6	219.2(5)	0.5(2)	219.0(5)	445
Ba_2PrWO_6	228.2(3)	1.5(1)	227.6(1)	503
$\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$	730.9(1)	0.9(1)	731.0(2)	422
$\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$	390.2(2)	0.8(7)	390.2(4)	500

3.5 High-temperature decomposition via XRD & Raman

High temperatures (investigated via XRD & Raman spectra) are equally important as LT from a technical standpoint, knowing that solar panels could work both in space and on the roof of the house at high noon. Especially since HT can be explored in the broader temperature range (mainly using XRD, which was not possible at LT), some axial and angular dependencies could be studied having a better perspective in mind. As foreshadowed in the previous chapter by DSC and TG, all our substances seemed unstable (DPs apparently more than the other two derivatives). The remaining task was to check to what exact materials they are decomposing to and are those changes also occurring in the protective N_2 atmosphere (are they negligible enough to withstand the influence of residually stored O_i)? If yes, can we use them in an oxygen-deprived environment, given that no HT phase transition transpires?^{65,142,146,387} One shall see.

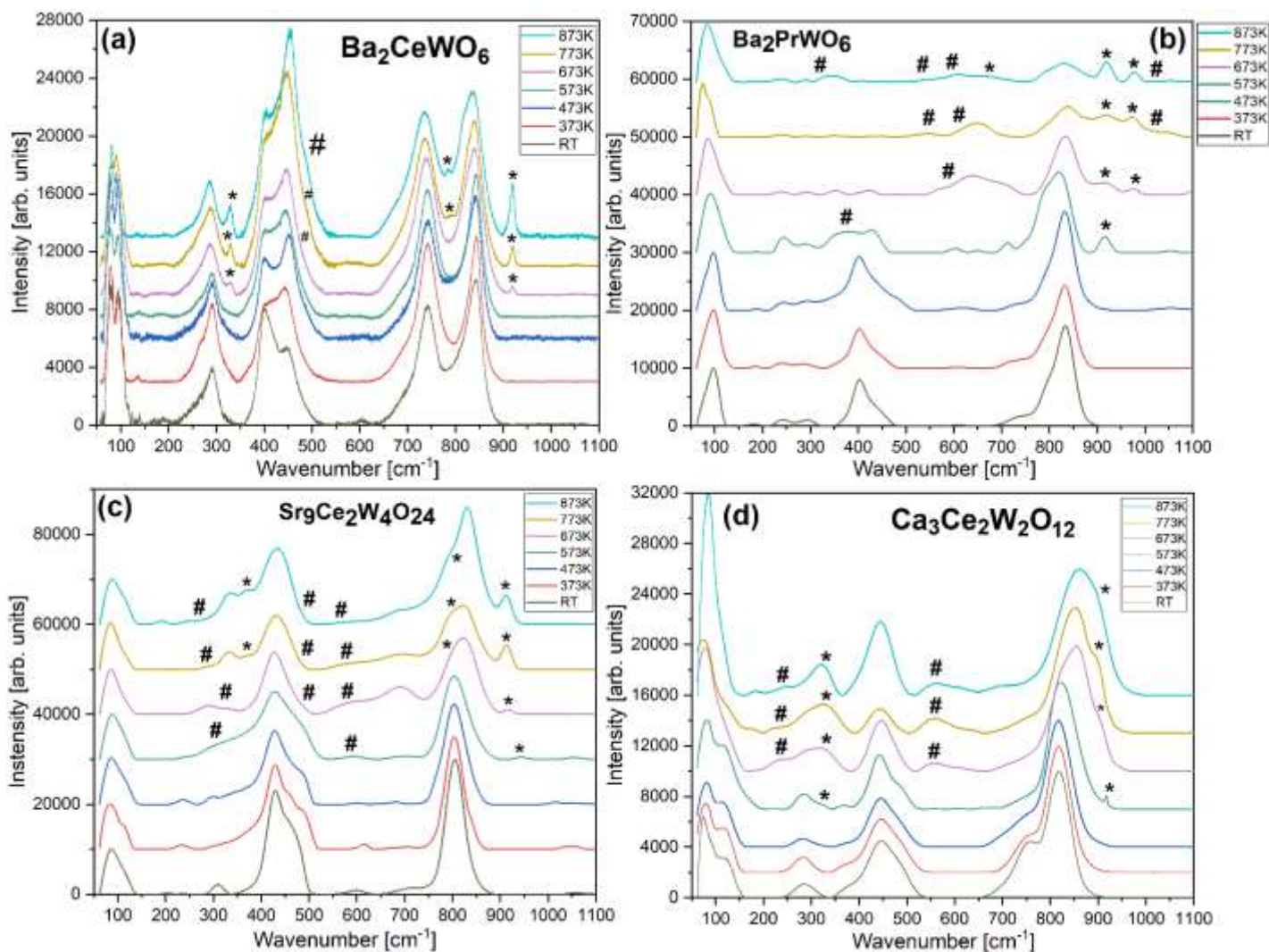


Figure 38 Raman spectra of all studied materials collected in the function of heating from room temperature up to 873 K in the air (without any protective, inert N₂ atmosphere) show noticeable features of decomposition mostly above 573 K. Asterisks mark modes originating from forming tungstates (A_xWO₄). Hashtags signify a group of various defective REO_{2-x} oxides detaching from the original matrix.

Raman HT spectra depicted in *Figure 38*, without N₂ protective atmosphere (air), show many features emerging at temperatures above 573K – they correspond to 1st stage of decomposition related to Ce³⁺ (*Figures 35 & 36*). Browsing the literature,^{319,320,322–326,417,431,432} they could be easily associated with mixed valency REO_{2-x} oxides (#hashtags) and later, A_xWO_{3+x} tungstate species (*asterisks) marking gradual changes towards 2nd stage of decay.^{2,15,22,23,26,28,203,433} Gradual, visual degradation of the samples could also be traced by the naked eye via prominent discoloration from originally reported colors (*Figure 13*) – from green or brown to i.e, yellowish or greyish powders as shown in *Figure A16*. In N₂, this change was not present – at least not macroscopically, as depicted in *Figure A17*. Since Raman detected something resembling the REO_{2-x} phase sweated off BPW and SCWO surface, that might happen in some highly-limited amount (O_i influence).

This is, however, not enough to pass the final verdict about the usage & actual phases. If there is no natural first or second-order transition in N₂ but pure decomposition, HT powder XRD diffraction patterns in *Figure 39* would be crucial to assess that situation. They were done to confirm that in the air, materials decay to more (barium DPs *Figure 39a,b*) or less (Ca and Sr derivatives *Figure 39c,d*) prevalent phases occur, taking into account all weakening & shifting peaks with additional, new features appearing spontaneously here and there (arrows guide the eye). Amorphization is also a big issue concerning all flattening patterns in the CCWO.

Peaks' FWHMs increase gradually with the temperature to the point where nothing is visible - chemical changes might still occur but are not distinguishable from the noise & glassy structure.^{434,435} Another derivative, SCWO quaternary perovskite, seems to undergo only some subtle changes – the least of them all. One could also see some broadening peaks swinging or disappearing here and there (vouching for slow, gradual decay to tungstates & oxides) put in big close-ups in appendix *Figure A18*. It seems that expanded O & W⁶⁺ content (featured in XPS *section 2.3*) benefits the stability of these structures in general. Measurements acquired later, at RT, from decomposed matrices at specific temperatures assess the status of samples in much greater detail regarding possible chemical mechanisms of decay - see *Figure 40* for reference. Together with already gathered data about minor phases in *section 2.2* & DSC with TG measurements presented in *section 3.4*, one can deduce (to some extent) all the plausible reactions that occur at HT in the investigated materials. Beneath is the list of all currently-recorded materials & impurities:

- Ba₂CeWO₆: BaO (~1000K),⁴³⁶ CeO_{2(-x)} i.e. Ce₇O₁₂,⁴³⁷ BaWO₄, Ba₂WO₅, Ba₂WO₄,^{2,313}
- Ba₂PrWO₆: Pr₂O₃ (PrO_{2-x}),¹⁶⁷ BaPr₂WO₇,¹⁹⁰ BaWO₄, Ba₂WO₅, Ba₂WO₄;
- Sr₉Ce₂W₄O₂₄: Sr₃Ce₂W₂O₁₂,⁴³⁸ CeO_{2(-x)}, SrWO₄,⁴³⁹ Sr₂WO₅,⁴⁴⁰ Sr₅W₂O₁₂;^{441,442}
- Ca₃Ce₂W₂O₁₂: CaO (~1000K),^{442,443} CeO_{2(-x)}, Ca₃WO₆,⁴⁴⁴ CaWO₄,^{445,446}

Collectively, the XRD database suggests that such decomposition products occur in various ratios using an unprotected atmosphere at HT. That leads us to establish some crude stoichiometric reactions of decay (from *Figure 35*) leading at specific intervals of temperature:

At decomposition series 1 (Cerium ~550K):^{416,417,419,420,425}

- BCW: 2 Ba₂Ce³⁺W⁵⁺O₆ → 2 CeO₂ + 2 Ba₂WO₄
- BPW: 2 Ba₂Pr³⁺W⁵⁺O₆ → 2 PrO₂ + 2 Ba₂WO₄
- SCWO: Sr₉Ce₂W₄O₂₄ → Sr₃Ce₂W₂O₁₂ + 2 Sr₂WO₅ + 2 SrO
- CCWO: 2 Ca₃Ce₂W₂O₁₂ + O₂ → 4 CeO₂ + Ca₃WO₆ + 3 CaWO₄

At decomposition series 2 (Tungsten ~800K):⁴²¹⁻⁴²⁴

- BCW: 2 Ba₂Ce⁴⁺W⁴⁺O₆ + O₂ → 2 CeO₂ + 2 Ba₂WO₅; 2 Ba₂WO₄ + O₂ → 2 Ba₂WO₅
- BPW: 2 Ba₂Pr⁴⁺W⁴⁺O₆ + O₂ → 2 PrO₂ + 2 Ba₂WO₅; 2 Ba₂WO₄ + O₂ → 2 Ba₂WO₅
- SCWO: 2 Sr₃Ce₂W₂O₁₂ + O₂ → 4 CeO₂ + 2 Sr₂WO₅ + 2 SrWO₄

Brief reactions that take place later (continue after +900K - 2nd stage of decay) or their intermediate products are just contemporary (mainly referring to alkaline impurities disappearing after cooling - probably reacting with REO_{2-x}, CO₂, or moisture if exposed to air)*:^{442,447-449}

- BCW: Ba₂WO₅ → BaWO₄ + BaO*; BaO + CeO_{2-x} → BaCeO_{3-x}
- BPW: Ba₂WO₅ → BaWO₄ + BaO*; BaO + Pr₂O₃ + WO₃ → BaPr₂WO₇
- SCWO: 2 SrO* + 4 Sr₂WO₅ + O₂ → 2 Sr₅W₂O₁₂
- CCWO: Ca₃WO₆ → CaWO₄ + 2 CaO*; CaO + CeO_{2-x} → CaCeO_{3-x}

Now, seeing some evidence about regional stability and all of the processes mentioned above taking place in the air, one can calculate axial and structural parameters (up to roughly ~550K), which would be necessary for future applications and modeling. Ba₂CeWO₆, Ba₂PrWO₆, Sr₉Ce₂W₄O₂₄ & Ca₃Ce₂W₂O₁₂ data are presented consecutively in *Figures 41, 42, 43, and 44* however, all their linear fits are gathered in *Table 21*.

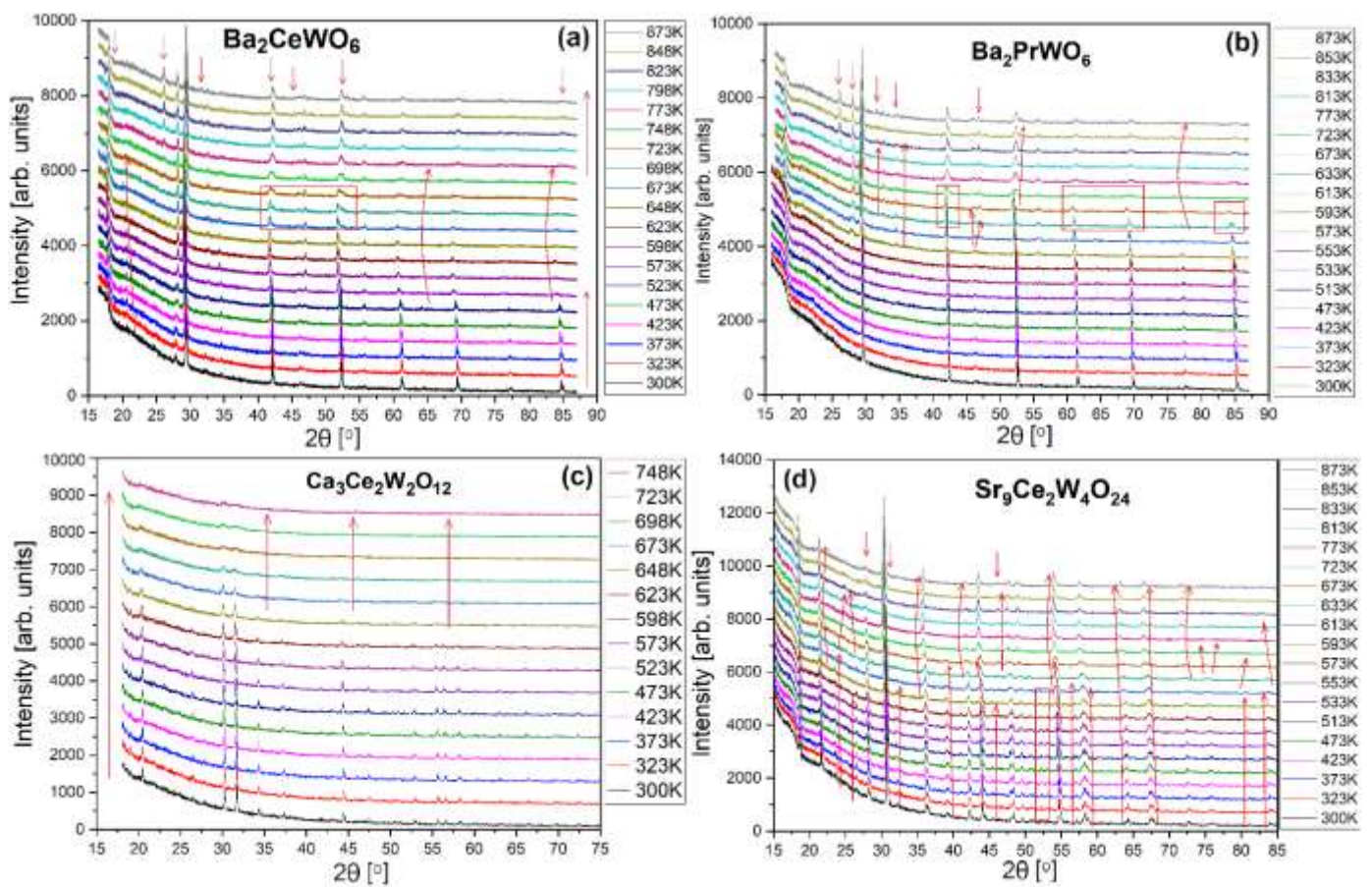


Figure 39 High-temperature powder XRD diffractograms presented the function of heating for all 4 investigated RE-tungstates. Each pair (double perovskites and derivatives) share a similar degree of decay after submerging in hot air - to emphasize those rates, red arrows were implemented to guide the eye regarding a few key features. In (a) BCW and (b) BPW, they are more abrupt, but in (c) CCWO & (d) SCWO more subtle & required assistance in the form of close-ups available in supplementary *Figure A18*.

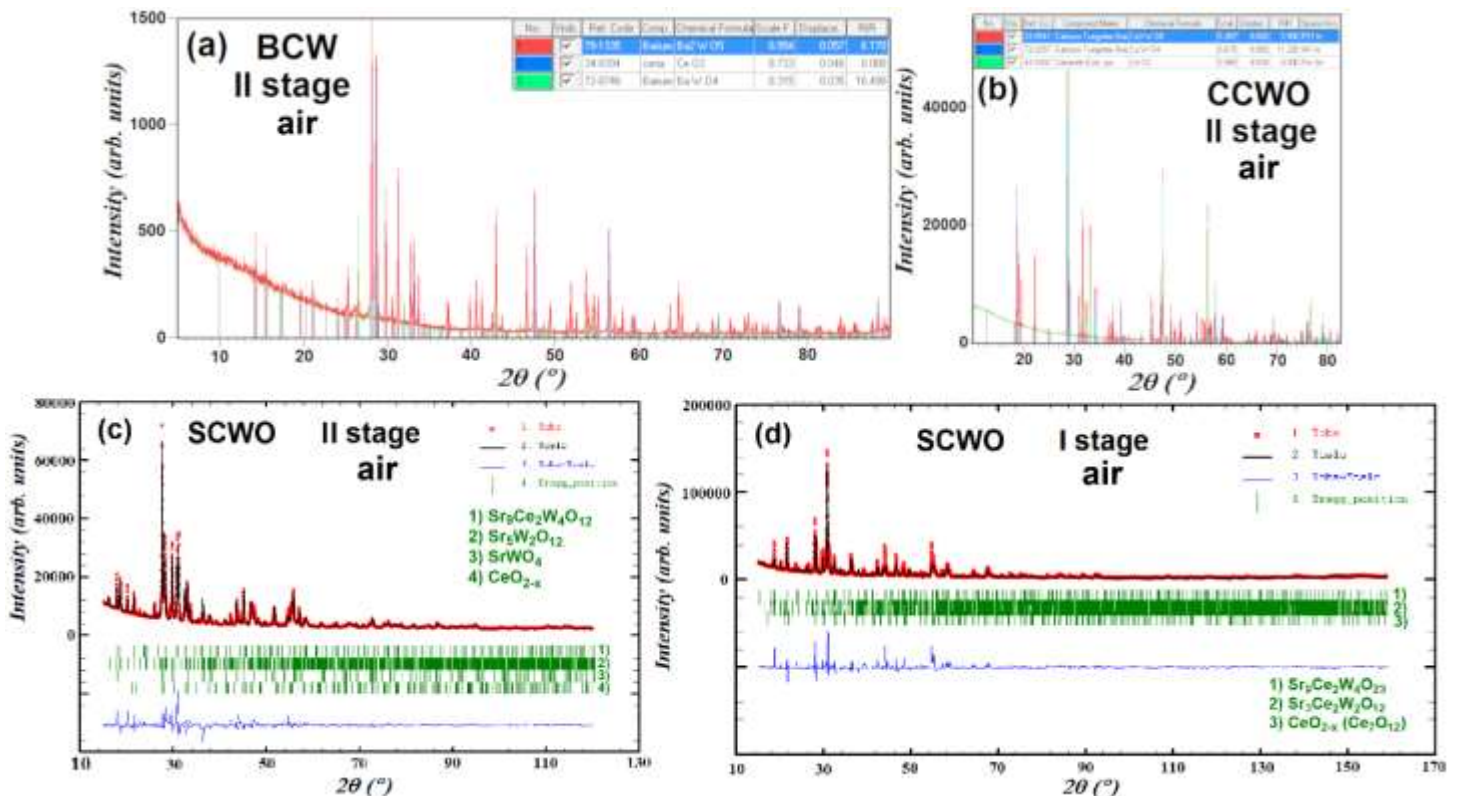


Figure 40 Powder XRD patterns after heating in the air (solid state synthesis without protective environment) at 2nd stage of decomposition temperatures for (a) BCW, (b) CCWO, and (c) SCWO while 1st stage (for the last-mentioned compound) is presented in figure (d). Fitting was performed using the HighScore databank.

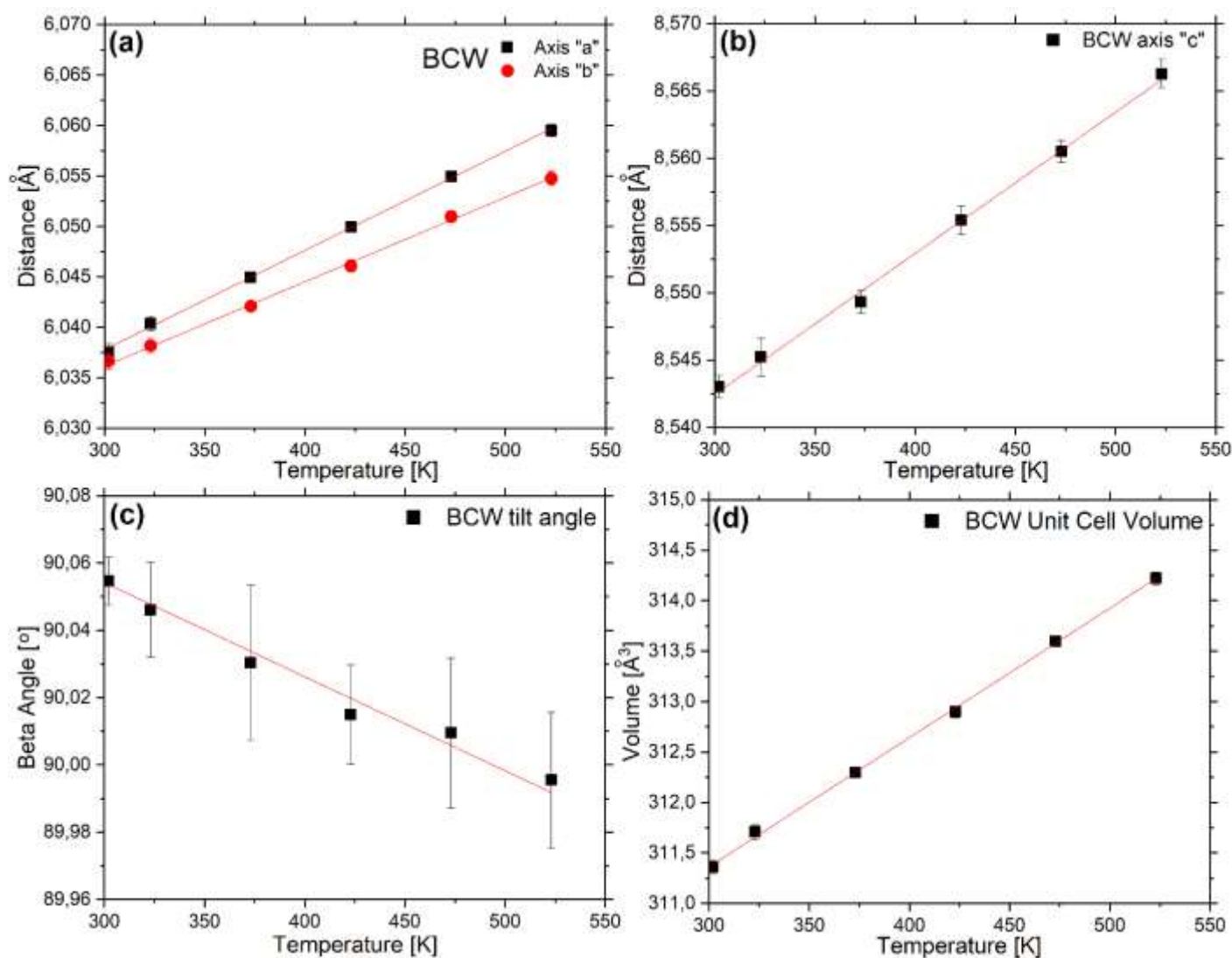


Figure 41 Various unit-cell structural dependencies determined through HT powder XRD measurements for the novel Ba₂CeWO₆ double perovskite. The crystallographic structure was investigated while heating in the air before any possible signs of decomposition were noticed. Figure (a) covers expansion along similar “a” and “b” axes, while (b) presents changes for a slightly longer “c” dimension. Graph (c) showcases a particular case of angular tilting being straightened towards a perfect, right angle, while in (d), a progressive, volumetric expansion can be seen. Red lines are linear fits, after which mean values were derived and stored in *Table 21*.

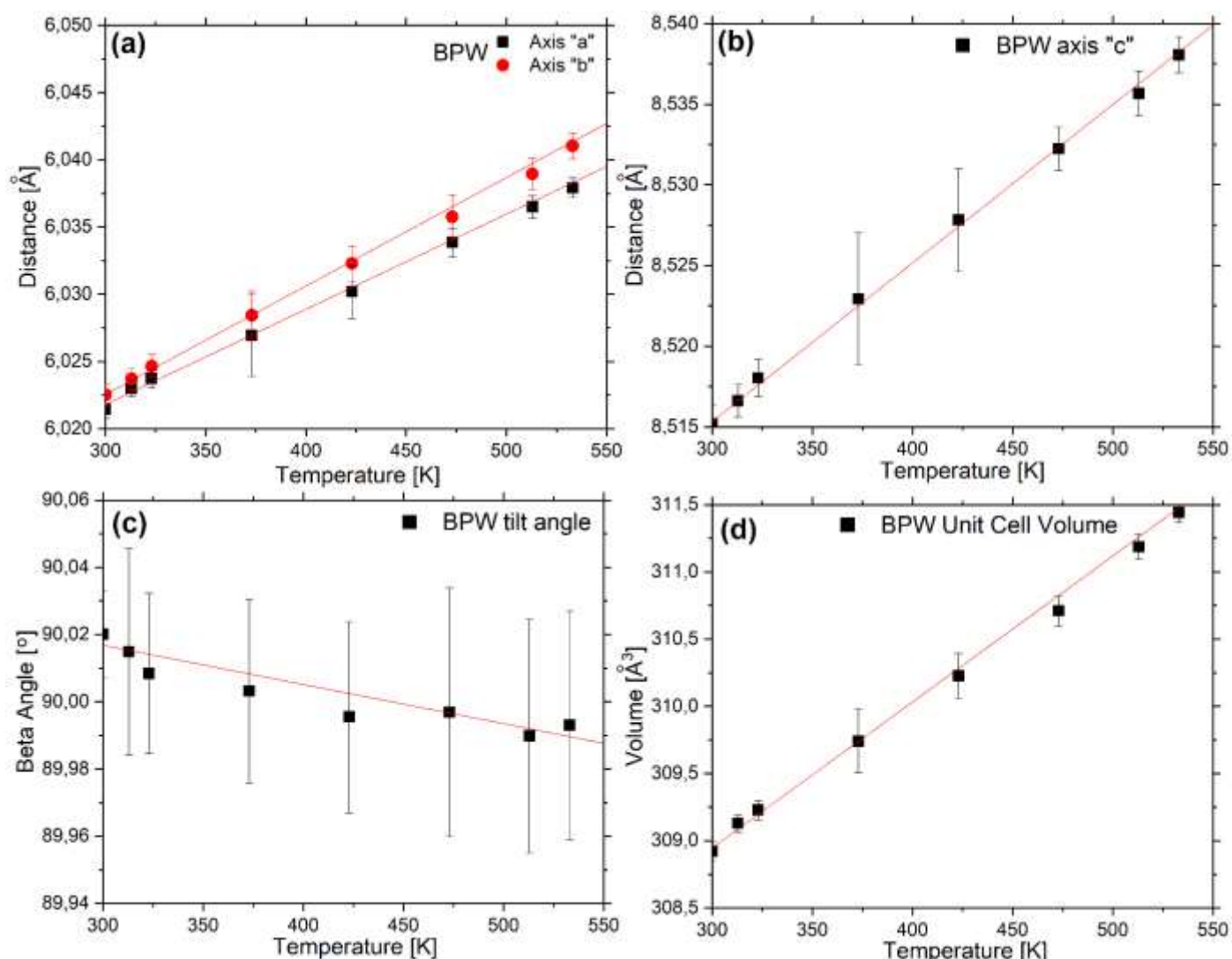
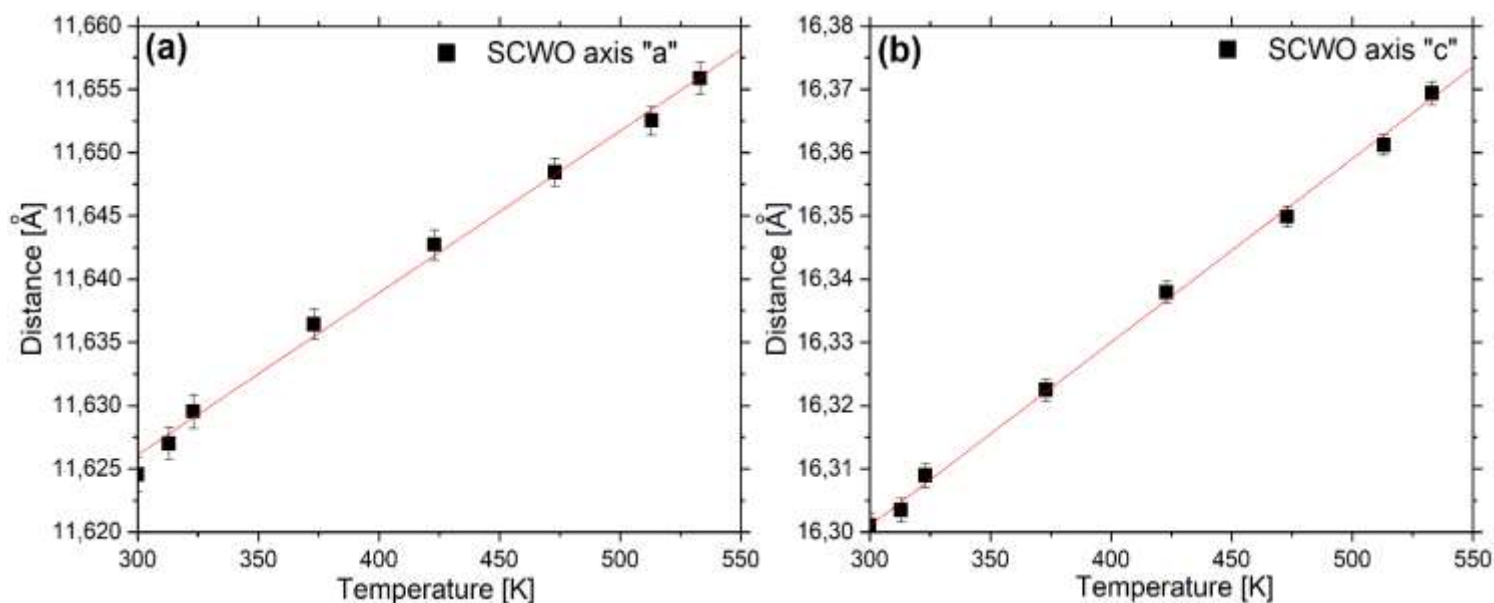


Figure 42 Various unit-cell structural dependencies determined through HT powder XRD measurements for the novel Ba₂PrWO₆ double perovskite. The crystallographic structure was investigated while heating in the air before any possible signs of decomposition were noticed. Figure (a) covers expansion along similar “a” and “b” axes, while (b) presents changes for a slightly longer “c” dimension. Graph (c) depicts a particular case of angular tilting being straightened towards a perfect, right angle, while in (d), a progressive, volumetric expansion can be seen. Red lines are linear fits, after which mean values were derived and stored in *Table 21*.



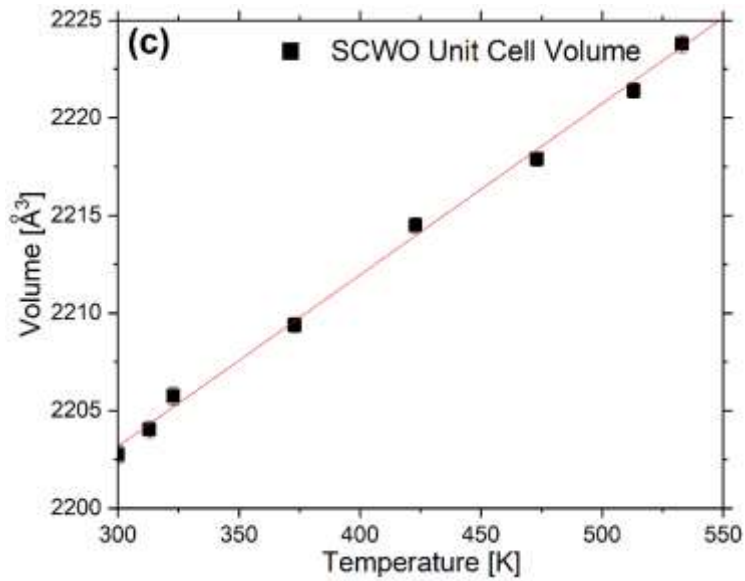


Figure 43 Various unit-cell structural dependencies determined through HT powder XRD measurements for the novel $\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$ quaternary perovskite. The crystallographic structure was investigated while heating in the air before any possible signs of decomposition were noticed. Figure (a) covers expansion along the “a=b” axis while (b) presents changes for a slightly longer “c” dimension. Graph (c) shows the same unit cell's progressive, volumetric expansion. Red lines are linear fits, after which mean values were derived and stored in *Table 21*.

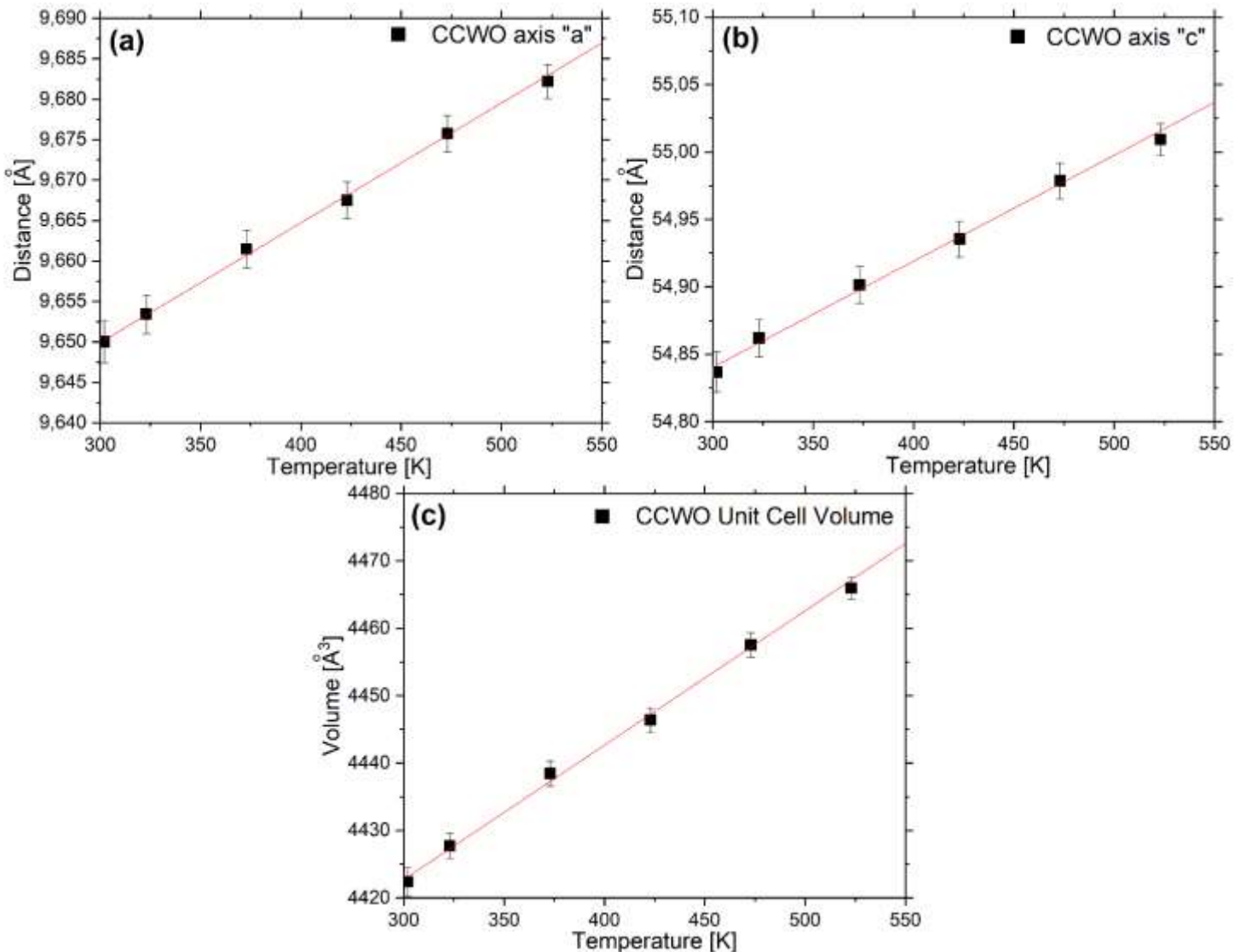


Figure 44 Various unit-cell structural dependencies determined through HT powder XRD measurements for the novel $\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$ ilmenite. The crystallographic structure was investigated while heating in the air before any possible signs of decomposition were noticed. Figure (a) covers expansion along the “a=b” axis while (b) presents changes for a much longer “c” dimension. Graph (c) presents the same unit cell's progressive, volumetric expansion. Red lines are linear fits, after which mean values were derived and stored in *Table 21*.

Table 21 Crystallographic expansion coefficients deduced from powder XRD data presented above in function heating while submerged in the air for all investigated materials just before their decomposition commenced (roughly above 500÷550K).

Sample	Axis direction [$\text{\AA}/\text{K}$] ($\times 10^{-5}$)			Angle [deg/K] ($\times 10^{-4}$)	Volume [$\text{\AA}^3/\text{K}$] ($\times 10^{-2}$)
	a	b	c		
BCW	9.83(0) \pm 0.14(0)	8.36(9) \pm 0.18(5)	10.48(7) \pm 0.28(6)	-2.79(8) \pm 0.17(7)	1.28(2) \pm 0.01(9)
BPW	7.08(0) \pm 0.17(2)	8.06(5) \pm 0.16(3)	9.83(3) \pm 0.23(2)	-1.16(7) \pm 0.10(5)	1.08(6) \pm 0.02(7)
SCWO	12.81(7) \pm 0.35(0)	---	29.01(7) \pm 0.48(7)	---	8.78(3) \pm 0.19(7)
CCWO	14.19(6) \pm 0.29(9)	---	78.63(1) \pm 2.17(7)	---	19.94(1) \pm 0.47(2)

As for barium DP's, the expansion coefficients seem quite unremarkable,^{449–452} but the presence of the progressive angle tilt somehow creates a solid base for a lowering symmetry phase transition at some temperature points – in our case, probably monoclinic noticed earlier at LT below 200K (see *section 3.1*).^{453,454} This also does not neglect any possibility of other transitions at HT if only these compounds were stable enough (even concerning inert conditions). Derivatives, especially CCWO, have substantially large coefficients but also possess giant unit cells.^{450,455} This fact, in comparison with abundant defects, vacancies, and partially-substituted A^{2+}/RE -sites, might also lead to some interesting properties via mending, rebonding, and lack of internal lattice-relaxation issues when exposed to stimuli in the opposite direction (like decomposition after compression or irreversibility while cooling from HT),^{456–458} however without any progressive tilt (angular dependency) the only thing which could change the unit cell (as LT-HT measurements seemed uneventful) is the packing order while axial expansion/regression transpires – those would however be only related to transitions within the same Laue group.⁴⁵⁹

4. Optical studies with and without RE dopants

4.1 Luminescence - excitation and emission spectra

Lastly, PL studies regarding possible future solar panels and sensory device applications. All this detailed research up to now set a firm base towards divulging possible phenomena that could be encountered in this most promising field of study. It has proven worthwhile since there is still a chance for those materials to benefit the industry despite them turning out to be weak, unstable phosphors of volatile/alternating structural nature. Some of them are thoroughly divulged upon, the others briefly speculated. The hypothesis suggests that with more sophisticated techniques, one could confirm a few of the most questionable properties but with more available time & resources reaching outside standard Ph.D. capabilities. Even though a lot must be done addressing all of the registered phenomena thoroughly, the chosen scope of literature and the given initial R&D approach must suffice for now.

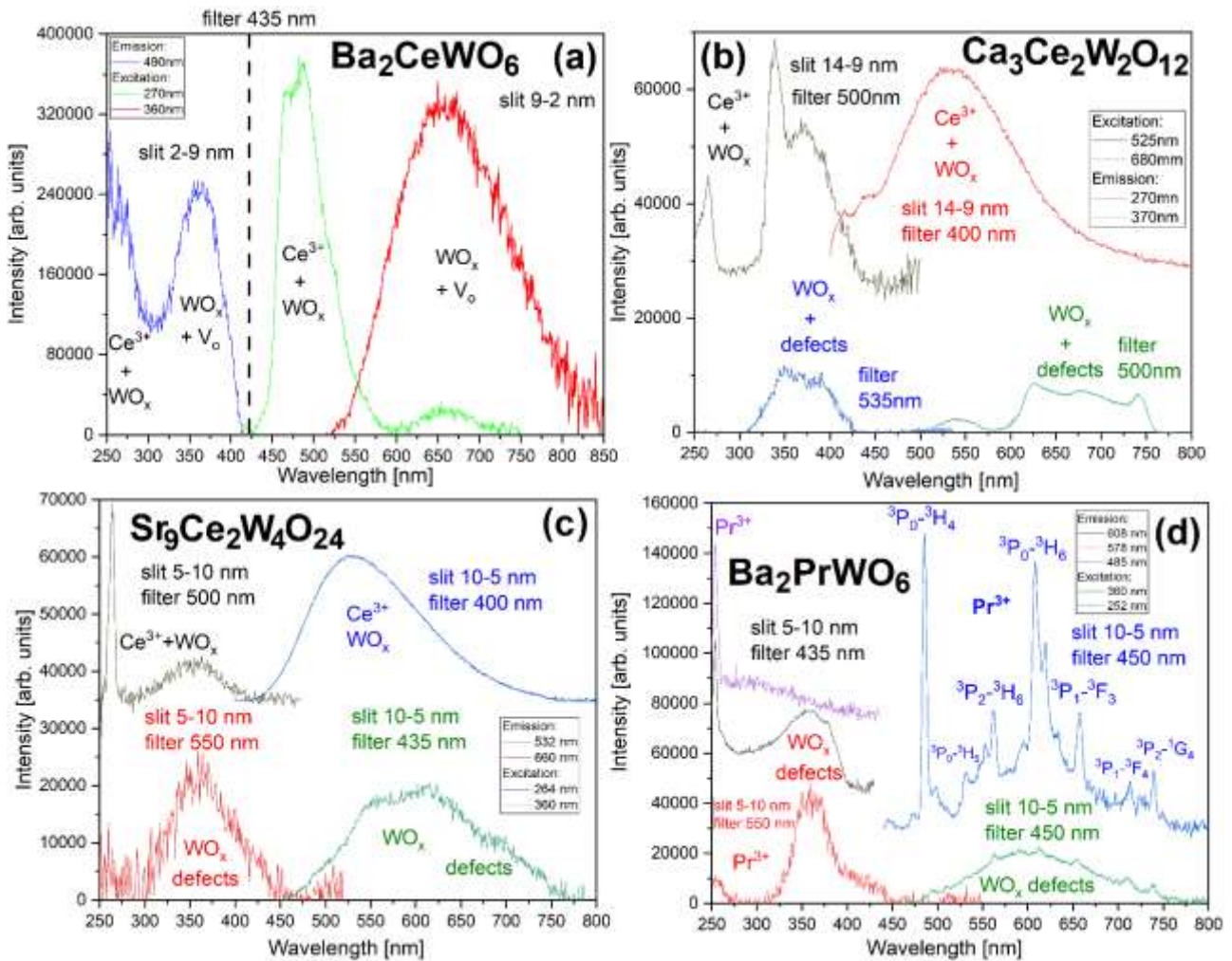


Figure 45 Excitation and emission spectra taken at ambient conditions using Xe lamp for all 4 investigated RE-tungstate materials: (a) BCW, (b) CCWO, and (c) SCWO cerates alongside with (d) BPW praseodymate. Adjusted monochromator's entrance and exit slits reflect the maintained, relative resolution of registered bands.

The excitation and emission spectra of tungstates (WO_x), visible above in *Figure 45*, are actually quite typical for many materials from this widespread family.^{2,5,8,202,460-465} Wide blueish-violet absorption band at around 360-380 nm emits further at an even broader, greenish-yellow 500-600 nm range. Sometimes V_o/O_i complexes introduce deep red components in low energy above +630 nm. This is even a typical, defect-related behavior encountered in our previous scintillator – bands positioning could change slightly, though with respect to SG & crystal field's strength (as seen in SCWO & CCWO).⁴⁶⁶⁻⁴⁶⁸ Since tungsten is a common component in every matrix of our novel materials, their appearance is anticipated & much appreciated, given that they could enhance future energy transfers considering the cooperation with RE ions. Speaking of which, RE ions like Ce^{3+} (*Figure 45abc*) often contribute to PL by enhancing the total intensity & broadening the range of excitation into UV.⁴⁶⁹⁻⁴⁷² The only issue might come forth from a sharpish (not in the case of BCW) excitation band at around 270 – 250 nm. In many oxides, it is referred to as a charge-transfer band that, if hit directly or above, could trigger an e^- transfer between RE-W and nearest O defects (not exactly being at 2- state).^{468,473-479} Pr^{3+} in *Figure 45d* has a very similar peak, closer to UV ~252 nm & behaves similarly, at least relatively, considering its multiple, lower electronic states creating much more sharp lines localized on the bulgy WO_x emission.^{470,480,481} (Ce, in comparison, has more pronounced non-radiative relaxation between higher and lower ^5D states entangled with W 5d/4f states and emits poorly towards ^2F levels).⁴⁸² More about that phenomenon later, but for now, it is worth noting that considering the general performance of these materials, it will be essential to work around those NUV bands (slightly above at ~228÷232 nm, as it turned out) since longer excitation

wavelengths, typical for such RE ions (such as 325 nm for Ce^{3+} or 430-460 nm for Pr^{3+}), did not perform as well in terms of acquiring any actual response from the samples as can be seen in supplementary *Figure A19*. Transitions registered above and beyond that NUV threshold were the only detailed & strong ones to hold any significance despite having smaller absorption cross-sections. They will also be crucial regarding planned down conversion with dopants (BCW, CCWO, SCWO: (sensitizers) $\text{Pr}^{3+} \rightarrow$ (mediators) $\text{WO}_x \rightarrow$ (activators) Yb^{3+} ; BPW: ($\text{Ce}^{3+} \rightarrow \text{Pr}^{3+}$) $h\nu$ donors \rightarrow (mediators) $\text{WO}_x \rightarrow \text{Yb}^{3+}$ $h\nu$ acceptors) – working with the constituents (BCW, CCWO, SCWO – Ce; BPW - Pr) and fighting against PL quenching issues (while changing their total mass % concentrations to optimize the set) as depicted in *Figure 46*. For true amounts of PL-active (3+) dopants, please refer to *Table 11* & *Figure 25* in *section 2.3*. All basic PL components were determined from all the convoluted spectra & verified through the literature.⁴⁸³⁻⁴⁸⁸ Each specific contribution is visible via colorful insets below. To avoid possible confusion and to make the graphs more legible in terms of upcoming PL qualitative analysis, all *Figure 46* transitions marked as RE (Ce – purple, Pr - blue)³⁺/ WO_x (orange) and V_o+O_i defects (red) were summarized in comprehensive energy schemes shown in *Figure 47*.

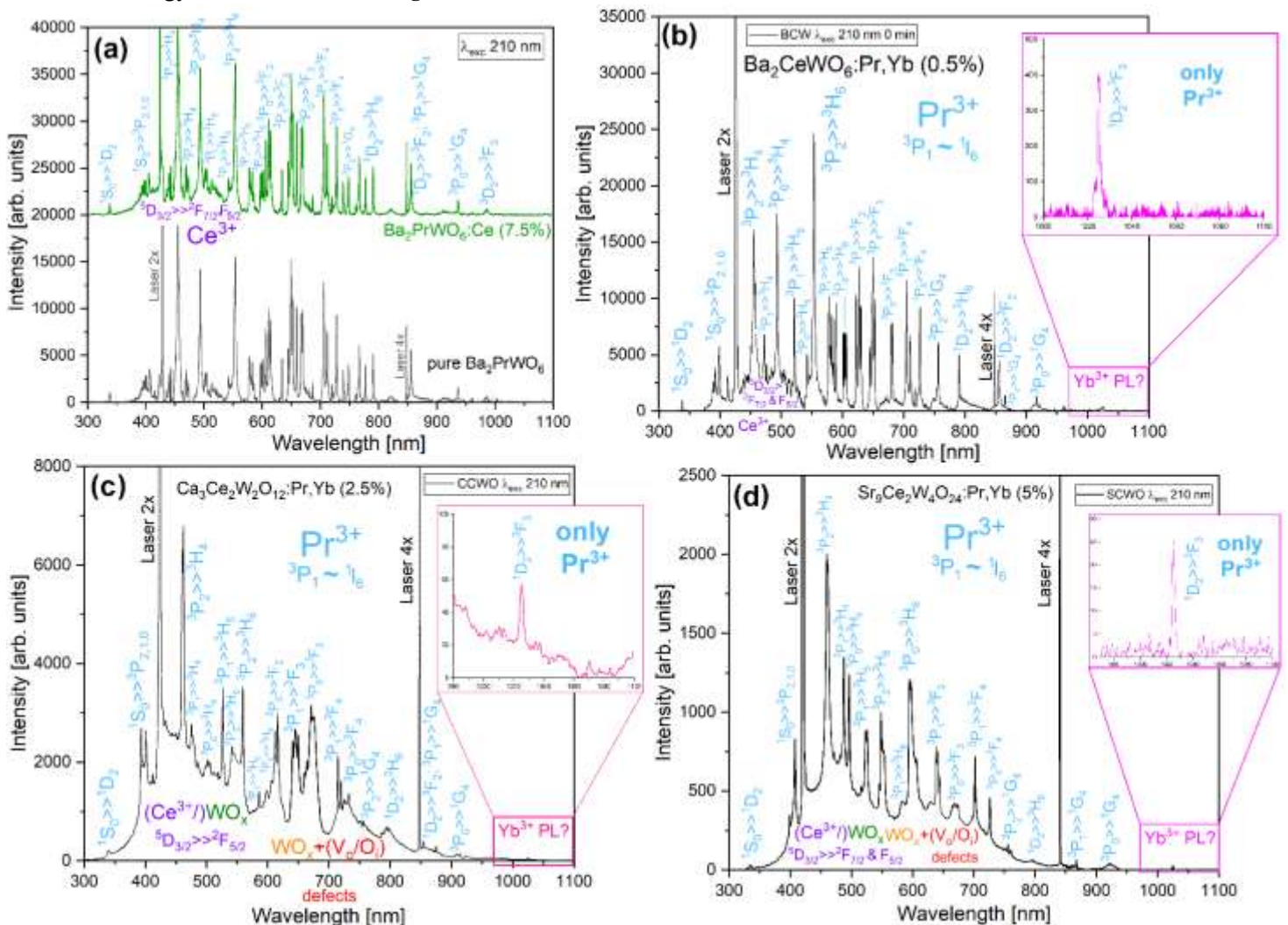


Figure 46 Fully described photoluminescence spectra of all RE^{3+} doped tungstates excited with 210 nm laser light under ambient conditions: (a) pure BPW double perovskite compared to BPW matrix hosting 7.5% of Ce, and the rest (b) BCW, (c) CCWO & (d) SCWO containing 0.5, 2.5 and 5% (Pr, Yb) respectively.

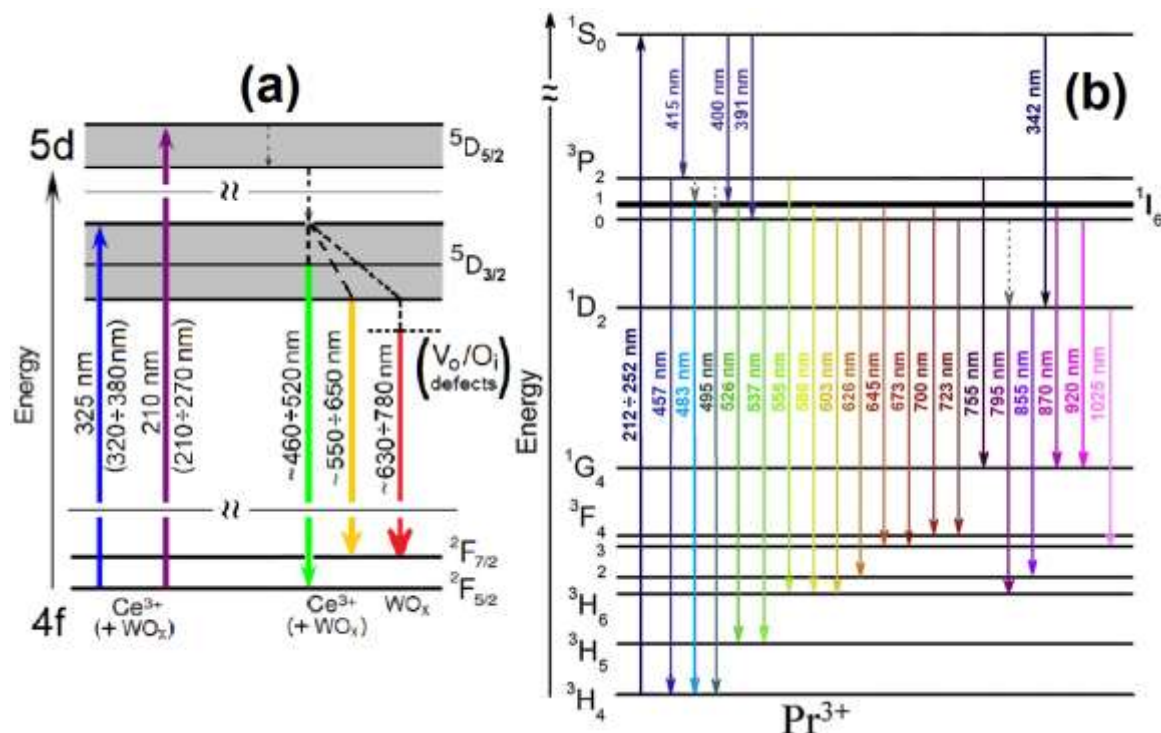


Figure 47 Energetical schemes of observable: (a) Ce³⁺/WO_x and (b) Pr³⁺ emission bands.

As Ce³⁺ emission is unremarkable, it would not be much elaborated upon; however, an interesting fact is that one can excite easily with weaker (not essentially narrow) near-violet light, but, Pr³⁺ requires more potent (intense) UV light to have any response at all. That last (preferably 215 or 232 nm) source brings us to the bottom of the ¹S₀ level in Pr-related compounds where e⁻ most likely relax to lower ³P states. The mechanism of relaxation is probably fully radiative since all ¹S₀→³P_{2,1,0} bands are visible just before double (2×) laser lines in all Pr-doped materials held in Figure 46, but they are quite weak – especially ¹S₀→¹D₂ which is not always or barely visible around 342 nm.^{16,489} Most energy is presumably used further in a cascade of deexcitation processes towards much lower levels. Here, worth noting is ¹S₀→³P₂(...)³P₀→³H₄ passage resembling quantum cutting in NUV-blue region close to 400-450 nm. Also, closely overlapping ¹I₆ and ³P₁ states - differentiating between one another would be quite a daunting task, so a broader ³P₁ mark was used instead as an abbreviation. Together with the ³P₀ state, they would theoretically be the sensitizing key of Yb (activator) by quantum-cutting red emission lines towards ³H₆→³F₄ levels.

Furthermore, the most problematic aspect was encountered here - the appearance/composition of the area shown in graphs as their pinkish insets. The last, almost vacant region where Yb³⁺ should have been, given theoretically planned energy transfer (Figure 2) and aforementioned quantum cutting (but instead of blue should be the red region).^{294,297,298,306,307,490,491} As it turned out in Figure 48abc, only direct excitation shows anything at all and, what is worse, only weak peaks at helium-level temperatures where they should be quite radiant.⁴⁹²⁻⁴⁹⁴ Although an interesting stark-splitting effect was spotted and several dependencies were observed: one, unexpected (6>4) being confirmed by two different pulsed laser wavelengths – visible under 917 but not during 978 nm excitation as seen in Figure 48d followed up by selection rules supported by weak kT effects; and the second – BCW time-resolved uprising of slowly strengthening weak Yb peaks.^{492,493,495-497} This ion did not act as a theoretically planned hv activator. The actual culprit was surprisingly not only weak effectiveness and unfortunate coloring of the samples (derivatives) but also a time-resolved Yb²⁺→³⁺ evolution (barium double perovskites) as seen in Figure 49. The first hint about the latter phenomena is suggested in the next chapter, where a direct, internal RE↔W charge transfer appears to be dumping electrons from RE to W during subtle NIR illumination, but, as it will turn out later on, another chemically oxidizing agent, localized in the interstitial gaps of the crystallographic plane, might be responsible for such oxidizing, PL-enhancing behavior. Inversely, it might also be accountable for the reduction of Yb during the synthesis by feeding its electrons through shallowly placed bands while being submerged in Ar:H₂ medium^{303-305,498,499}

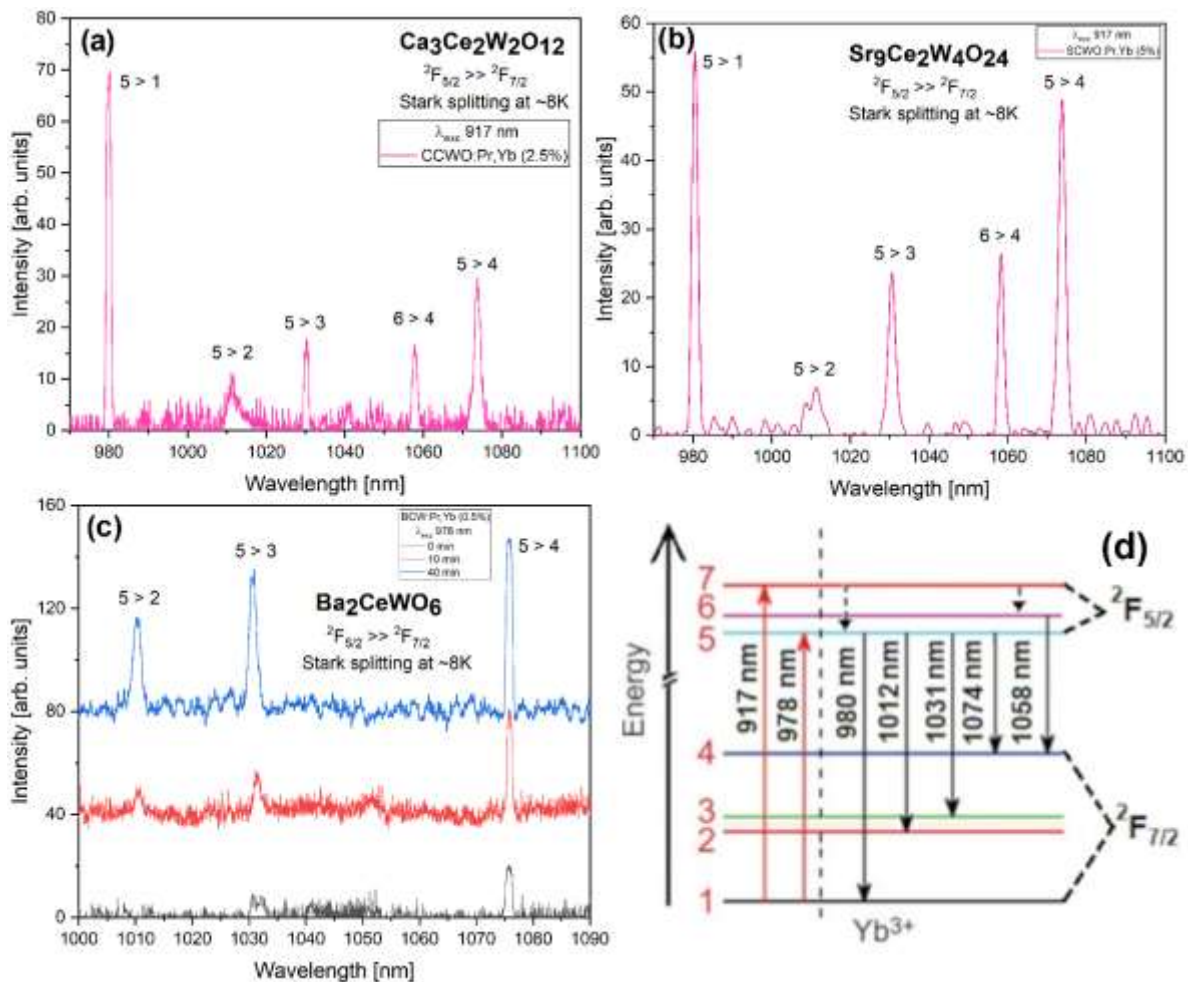


Figure 48 Weak but stable Yb^{3+} emission registered solely under characteristic NIR light excitations for tungstate derivatives at LT: (a) CCWO, and (b) SCWO featuring prominent Stark splitting under 917 nm wavelength. Unexpectedly strengthening time-dependent emission of BCW double perovskite is shown in figure (c) using a slightly weaker 978 nm excitation. Time refers here to the relative period of 210 nm Pr^{3+} wavelength exposition measured from the start of the experimental cycle. To emphasize exclusive Yb^{3+} PL bands, an energetical scheme of observable transitions was constructed in figure (d).

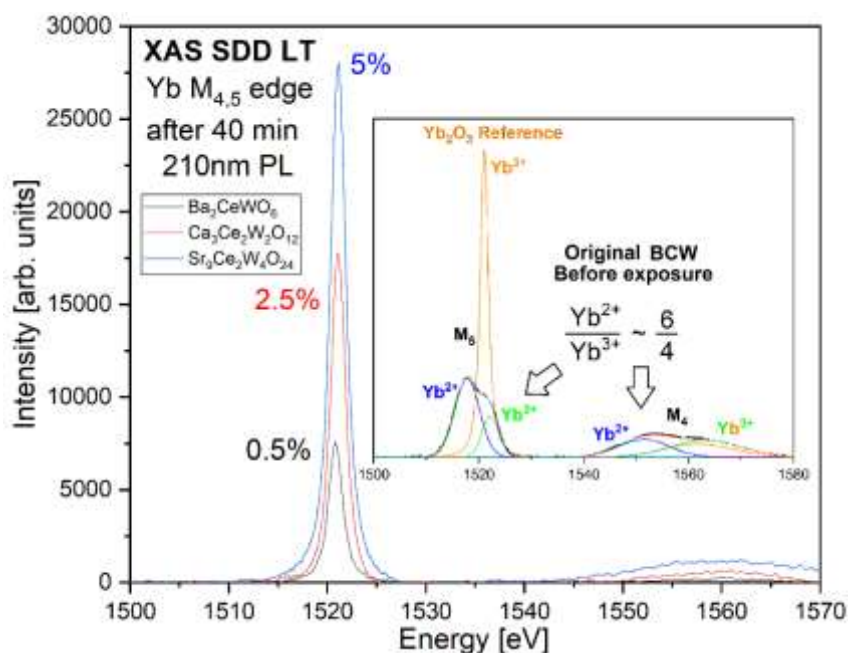


Figure 49 XAS Yb LT M-edge spectra presenting prominent changes in charge states after 40 min of exposure to UV $\lambda_{\text{exc}} = 210$ nm light in comparison to the original composition & Yb_2O_3 reference registered at RT (as inset from Figure 25b shows).

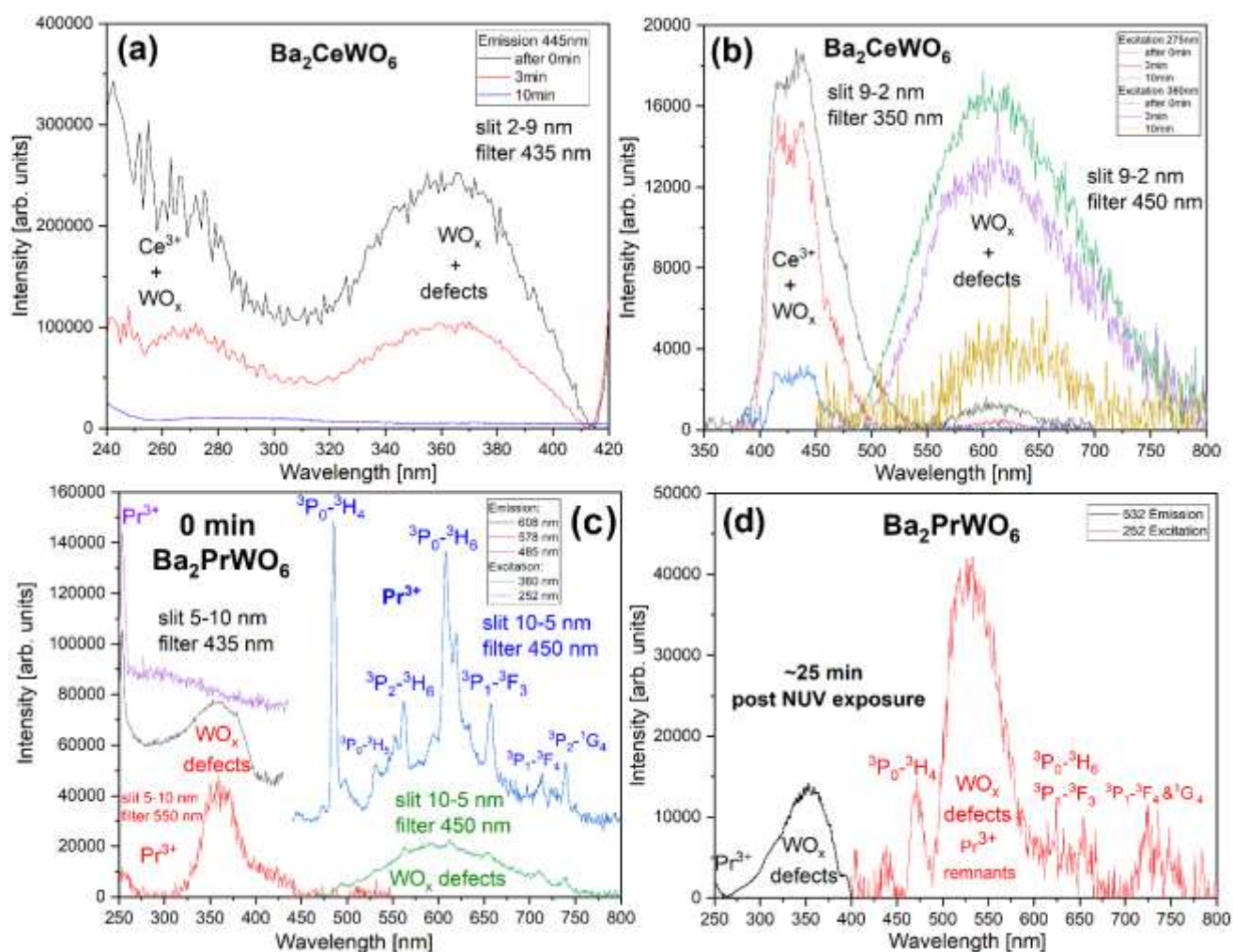


Figure 50 Time-resolved evolution of excitation end emission spectra in barium double perovskites prone to charge-transfer/photobleaching effect: simplistic (a) excitation and (b) emission bands of BCW are presented on the top while more complex BPW on the bottom (c). Post-NUV exposure data was presented on separate figure (d) to avoid illegibility final.

4.2 Charge transfer & photobleaching phenomena

Ytterbium bands in *Figure 48* were not actually the first signs to pinpoint us in the direction of charge-transfer phenomena – they were just luckily encountered cases of strengthening PL in contrast to previously recorded excitation/emission spectra in *Figure 50* while searching the area for stronger, repeatable signals. Only by considering such bizarre behavior were the measurements once again repeated, and some key features were quickly distinguished during weak, time-resolved illumination with the Xe lamp. It was first considered an equipment error while investigating specific excitation/emission bands. However, this assumption was quickly discarded since moving from place to place/exchanging the sample (placing a new portion) had similar outcomes. As PLE & PL spectra in *Figure 50* indicate, in both barium RE tungstates, all well-distinguished bands dropped with time – some faster (less detailed BCW in *Figure 50a,b*) than the other (BPW - *Figure 50c,d*) and did not recover if left in the dark. That was strange, considering that acquired derivatives (SCWO & CCWO) did not show such tendencies and remained the same, as seen in *Figure 45*. The next logical step was to check what happened to the PL-active ions since the sample visually did not change. As X-rays and EPR spectra of exposed

DPs samples shown in *Figure 51*, something regarding B-site ions indeed did change. When compared to the original (not exposed to light), signal-rich XPS and XAS spectra in *Figures 19 & 20* from previous *chapter 2.3*, it becomes evident that previously balanced matrices hosting a mix of $\text{RE}^{3+}/\text{RE}^{4+}$ cations were drained entirely out of electrons changing into PL-inactive (Ce and Pr) $^{4+}$ species while W^{5+} ions were flooded by them turning completely into W^{4+} . EPR seems to follow such claims when compared analogically to unexposed samples in *Figure 34, chapter 3.3*. In BCW, the Ce^{3+} signal completely disappeared at LT, and, given the activation mechanism of RE ions, that could only prove its conversion to Ce^{4+} . Pr^{3+} 's case is slightly more complicated since it is visible only in horizontal fields. Pr^{4+} would be visible under given perpendicular fields only if the signal was not jammed by antiferromagnetic phase transition below $\sim 90\div 110\text{K}$. RE bands are already broad & weak thanks to polycrystalline, anisotropic domains, and defects, but at least some evidence could be told what is actually happening in BPW at RT – after bathing in reducing 1:1 Ar/H_2 conditions, both Pr^{4+} and O_i would disappear but, RE are not visible at that temperature point so these sharp signals are definitely related to encapsulated O species. As mentioned, oxygen could be partially involved in the charge transfer/oxidization procedure. Here, after NUV irradiation, interstitial oxygen or dangling O^\cdot species disappear (just like ozone) and presumably integrate with the lattice. This process makes the matrix enriched in oxygen – typical for antiferromagnets - where new O^{2-} ions are presumably also created by devouring electrons from Pr^{3+} , making the matrix further PL inactive. Thus, the aforementioned drop in excitation and emission spectra. There is most likely some competition between O and W ions since the most unstable W^{5+} (being already in the minority) disappears almost instantly.

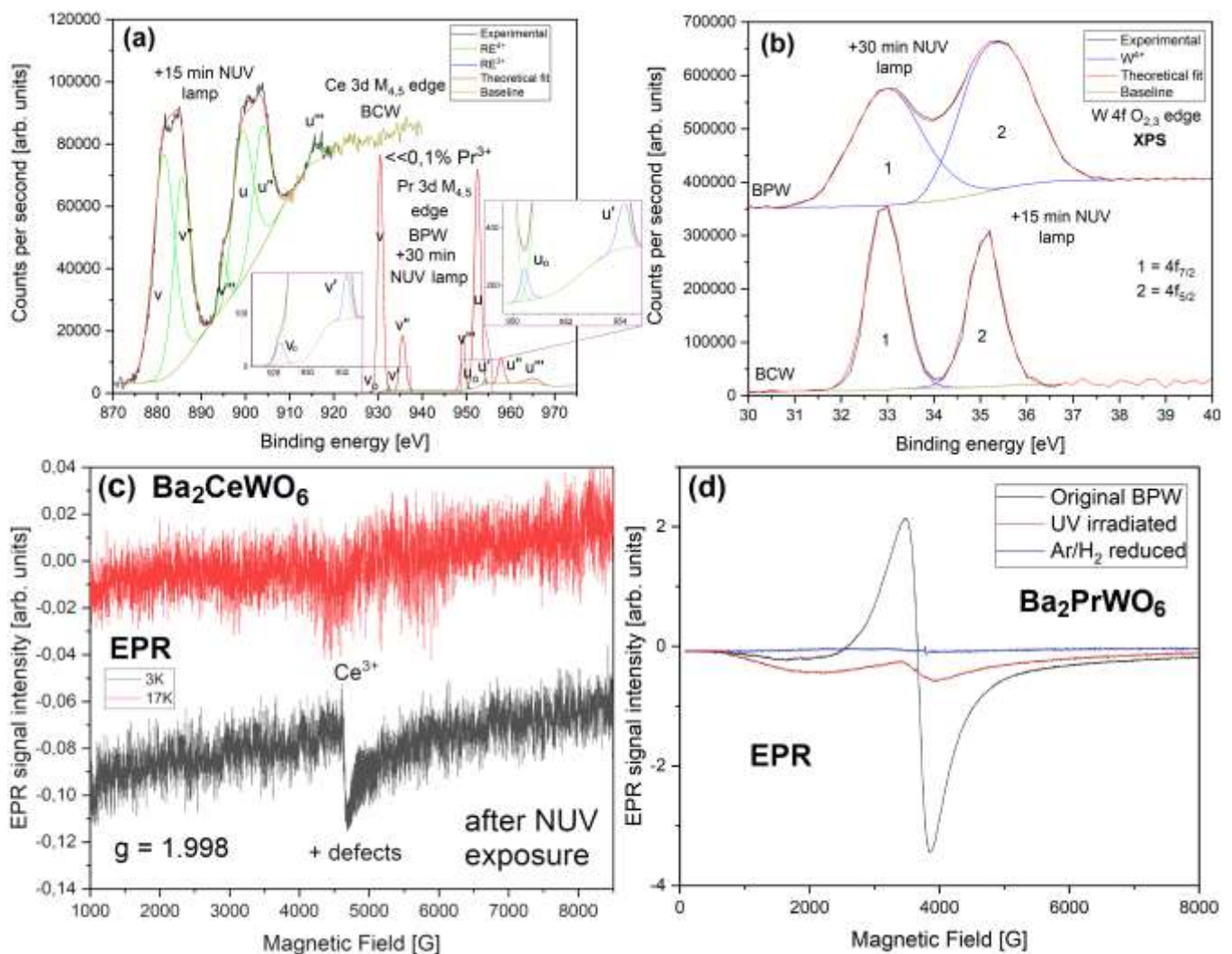


Figure 51 XPS ambient spectra of both, undoped tungstate double perovskites (Ba_2CeWO_6 & Ba_2PrWO_6) investigating changes in ionic states of crucial B-site ions after brief (15 & 30 min respectively) NUV exposure: (a) Ce/Pr 3d spin-orbit multiplets, and (b) W 4f doublet shells. EPR was done in consecutive order regarding the same issue for (c) BCW at LT & (d) BPW but at RT.

Further discussion revolving around changing PL requires some clarification regarding the actual chain of events and the causality of facts that transpired during this work. Since the data presented here are ordered arbitrarily, to truly understand the final outcome, one must piece together scattered evidence from:

- RT PL spectra at weak, Xe lamp excitation in any gas atmosphere (done first & already presented above in this and previous *section 4.1*),
- HP PL experiments in inert PTM (Ar) using lasers (described below),
- LT PL in vacuum-pumped or inert (He) conditions using lasers (described further below),
- RT PL spectra at strong laser excitations in the air (elaborated in the next *section 4.3*).

Starting with HP PL, there should be no surprise why such in-situ investigations presented in the appendix *Figure A20* came up short with not many remarkable results and were dropped not long after considering the aforementioned fast PL quenching.^{161,500–502} While measuring Raman spectra with simultaneous access to 325 nm excitation wavelength (stimulating broad $\text{WO}_x + \text{V}_o$ bands in undoped samples), the changes that transpired during compression (in air-deprived, Ar PTM) were quite subtle, noisy, and rapidly diminished to almost 0 intensity. The emission in more stable derivatives persisted until the investigable range's top pressure limit/chosen GPa aim. While approaching the region of possible phase transition (*Figure A20a,b* SCWO ~15 GPa, and *Figure A20c,d* CCWO ~9GPa), both sets of data swing slightly in positions changing the progression trend from weak blue to red-shift and vice-versa when going back. CCWO's PL even increased slightly until 12 GPa, which is the presumed endpoint of the eventual change to $C2/c$ or $C2/m$ SG.^{71,72} At least both changes are reversible and to a similar degree considering the noted hysteresis on Raman spectra. In *Figure A20e,f* the BCW double perovskite is presented as a contrasting issue. No phase transition was observed, but the same problem, consistent with results acquired in *Figure 50a,b* takes place. Not only the PL is quickly diminished during compression but, it is also irreversible. Moreover, it is indefinitely quenched when exposed to radiation at the same place for an extended period of time (while maintaining a certain level of pressure). That proves that the charge-transfer process is at least partially responsible and takes place mainly during light stimulation. Not much can be experimentally said & done to estimate how HP influences such a phenomenon (it definitely does, as can be seen in literature)^{503–507} but, for sure, it seems hopeless to look into energy-transfer and band-gap-related phenomena if no actual response from the sample is received. Neither full range of PL interest could be registered and analyzed before photobleaching commences, so the only natural way to progress further is to drop HP & investigate the region at lower temperatures where the emission should naturally increase.^{508–513}

While exposing undoped cerates at LT (in a function of heating while being vacuum-pumped in helium-filled cryostat), nothing unusual happened in the formerly presented region of spectral interest (WO_x), as seen in *Figure 52*. If there was any PL, it was weak, broad, and bared the same 2-component ($\text{RE} + \text{WO}_x$, and $\text{WO}_x + \text{defects}$) characteristics as those seen before in *Figure A20*. Only results for the two most extreme temperatures were shown for derivatives regarding their borderline cases in *Figure 52a,b* (solely related to intensity changes). PL spectra expectedly gain strength when approaching cold helium temperatures. Especially BCW exposition had an

interesting outcome, so it was expanded upon in *Figure 52c* to d. At the lowest temperatures, PL was relatively stable (at least time-wise compared to RT quenching). Presumably, the electron charge carriers were frozen (as lack of any features at LT XAS TEY measurements shown in *Figure 33*)⁵¹⁴⁻⁵¹⁶ or the phase transition had some stabilizing effect on PL-active ions ($\text{RE}^{3+}/\text{W}^{6+}$) - probably by manipulating fixated states of the electronic structure.^{363,517-520} Considering that outcome, those cryogenic conditions could scientifically enable a more detailed approach to efficiently investigate other bands while using more befitting, intense, but pulsed illumination (on OPO). For sure, the duality and the constant presence of registered signals together with their significant improvement/enhancement when joined with Ce^{3+} (now not only dopant but constituent) signifies that they are at least partially tied/codependent, as stated in our previous work where it was still a dopant.^{2,521} This general behavior will benefit the amount of gained intensity and serve as a reference background while stimulating more complex matrices containing other RE (Pr, Yb) dopants for any signs of efficient down-conversion.

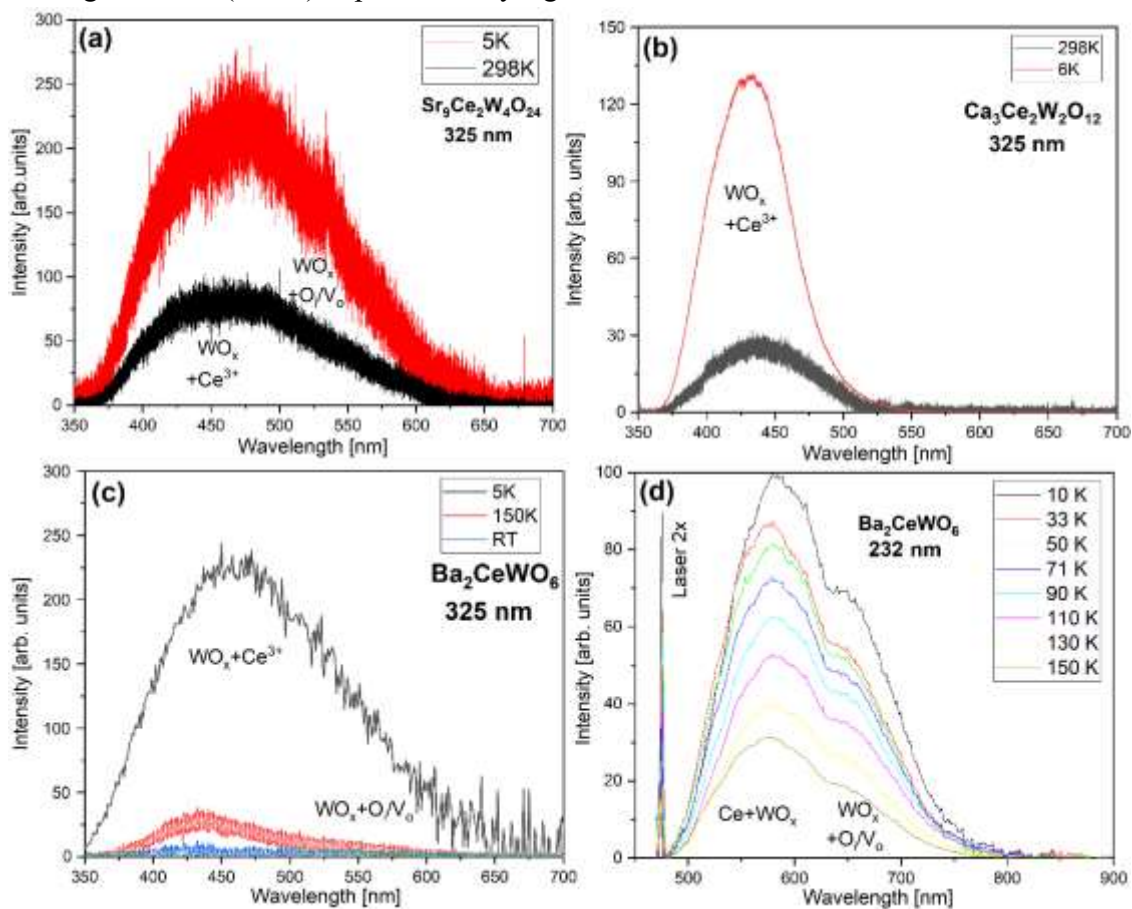


Figure 52 LT PL spectra of undoped cerium tungstates in the heating function present only weak WO_x emission bands influenced by various structural defects or Ce^{3+} (double component/shoulder feature). Figures (a) & (b) present both SCWO & CCWO derivatives, respectively, with (c) BCW double perovskite measured under NUV 325 nm laser light while being investigated in situ by Raman spectrometer in the most extreme temperatures. Figure (d) shows the same last compound under more effective 232 nm excitation with more detailed, intermittent steps.

In Pr-doped samples, the situation indeed became more complicated. While derivatives' PL decays smoothly, as *Figure 53a,b* shows during heating from helium-level temperatures up to ambient conditions, barium double perovskites in *Figure 53c,d* possess two types of behavioral patterns that differentiate above and below possible phase transition region (150-100K) – one

addresses the recoverable PL below 90K (as presented in supplementary *Figure A21a*), and the other, unrecoverable above 150K (*Figure A21b*). The first situation, however, has a limited number of cycles (3) that could be reapplied while moving back and forth between each borderline temperature – *Figure A21c*. That means that whatever is responsible for LT $\text{Pr}^{3+}/\text{W}^{6+}$ transformation and recovery during illumination is not persistent, limited to a certain degree, or at least its fueling sources (traps) are irrecoverably depleted. There are most probably two phenomena working almost simultaneously at the same conditions: a broad low-symmetry SG transition (presumably to $P2_1/n$ SG around 150K) which was already reported in the literature considering several DPs;^{183,356,360,362} And, presumably, the indirectly-observed magnetic transition set tightly somewhere in between 120-90K. The latter sometimes could significantly impact overall optical properties at a varying degree. By this, I am not only simply referring to trivial changes in certain B-site states (charge-transfer and freezing e- carrier effects that manipulate SGs, local Bruloin zone, and effectively the whole band gaps structures)^{237,365,368,478,514–516,522} but also imply the existence of more subtle phenomena like metal-insulator (Mott-Hubbard-like) antiferromagnetic transition changing the whole electromagnetic arrangement within the samples which would be, in particular, BPW.^{397,398,400,401,514} That would surely add much commotion (additional factors) amongst already data-rich experiments between all the 150-100K related phenomena, but at least it would confirm the change to be second-order in origin. There, the general temperature quenching stops and the blue region of emission loses its intensity abruptly in contrast to the red spectral zone. Prominent erratic swings in regional PL intensity also appear. PL barely recovers when cooled down from 150K back to ~10K, as depicted in supplementary *Figure A21b*.

Contrary to the 90K dependence in *Figure A21a*, it looks almost the same as before - weak, noisy, maybe closer to 130K. That provides clear evidence that all previously-described PL problems are directly connected to the ionization through a specific CTB illumination. Certainly, apparent conditions (like current temperature kT effect) also apply - at least to some extent, given emerging SG/electromagnetic transitions interfering with the width of the local band gap. Unfortunately, complex, structural fluctuations (i.e., mostly from stimulating the aforementioned CTB between RE & W) have yet another factor to be acquainted with - a time-resolved charge-transfer component.^{370,371,479,514,517,523–527} That last phenomenon indeed progresses parallelly within the sample. Sometimes at such a rate that its contribution is difficult to assess quantitatively. Since ionization happens so fast at LT (within 1 or 2 seconds before some discrete equilibrium is established as observed by the microscope-aided eye), the PL spectra presented in *Figure 53* can only refer to the state where that last phenomenon progressed thoroughly and already reached its own balance at certain conditions.

The issue further evolves while using different media/gases in which the substances are submerged. That is why most PL experiments performed in the air on Barium DPs were cut off above 150K (in contrast to derivatives or other LT measurements where samples were closed in controlled, evacuated, He-filled cryostats). Air's (crucial yet destructive) influence via oxidation and local overheating of O_i creating additional V_o complexes in polycrystalline lattice further complicated already complex interpretation – the number of stimuli needed to be reduced in order to analyze what was actually happening. On the other hand, the air's impact later added an interesting twist to overall observable PL effects. More about this in the next *chapter 4.3*. For now, it is important to be aware that there is some clear evidence of 2 simultaneous processes influencing PL. One is time-dependent, directly related to CTB stimuli (affecting not only $\text{RE} \leftrightarrow \text{W}$ but also O_i/O^- presence). The other, temperature-activated or rather dependant transition through some sort of barrier from shallow RE (Pr^{3+}) states up to CB and ionizing RE^{3+} ions from shallowly placed traps (presumably oxygen related).^{370,384,387,475,478,479,528–531}

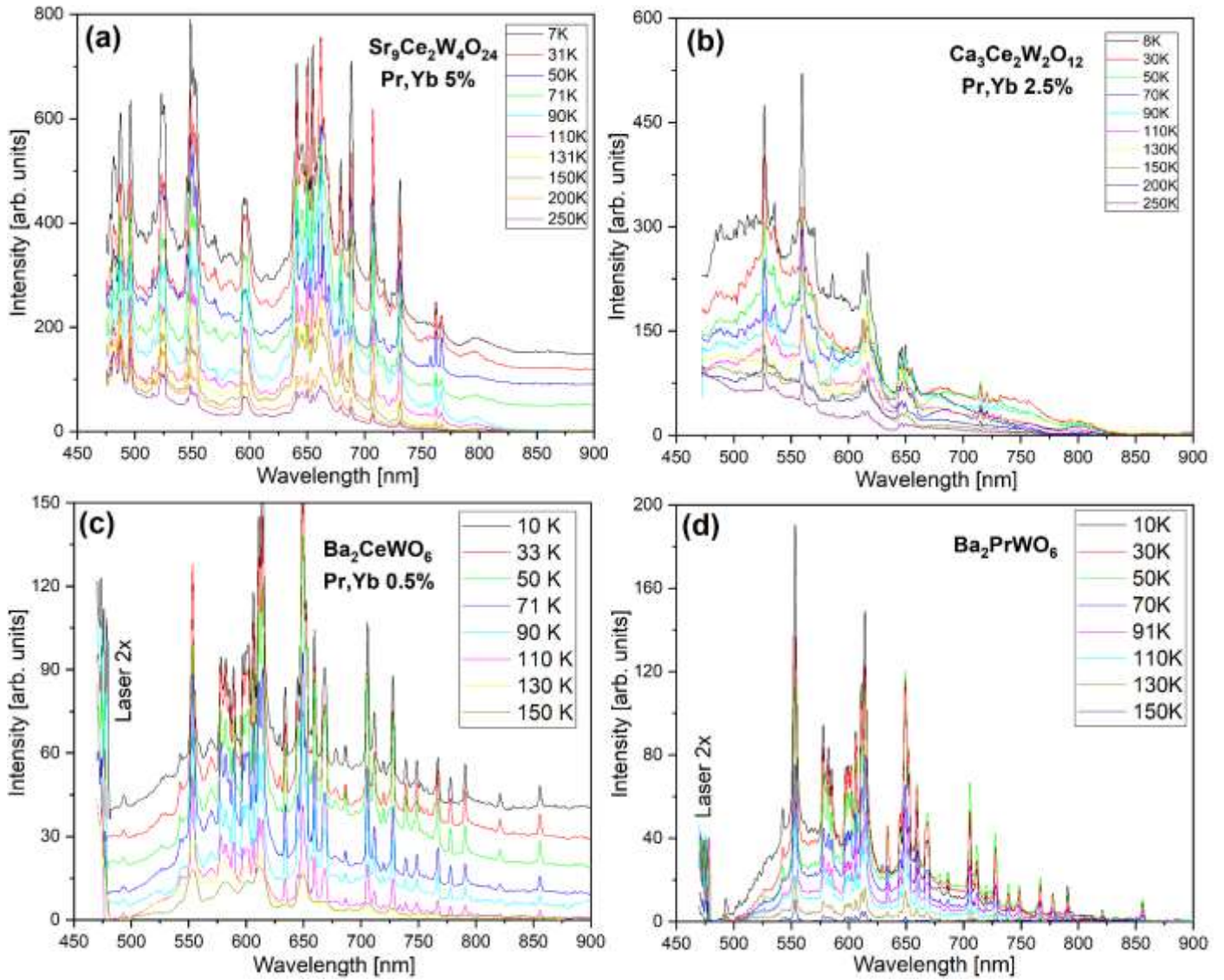


Figure 53 Low-temperature PL dependency of all (Pr,Yb)³⁺ doped materials: derivatives (a) SCWO, (b) CCWO, and double perovskite (c) BCW compared to (d) pure Pr archetype BPW. Data was collected in the function of heating towards RT with $\lambda_{exc} = 232$ nm from the optical parametric oscillator.

From the last-mentioned process and its thermally-quenched LT PL dependencies depicted in *Figure 53*, unique activation energies (E_a) could be obtained. They could be useful to interpret what is actually happening with the electrons – ionization through some sort of thin barrier from shallowly placed RE³⁺ states or traps by exciting them up to the CB. They could be roughly calculated from empiric equations such as Arrhenius law⁵³² and its modifications depicted in *Equation 4* below:

$$\text{Equation 4: } \frac{I(T)}{I_0} = \frac{1}{1 + C \exp\left(-\frac{E_a}{kT}\right)}$$

Such an equation implies balanced, stable materials in their chemical (nonreactive) equilibrium state during illumination. Everything we excite with light from the ground to the excited states should return, whether as radiative or non-radiative (temperature-dependent) recombination. Both during and after turning off the light, we should have a constant concentration of the admixture, independent of the temperature. So, given the law and assumptions derived from it in the appendix (*Equations A1*), they also imply that if there are more energy levels (i.e., from the impurity), the whole process must concern solely the same ion. Considering our materials' apparent interactions with the conduction band, that is not it.

The other scenario and set of equations are based on the case of one dopant in the only available charge state ionized to another with no traps present in the way. The equation ends in a quadratic state

incorporating concentration values, even in the most straightforward cases (as shown in supplementary *Equations A2*). $r(T)$ very rarely depends exponentially on temperature; it is usually a dependence to the power of X (X usually = 2), but it is again not the case since when the light is turned off, all electrons return to the impurity, i.e., concentration $N=N_0$.

The third case is related to ionization from the excited states, including traps (V_o , O_i , W^{5+} , RE ion shallow states), so justifiably, our case. Still, many assumptions and simplifications must be made, but at least there are as many traps as imbalanced RE^{3+}/W^{5+} & RE^{4+}/W^{4+} ratios state. Amongst them, interstitial oxygen is placed. The EPR O^- signal is extinguished after UV illumination at room temperature, but surely not completely. The traps are presumably deep and not thermally ionized at RT since the pairs would be processed on their own towards one of the presented forms. If we assume that all the traps are M_0 and they are not ionized by light, then the formula derived from all the presented in supplement *Equations A3* would finalize themselves in the form of *Equation 5*:

$$\text{Equation 5} \quad N(T) = N(0) \left[1 - \beta \exp\left(-\frac{E}{kT}\right) t_0 \right]$$

Where $N(T)$ is the concentration of a specific ionized dopant in certain conditions. $N(0)$ is the whole concentration of specific ionized dopant available at the start/beginning of the experiment at the most extreme temperature in our case. β is the probability of ionization (in our case, high). k – Boltzmann constant [meV/K]. T – Temperature [K]. t_0 – the time of reaching charge transfer equilibrium within a specific temperature (in our case, much shorter - 3÷5 seconds - than the measurement lasting approximately 1 minute); and finally E – the desired activation/barrier energy in meV obtained after logarithmic linearization from the slope of *Equation 6*:

$$\text{Equation 6} \quad \frac{Ea}{k} \times \frac{1}{T} - A = \ln \left\{ \frac{N(0)}{[N(T) - N(0)]} \right\}$$

$$a \times x + B = y$$

‘A’ is a linear constant B containing $\ln(-\beta \times t_0)$, and N_0 and N_T could be vaguely substituted with the normalized intensity of I_0 at T_{\min} and I_T at specific temperatures. However, since each step contains two (rapid time-resolved and longer temperature-resolved factors) processes, the $N_0=I_0$ equilibrium must be derived from the previously acquired step, so I_0 will now be marked as I_{T-1} . That holds true if the repeatability of PL is maintained. Given the results (3 repeatable cycles below 90K) in supplementary *Figure A21*, this could stand true to some degree, and calculations done in *Figure 54* could be made.

Original yet small, all acquired values were compared to BPW as reference (a) being approximately 7.9 meV. Other 5%, 2.5% & 0.5% Pr (and Yb) doped (b) SCWO, (c) CCWO cerium-tungstate derivatives with (d) BCW double perovskite follow a particular pattern suggesting the deeper the E_a (barrier by ± 5 eV) with lowering radius of A^{2+} ion: 19.0 (Ca), 15.9 (Sr) & 10.2 (Ba) meV respectively. The results seem rather consistent despite all the above approximations and simplifications (standard deviation is not bigger than 6÷10%). For pure BCW, in supplementary *Figure A22*, E_a is ~ 7.3 meV. This value is slightly smaller than for BPW but still within the error bar range. The answer for this might lie in measuring undoped cerate crystals hosting fewer defects.⁵³³ Of course, independent confirmation from outside sources is still required since the results could differ from batch to batch, given the pronounced instability of RE ions.

The $(E_a=kT)$ thermal effects and the expected placement of the shallow Pr 1S_0 states could be deduced from those slopes and these values. Not just for BPW but all Pr-doped materials alike. Evidently, they signify the thermal ionization energy ($a = E_a/k = T$ [K]) to eject electrons from the 1S_0 state to CB.^{489,491,534} Judging by LT experiments, those CTB ionized electrons can be retrapped on Pr states but also could fall on other traps, of which there could be many (meaning O_i/O^- or W^{5+} states reported previously by XPS & EPR in *sections 2.3 & 3.3*). If we heat the samples slightly from low temperatures, we will ionize the shallow traps or breach the barrier of the shallow RE levels. The

more heating-cooling-heating PL cycles we carry out (especially DPs in *Figure A21*), the more Pr^{3+} ions would be depleted. Some compounds seem more stable than the others and resist that process more, as seen in this and the previous section. Ca & Sr derivatives triumph in that matter much more than Ba cerates and praseodymates – that stems from the depth of ionization temperature values derived from the ‘a’ slope coefficient in *Figure 54*, indicating at which conditions those barriers could be naturally crossed. These temperatures equal 92, 184, 220 & 117K for BPW and Pr-doped SCWO, CCWO, and BPW, respectively. At room temperature, kT energy would equal roughly 25 meV, so we would have only deep traps working and a significantly high probability of self-ionization (by any given stimuli). As a result, we should have reloaded all electrons from Pr^{3+} and irreversibly changed the whole material composition. However, it is not precisely the case as can be seen on the next page - only after prolonged exposure Ba DPs do it, again, proving the existence of varying-deep traps stabilizing them at certain conditions until they get exposed. This phenomenon could be useful for dosimetry purposes^{535–540} but not in down-conversion.

One more remark. No wonder only ~30% of the grains emit so differently when shallowly placed Pr sublevels deviate subtly by several meV from the CB. Most of them are probably ionized. If more traps are localized near the bottom of the conduction band, this could severely impact the efficiency of the PL.^{541–545} With significant structural fluctuations taking place in the sample, simple thermal changes could change the whole PL intensity or diminish bands altogether by manipulating the E_g . Such a small photoelectric work function would also easily stimulate $\text{RE}^{3+/4+}$ charge transfer to or from the W or O_i/O^- taking a nonradiative route of energy dissipation.^{372,524–526,533}

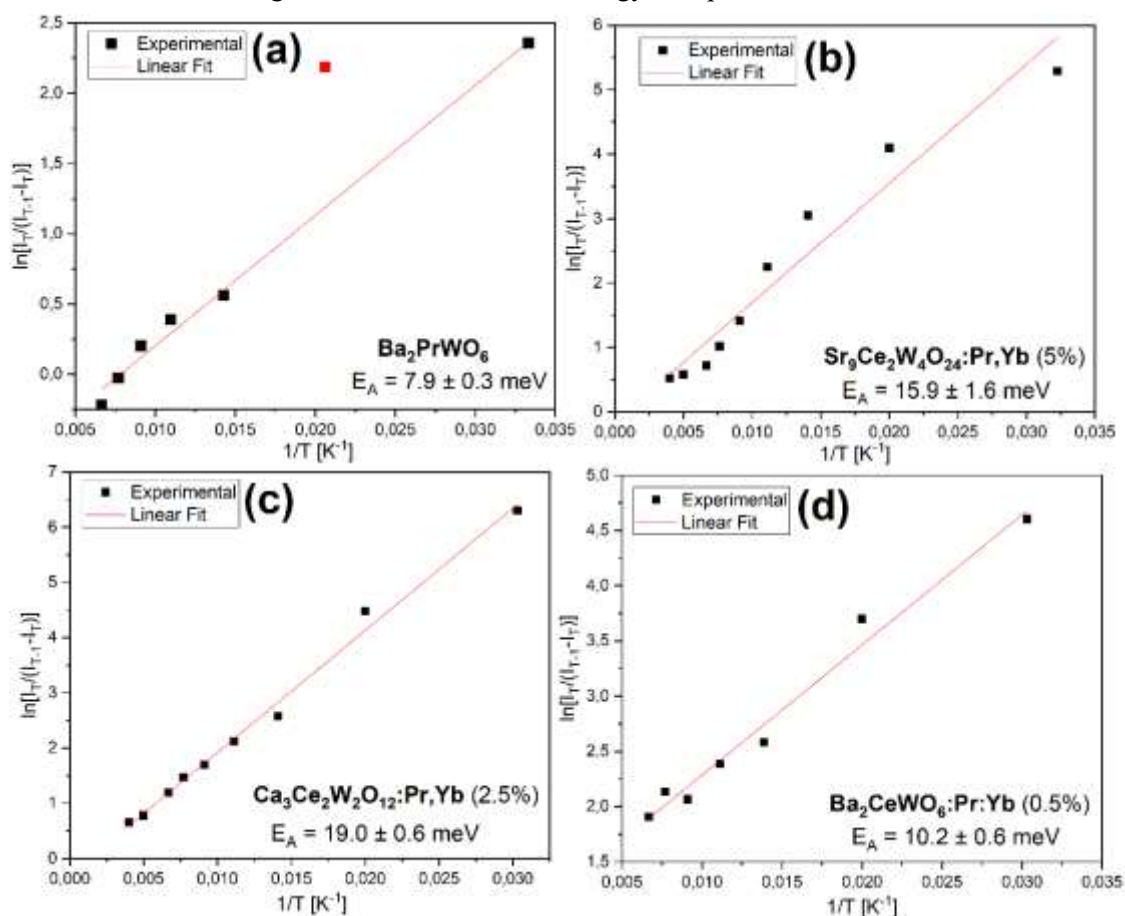


Figure 54 Activation energy estimates done for all investigated compounds based on the modified Arrhenius equation and measurements at LT temperatures in a pumped, helium-filled cryostat. Undoped barium tungstate DP reference - BPW is presented in figure (a), while other Pr,Yb-doped tungstates having similar E_A values are shown in graphs (c) & (d) regarding 5% SCWO & 2.5% CCWO derivatives. Graph (d) hosts data about 0.5% doped BCW double perovskite.

4.3 Oxygen evolution & lack of the energy transfer

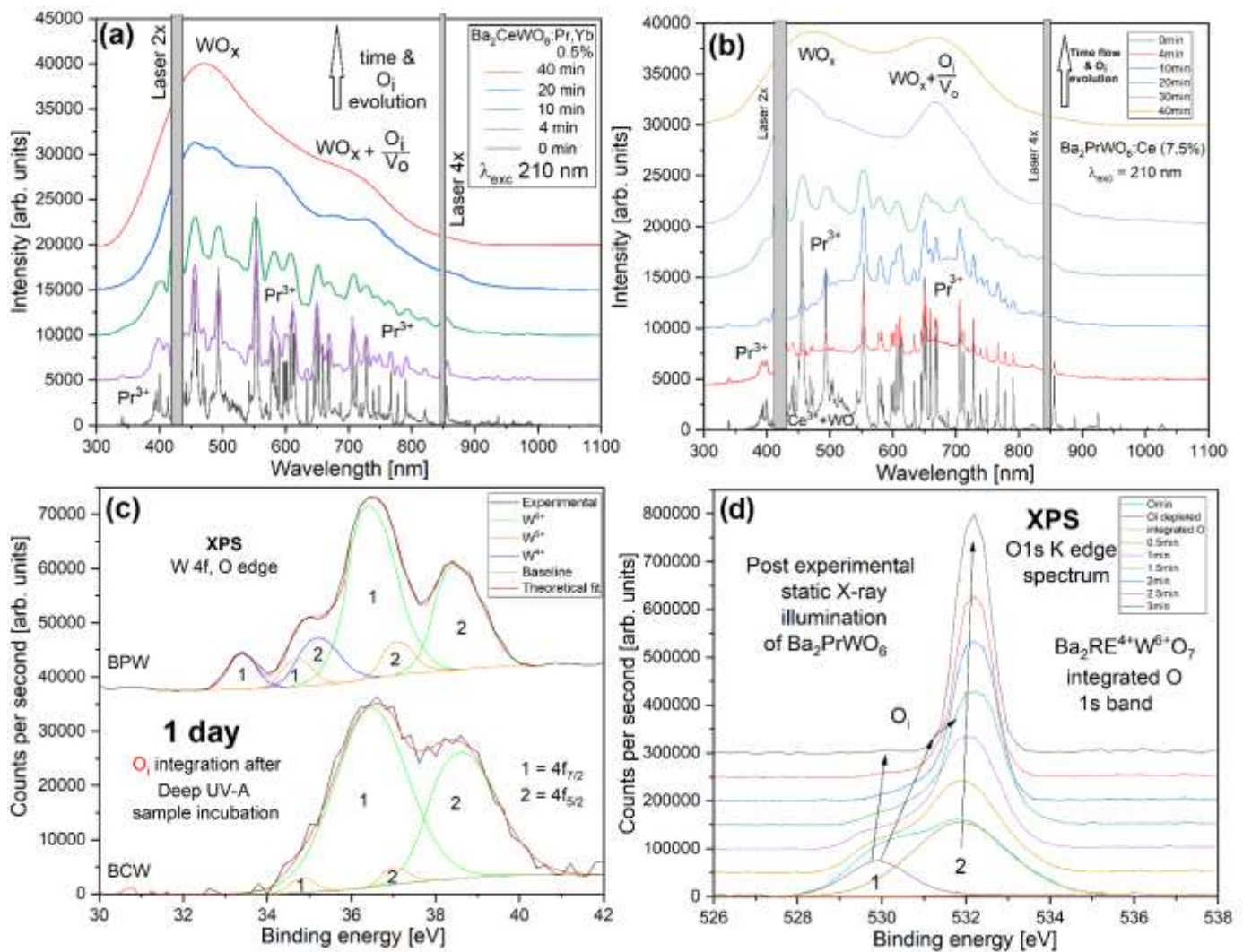


Figure 55 Time-resolved PL evolution of both doped barium double perovskites (a) BCW:Pr,Yb (0.5%), and (b) BPW:Ce (7.5%) in RT while exposed to intense $\lambda_{exc} = 210$ nm radiation in the air showcasing progressive transformation from sharp, Pr-specific to broad $BaWO_4$:Ce/Pr-like bands. Those changes originate from (c) prominent W oxidation featured on XPS supported instead by (d) O_I integration than RE-supplied charge transfer.

For a good reason, barium double perovskites' photoluminescence was investigated at LT only up to 150K. Not only was it already weak, but a strange phenomenon occurred while heating towards RT – all Pr^{3+} lines rapidly declined, broadened, and turned into two broad bands, completely not resembling anything from the initial spectra. The same phenomenon occurred when fresh samples were exposed to air for a few minutes at RT, as seen in *Figure 55a,b*. During HT measurements (up to 573K), those changes progressed even faster while being very weak (similar to HP results shown in *Figure A20*) during in-situ Raman investigation using a 325 nm laser. Above (till 873K), only WO_x bands were visible, if any at all, so the first assumption was probably just oxidation due to nonradiative processes such as rapid overheating or straightforward absorption. Samples' unfortunate coloring contributes to that factor, but that would not explain why those changes would also progress (though much, much slower) in oxygen-deprived conditions. An XPS check was made on a small amount of recovered material and, as it turned out, in *Figure 55c* composition of tungsten B-site ions has completely changed.

Given the exact nature of registered changes, W was initially to blame since the final form of the PL resembles the bands registered for BaWO₄ scintillators (possibly another instance of RE \leftrightarrow W charge transfer),^{2,202,521,546} but while checking oxygen's O1s spectrum a time-resolved rapid evolution of its content was discovered.

So, not only was the material oxidizing from the air while exposed to radiation spot on or above the energy of the CTB band,^{370,475,478,479,528,529} but also internally kept O_i/O⁻ species were involved in the whole procedure.^{384,387,530,531} Moreover, complementary XPS RE edges provided analogical evidence about the state of all given RE ions and another barium compound being treated the same way, as seen in supplementary *Figure A23*. This, unfortunately, meant that the materials were not just unstable (like many other known perovskites),⁵⁴⁷⁻⁵⁴⁹ but also compromised most of the initially made measurements due to exposition - they were more unstable than expected. Only those firstly-collected edges were valid. Each other shell needed to be measured on a different sample portion or rather 'spot' to avoid changing the initial composition of the material. That problematic outcome needed some deep tuning of the measuring equipment and much patience with frequent sample exchange to have proper, reliable conclusions; however, it was dealt with the utmost care, and all further presented spectra (i.e., in *section 2.3*) were attuned to that matter accordingly.

Initially, the proposed mechanism for mixed valency and all charge-transfer-related phenomena indeed involved both RE \leftrightarrow W interactions (as seen in the previous *chapter 4.2*). But now, O_i influence could not be entirely disregarded in general. Some circumstances, though, facilitate one of these two paths of matrix evolution when exposed to certain external stimuli. *Figure 56* proposes such conditions given all collected evidence until now – from the whole work considering both X-ray absorption and photoelectron edges with emission and excitation spectra. It is more evident that weak sources of light and pumped or inert atmospheres process O bands negligibly slow, but RE-related luminescence still changes based on photobleached Xe lamp spectra provided in *Figure 50*. That suggests that one could presumably stimulate the structure (and probably also the band gap width) only in a subtle way suggested by the inserted blue bracket. Nevertheless, when exposed to stronger radiation like UV or deeper (purple area), not fully reduced O²⁻ species (so O_i or dangling O⁻) oxidize everything in the vicinity - to not emitting RE⁴⁺ species while slowly progressing through 5/6+ tungstates to familiar species like insulating BaWO₄ scintillators (which emit primarily at HP when its E_g is lowered through compression and double phase transitions to a monoclinic phase via metastable fergusonite).

The process would seem irreversible (at least in the air or at HT), given the chemism and excessive oxygen supply. Both compounds (BCW & BPW) would heat up, rebond and uncouple their various BB'sites to REO_{2-x} + BaWO₄ species. Prominent HT color bleaching (to yellow) and several orangish grains acquired after exposure (*Figure A16*) would support that scenario with similar XRD patterns as those stated in *Figure 39* or *40* in *chapter 3.5*. However, since in air-deprived conditions, the original color of the samples seems to be maintained (at least under the microscope while PL properties still change), a plausible outcome of Ba₂(Ce/Pr)⁴⁺W⁶⁺O₇ system would also be achievable since the XRD diffractograms stay relatively the same. Just as Ivanov and Carmine suggested in their analogical work with lead-based Mn²⁺W⁶⁺ DPs,^{356,398,550} or many other scientists referring to similar monoclinic (P2₁/n) BaRE₂WO₇ or (C2/c) BaCr₂O₇ structure archetypes (which appear structurally sound to our BaPr₂WO₇ impurity from the chemical standpoint).^{191-193,430,551} In other cases introducing more oxygen, especially after a long exposure time, would upset the B-site charge balance with non-stoichiometric values and decouple the structure to cerates and tungstates as stated before if not create something resembling Ca₃Ce₂W₂O₁₂ ilmenite or Sr₉Ce₂W₄O₂₄ quaternary perovskite since these compounds are popular side products.^{97,98,114,401} These derivatives have more O, are more stable and are not prone to such prominent changes as barium DPs – reasonably because of the presence of fully

oxidized W^{6+} , which is more stable & would expand the band gap exponentially given the odds presented in available scientific literature.

Reconstructing these bonds (to the previous, mixed ion, less stable state) and oxidation to RE^{4+}/W^{6+} species is a contradictory process. It would need subtle heating in highly reductive atmospheres proportionally mixed with artificially added inert media.⁵⁵² An attempt was made to recover them, but the outcome was doubtful. The final batches had a completely different color (dark blue) than the initial ones (BCW – dark greyish blue, BPW – greyish green) and were problematic to analyze. They were wet (a lot of OH formed, not only on the surface), what required middle-range temperature treatment in pure inert gases (between 800-600°C just like in suggested W^{5+} acquiring methods in WO_{3-x}) to make them relatively dry - EPR indeed did not register any moisture.¹⁷³⁻¹⁷⁶ Solely Ce^{3+} with paramagnetic defects in BCW since Pr^{3+} in perpendicular fields is not detectable – Pr^{4+} was not seen in LT, probably due to antiferromagnetic phase transition at LT >90K or complete transition to Pr^{3+} , so the water was gone. However, XRD & Raman spectra showcased a mix of particularly everything in appendix *Figure A24*. There, many substances like highly distorted REO_{2-x} oxides (hosting mainly RE^{3+}); AWO_{3+x} tungstates and their derivatives - most prominent, e.g., $BaWO_4$; W/WC, WO_2 , and WO_3 oxides emerge joined together with main double perovskite phases. Optimizing the recovery process on such a batch was deemed futile, considering the lack of sufficient purism, but a point has been made. Reducing, breaking & correctly reassembling all $RE^{3/4+}$ bonds together with abundant defects (stemming from different -O- bridges like -OH & =O) to their prior RE/W charge-balanced states seems just unnecessarily time-consuming considering overall poor effects; however, not entirely impossible.

The conclusion is that only signs of anything recyclable would be present in the blue region of *Figure 56*. Especially if introduced to PL stimuli of reasonable strength at LT (for sure below 90K) in pumped atmospheres without any traces of air. In that regime, everything seemed repeatable and consistent regarding measurements – at least to some extent, as given by the number of cycles introduced in supplementary *Figure A21*. This temperature region of presumed recovery (and stability) could be somehow connected not only to the coexistence of coupled, wider band-gap $(Ce/Pr)^{3+} / W^{6+}$ ions but also to change towards lower symmetry SG (presumably to monoclinic $P2_1/n$ joined exclusively by antiferromagnetic transition just in BPW).^{71,72,397-400} This makes those materials bad and unstable phosphors with limited recovery capabilities. There are just too many conventional obstacles and competing phenomena happening simultaneously for them to serve as efficient downconverters. No wonder any of the theoretically predicted mechanisms in *Figure 57* will not work.^{3,306,490,491} Surely, they are interesting to investigate (i.e., UV→blue light quantum cutting) but do not possess any conventional means. They might have some potential in dosimetry or as a one-way, low-temperature radiation sensor; however, that last statement is yet to be verified by other extensive measurements concerning repeatability.⁵⁵³⁻⁵⁵⁵

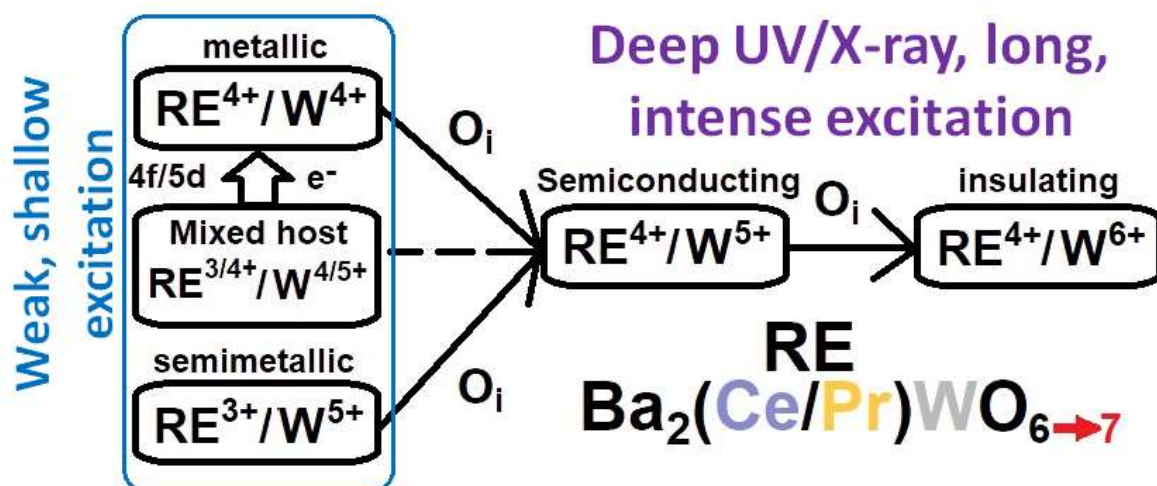


Figure 56 Exclusive mechanism proposal of aforementioned 4f/5d RE÷W charge transfer phenomena including a mixed path involving O_i interference followed by different strengths of Ba DPs illumination.

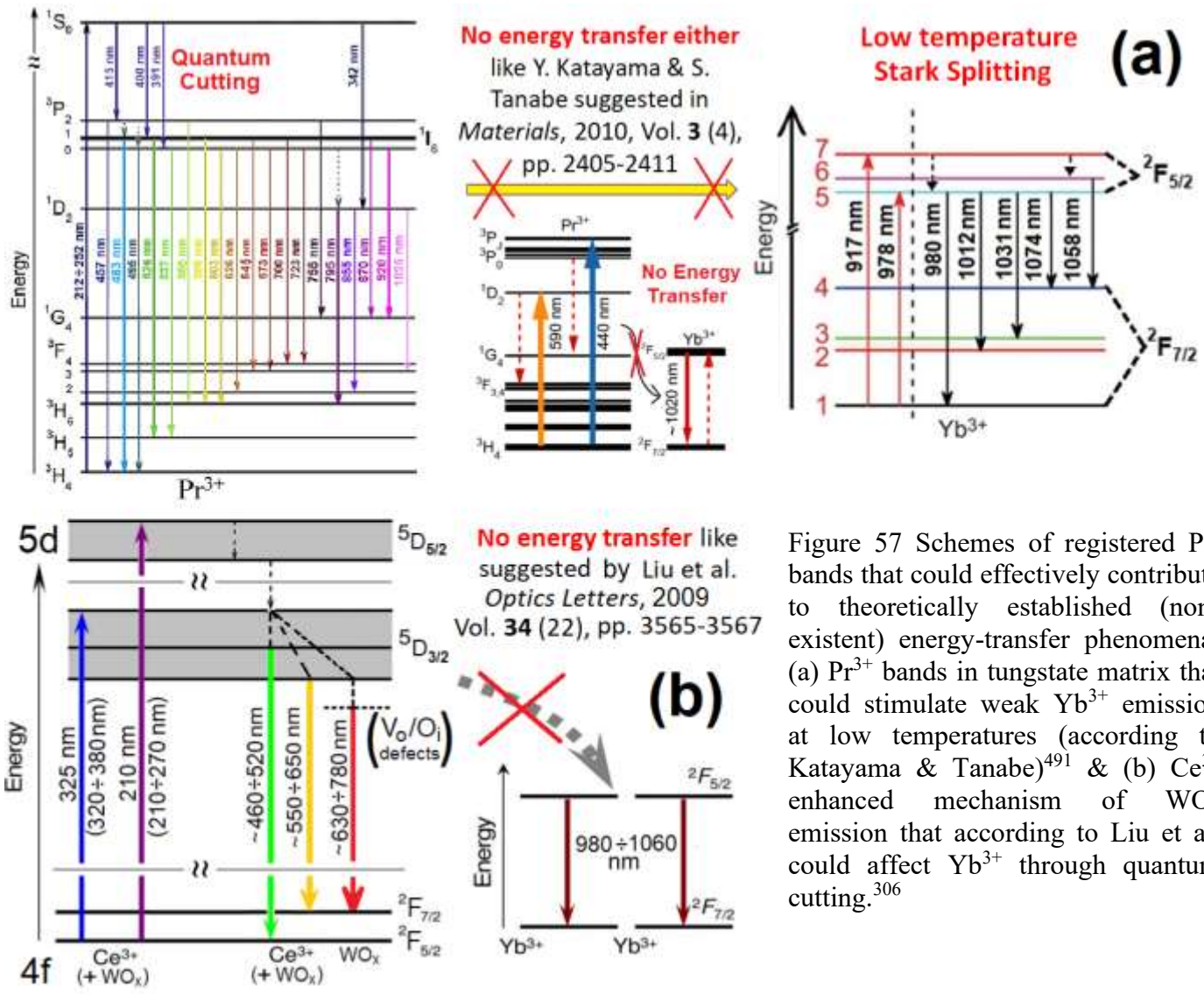


Figure 57 Schemes of registered PL bands that could effectively contribute to theoretically established (non-existent) energy-transfer phenomena: (a) Pr³⁺ bands in tungstate matrix that could stimulate weak Yb³⁺ emission at low temperatures (according to Katayama & Tanabe)⁴⁹¹ & (b) Ce³⁺ enhanced mechanism of WO_x emission that according to Liu et al. could affect Yb³⁺ through quantum cutting.³⁰⁶

4.4 Band gap estimations & absorption edges

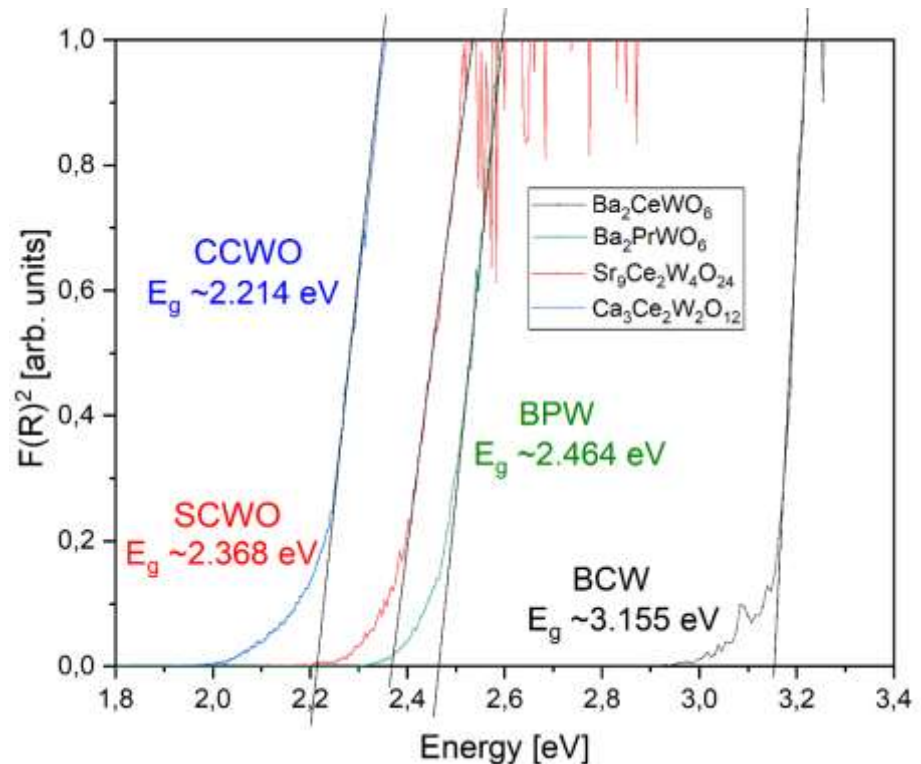


Figure 58 Kubelka-Munk equation incorporated into experimental reflectance spectra collected for all undoped samples at ambient conditions to provide practical insight about possible band gaps.

Experimental band gap (E_g) estimations were done by performing RT reflectance measurements on powders placed inside an integrating sphere in the widest possible range. Their results were processed through Kubelka-Munk equations^{556,557} for direct band gaps and are shown in *Figure 58* regarding all discovered undoped tungstates. As one can see, they progress quite smoothly regarding the shrinking of ionic radii. The gap will downsize substantially if one substitutes B-site, Ce central ion with Pr (while maintaining the charge state ratio). The same goes for alkali A^{2+} ions when Cerium is maintained inside – E_g closes going down Ba→Sr→Ca chain. These tendencies are to be expected, but other band gap ‘manipulation factors’ must be considered, like changes in temperature, stoichiometry, SG, or B-site charge ratios with partial site occupancy.^{517,558} Testament to this fact could be E_a values calculated for the doped materials in *Figure 54* where, beneath the CB, shallowly placed 1S_0 Pr state utilized mainly for the ionization process are ordered inversely (Ca→Sr→Ba) to declared tendency – the E_a value (by around +5 meV) rises as the E_g decreases (by +100 meV or more).

A more problematic issue is that those (experimentally & theoretically established) gaps are too small (by at least $2\times$) to be realistic even if they match & progress as expected when compared - *Figure 58* vs. *59*. While their width matches the values mentioned in common literature for such systems, they are rather improbable given that we observe all this luminescence described thoroughly in the previous 4.1-4.2 sections. Since the lowest PL excitation lines are around 4.92 eV for all Pr-containing samples (i.e., $^1S_0 \sim 252$ nm in BPW *Figure 45d*) and they are neither broadened nor Fano-shaped,⁵⁵⁹⁻⁵⁶² it means that these states are not in the conduction band but in the gap. The ground Pr^{3+} states are also in the gap (otherwise, there would be no overcharging), so E_g must be greater than the aforementioned 4.92 eV. We do not know where the ground state (3H_4) exactly lies (relative to the valence band,) but we can infer that it lies $4.92+0.0077$ eV below the bottom of the conduction band. Also, seeing bands from 1S_0 states gives the impression of gaps ranging to 5.5 eV or higher (not three or below), placing all compounds deep into the insulating range.^{398,550,563,564} Cerium hosts doped with Pr would also remotely adhere to that scenario by precedence. Not all samples gave out prominent PL, though – roughly 25-30% of grains emitted at all in the VIS range if doped with Pr. There are several hypotheses to answer that issue - one being based on theoretical Wannier/Tight-binding equations given available CIF files and discovered coexisting BB’-site cations balanced in (*section 2.3*) reported (4/4+ & 3/5+) charge state ratios.^{565,566} It would be most convenient to start with this concept and progress on using it as a base for further discussion. The divagations will be primarily based on barium double perovskites since they possess elementary, small unit cells without notable vacancies or partial substitution, making calculations reasonably processable. As seen in *Figure 59ab*, normal band gap estimations for $RE(Ce,Pr)^{4/3+}/W^{4+}$ systems place Fermi level (E_F) shallowly buried inside the conduction bands. These systems are metallic and would not emit at all (without any spontaneous defects or vacancies). That is why probably most of the samples’ grains do not give any response. Exposed to typical excitation wavelengths (as seen in *Figure A19*), they would be prone to fast, gap-closing charge transfer phenomena (which seems to take place while comparing *Figures 19* and *20* with *Figure 51ab*).^{363,366,369,371,523,529} However, when W starts to slowly oxidize through spontaneous charge transfer (initially to W^{5+}), i.e., during synthesis or illumination (weak CBT band stimulation), the E_F level would move out just slightly below the conduction band, making the whole matrix semi-metallic.⁴⁷⁴⁻⁴⁷⁹ That brings us to the previous discussion about E_a estimates. If such a small band gap with subtly placed sublevels is introduced to the system together with major structural fluctuations, simple thermal kT stimuli (between 90÷120K) could easily eject electrons from shallower states ionizing the RE dopants and constituents.^{372,524-526,533} Subtle electron work function (in the range of 8-11 meV) would influence $RE^{3+/4+}$ states, moving the charge carriers to nearby, i.e., W^{5+} traps and further complicating the electronic structure. That phenomenon could take a non-radiative turn, taking the energy's portion and proceeding further through a semi-metallic band gap towards the bottom of

the conduction band via nearby O_i sublevels. That cascade could also propagate through widespread vacancies all the way down to the valence band, hampering deeply desired chance for (sharp) radiative emission and, consequently, effective energy transfer.^{475,529,533,542,544,545}

Contrary to non-emitters (~70% of grains), the influence of CTB in these (presumably) semi-metallic multisite oxide structures could easily explain the small amount of initially discovered, not-quenched grains resisting nonradiative processes by exploiting brief RE↔W or O sub-band evolution moments. Changing RE charge states with W ions or O devouring excited electrons would (by chance) inadvertently & irreversibly expand the E_g by not a small amount - mainly due to O^{2-} injection from interjected O_i/O^- if not supplied from the air that binds with W (but, more about that mechanism in a moment).^{530,531,567-569} Different radiative responses could develop via crystallographic lattice evolution - uncontrolled expansion of local defects could separate many interjected states in the fluctuating E_g by creating an artificial ladder. Additional PL lines could emerge, split, or resurface from underneath the CB. At the same time, the gap is being modulated (especially inflated in shared or half-occupied atomic sites like in reported derivatives).^{498,503,518,519,527,558} All inclusions and erratic interatomic interaction between various ions (especially $Pr^{3+} \leftrightarrow W^{5+}/O_i, O^-$) in the lattice would unintentionally manipulate the crystal field universally changing clear, theoretically-established E_g into a more complex construct - rich in many sub-conductive (RE^{4+}, W^{6+}, O_i), intermediate (RE^{3+}, W^{5+}) and over-valence (W^{6+}, O^-, O^{2-}) states which would make all simply-modeled tight-binding estimations & PL-related spectra not trivial to interpret jointly. The final outcome of this evolution mechanism would be a semi-conducting or even insulating system.^{398,550,563,570-572} Then, E_g values at Γ , while still greatly underestimated, would at least match those received from previously done reflectance measurements. A visual description of the $I2/m$ Brillouin's high point symmetry zone can be seen in the appendix *Figure A25a*.

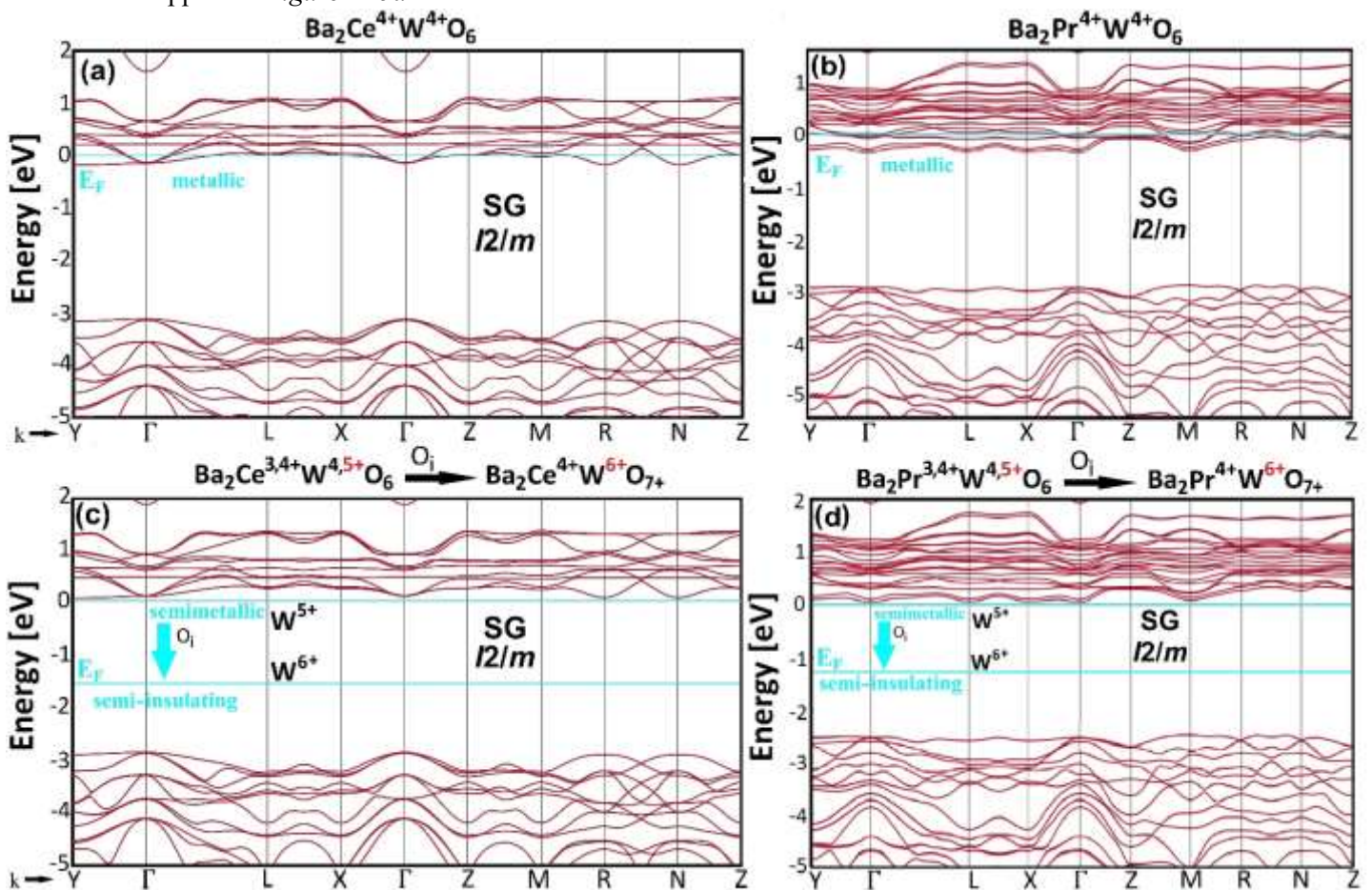


Figure 59 Theoretical band gap estimates provided by VASP software using Wannier/Tight-binding equations for both barium double perovskites (BCW & BPW) hosting primarily-selected, simplistic $I2/m$ unit cells with: (a,b) only pure $(RE/W)^{4+}$ ion couples, and (c,d) those being imbalanced by W^{5+} ions further influenced by the O_i presence towards W^{6+} .

The O_i interference or O^- from internal V_o complexes is the most plausible scenario. Those species mediate in the E_g spreading process from ground $2p$ up to near-CB Pr states by involuntarily oxidizing the sample's ions (ionizing not intentionally applying various stimuli/changing conditions), i.e., tungsten to the $6+$ state and moving the E_F level as shown in *Figure 59c,d*. That could exponentially distance conduction from the valence band (VB) towards the deeper UV region.^{527,530,531,568,569} Usually, interstitial forms of O like O-O dimers in a neutral state break upon capturing a traveling electron (thus quenching PL). When bond-breaking relaxation occurs (to dangling O^- or fully integrated O^{2-}), the antibonding defect level lowers itself from the conduction band (CB) towards valence states.^{531,569} The implication suggests that those O species act as temporary electron traps for the E_F close to the CB; thus, the excessive amount of interstitial atoms could positively shift the threshold of E_g towards values, as presented above in Wannier-shaped diagrams. Here, under weaker light illumination or negative stress stimuli (stretching), the Fermi level could lower itself toward the edge of the VB where O-O dimers could feed upon capturing free hole carriers with neglectable energy barriers (a theoretically recoverable process), but that might not be the case (as tested, the process is rather irreversible considering most grains are non-PL-active / semi-metallic in origin).^{531,568,569} Shallowly placed RE states might be responsible, especially considering the presence of RE^{3+} charge states even if in minuscule amount after exposure – when O-O bands would move back up (looping) after feasting upon the holes crossing Pr states would drain them up quickly.⁵⁶⁹ This phenomenon can be temporarily preserved, given that it originates from defective REO_{2-x} substrates (mainly referring to Pr). PL from them is also visible in our doped materials & is partially recoverable for a short while in progressing through a few cycles - *Figure 50*. However, one must maintain inert, protective atmospheres in LT (stored without an additional income of outside O to be sure of what one is dealing with). So probably RE partially hijacked this mechanism at the very start, given their shallow-placed states exist nearby with a reasonably small gap till they are also 'fed up' with oxygen-processes holes ^{363,479,531,568,573,574} - competing CB to VB processes of O^- , O^{2-} , RE^{3+} , $W^{4,5+} \leftarrow e^-, h^+ \rightarrow O-O$, O^- , RE^{4+} , W^{6+} would suggest some discrete balance without and distinct priority since all these forms coexist with each other; however, less detectable O^-/O_i forms seem to dominate the process at the end. This last statement would explain what happens during the HT synthesis (disproportionation between $3+/4+$ species, $2+/3+$ in Yb due to kT influence, and O_i/O^- coexistence). However, eventually everything would be forced to oxidize and turn into RE^{4+}/W^{6+} states by changing the stoichiometry towards $Ba_2RE^{4+}W^{6+}O_7$ systems and adding $+1O^{2-}$ to the O_6 subscript (like in air - *Figure 55*) during strong laser excitation.⁵⁵⁰ In this scenario, the E_g could grow even bigger and more insulating than predicted. E_F placed in the middle of such vast E_g would stop all these aforementioned phenomena - at last, some long-term current-voltage stability would be maintained between RE/W B-site ions. The whole process would resemble HT annealing, given that all the additional oxygen would integrate itself with the crystallographical lattice (O^{2-}) and forge/retrofit some sort of effective band gap in temperature-induced multi-oxide systems.^{527,530} That total outcome was initially unexpected and would require excessive band gap engineering.^{527,530} Divulging upon this subject could create a new, stand-alone topic for a separate dissertation and will not be expanded upon here. This field seems to be a rather hopeless expansion at that point while tackling basic BB'-site instability issues (this is a fundamental study for newly discovered materials), and energy transfer will always be affected by it ^{370,371,479,528,573} considering the chosen concentrations of remaining, scattered emitters (Ce and Pr)³⁺ will always diminish with respect to the planned amount of receivers ($Yb^{2+} \rightarrow Yb^{3+}$ - refer to *Figure 25* for details) but, at least PL quenching could be partially taken care of here given the limited supply of oxidizing agents during vacuum-pumped or $N_2/Ar/He$ -filled synthesis/experiments.

The other issue to address probable large E_g (stationary, not only based on acute broadening) could be the presence of intrusions and the influence from the nonuniform distribution of RE dopants or, rather, their $3+$ ions within all the investigated matrices given randomly scattered PL-active grains and

possible minor phases visible on XRD diffraction patterns (summarized briefly in *Table 6*). Since we are investigating (novel) powders, not crystals, having RE ions as their constituents scattered in different charge ratios throughout doped matrices is a widespread problem - researchers usually encounter such issues during first stages of synthesis and either try to develop a perfect purity/RE distribution ratio by improving diffusion rates, prolong the sintering time/cycles or switch to highly-ordered crystals if possible.^{567,575,576} In our case, the syntheses needed to stop at some point, considering every cycle introduced slightly more oxygen defects to the isolated system and too high diffusion rates facilitated segregation of AWO₄ systems from REO₂ oxides (as *section 1.2* states), thus resignation from Pechini method and backtracking to SSR. The percentage of the luminous grains could showcase such struggles, but their constant spatial and behavioral repeatability (at least visually for the measurements performed under the microscope) would suggest some effort towards mitigating those problems. Those grains seem not to be just some random entities (physically unidentified phases) scattered around. They are too abundant (easy to find ~ 30 volumetric %) and consistent in their performance to be undetected.

That brings us to the last, most plausible solution of the alternating E_g problem, the other, most probable *R-3* phase. As discussed at the beginning of paragraph 2.2, both barium RE tungstates could possess one of the two indistinguishable SG - just basing on the powder XRD and Raman spectroscopy was insufficient (- access to NPD would be required). By order of elimination using spectral techniques in non-ambient conditions, the trial concluded that we most likely deal with *I2/m* (*C2/m*) SG system, and the discussion proceeded with that conclusion further on. However, the ‘1-0’ disqualification procedure might not be a valid approach. Both materials could hang in a discrete balance within a specific temperature given that the alternate *R-3* phase is the less stable (as the literature suggests, it persists at around 353K or even more).^{29,71,72,357,359,360} That theoretically assumed coexistence might be the actual culprit for not distinguishing any of those two primary components of the batch. That rhombohedral phase (presumably being in these conditions in the minority given the low number of emitting grains) might be responsible for band gap differences given the changes in the Brillouin zone (depicted in *Figure A25b*) with respectable RE/W BB'-site ions. Adequately, similarly-tailored Tight-binding projections of E_g were made and presented in supplementary *Figure A26*. They apparently still disprove our suspicions by at least half the value, although slightly bigger than in *I2/m* (especially in the case of BPW, whose gap still rose by ~0,5 eV). Even if the calculations underestimate the gap (they usually do), the differences are just too big to be neglected and left without any answers. Sometimes E_g width progresses in monoclinic > rhombohedral ≥ cubic SG order while other (p, T) conditions are maintained but not as much.^{518,519,577-580} The reason for such PL must lie elsewhere, presumably in disproportionate RE³⁺/W⁶⁺ intrusions. At LT (below 150K), both barium tungstate systems also undergo a phase transition to most probably monoclinic *P2₁/n* SG,^{359,361} which usually widens the band gaps even more,^{517,576-578} but BPW additionally goes through antiferromagnetic transformation slightly below 120K (stretching down to 90K) which might contribute to the effect in many unpredictable ways – probably would make it an antiferromagnetic insulator after ultimately moving the charge to Pr³⁺/W⁶⁺ states.^{399,581}

All aforementioned structural fluctuations play a pivotal role in meddling with E_g and PL-related phenomena: mixed or changing ionic radii after stimulating Pr↔W, O sublevels by CTB, mobility-related defects, vacancies from redistributed oxygen, shared octahedral or half-occupied dodecahedral sites were reported here as well as in other less complicated structures having similar issues. This alone could explain all phenomena mentioned above, E_g evolution and potential ‘fluidity’ or rather instability of excitation and charge-related states.⁵²⁴⁻⁵²⁶

Conclusions

- Not all four newly-discovered compounds did turn into desired double perovskites due to apparent tilting from shrinking of A^{2+} polyhedral sites as the worsening Goldschmidt factor predicted. Spacious barium accommodated truly desired perovskite matrices, although slightly tilted from the perfect cuboid cell. Strontium created an intermediate structure called quaternary perovskite riddled with many defects - partially occupied $A^{2+}/RE^{3/4+}$ sites and half-empty vacancies. Calcium made densely-packed ilmenite, which in comparison to Sr, expanded two-fold while maintaining partially-substituted sites but did not possess any noticeable A-site vacancies.
- Barium cerium tungstate (BCW = Ba_2CeWO_6) and barium praseodymium tungstate (BPW = Ba_2PrWO_6) turned out to be compact but slightly-tilted, cuboid $I2/m$ SG, although (through powder XRD) the alternative $R-3$ SG can not be entirely discarded. The same issue concerns huge unit-cell of calcium cerium tungstate (CCWO = $Ca_3Ce_2W_2O_{12}$), which is probably centrosymmetric $R-3c$ than (noncentrosymmetric) $R3c$ – neutron powder diffraction studies are required at reputable nuclear reactor sites to be sure of that matter. The least-concerning compound (regarding SG assignment) is strontium cerium tungstate (SCWO = $Sr_9Ce_2W_4O_{24}$) which, although belonging to a hotly-debated group of minerals, has a pretty accurate fit – a tetragonal $I4_1/a$ SG without any close-call alternatives. (The other, much worse, is $Fm-3m$)
- All compounds present themselves with some internal B-site cation duality balance making them more stable (Sr & Ca derivatives) or not (Ba double perovskites). Derivatives such as SCWO and CCWO seem to last longer at ambient conditions, given that no external NUV or deeper-energetic stimuli are applied, probably because of their $Ce^{4/3+}$ vs. $W^{5/6+}$ 9+ charge balance. Their W^{6+} -based SGs seem unfazed since the ions' highest oxidation state does not have an alternative to go any higher. In SCWO & BPW, the general B(RE/W)-site ratios seem to be upset and do not fit as well as in their BCW & CCWO counterparts. Barium compounds seem to behave differently. Not only do they bestow internal 8+ B-site balance – $RE^{3/4+}$ ions are artificially 'paired' most likely with $W^{5/4+}$ counterparts (not closely but distanced, presumably scattered - balanced in the whole volume of the sample). Their response to the external radiation is also different – (at first) weak photosensitive stimuli seem to change them towards other non-radiative species like RE^{4+}/W^{4+} phase irreversibly. This might be because of their shallower ionization levels ($E_a \sim 7-11$ meV) compared to cerate derivatives, which reach 15-20 meV. That correlation would definitely reflect in some of their kT thermal properties vs. PL dependencies at LT.
- This situation is, however, brief. If exposed longer to more intense illumination (highly energetic +250 nm source or through overheating), this transformation proceeds much further. Both barium compounds react irreversibly with interstitial oxygen (O_i) stored in nearby vacancies (V_o) or complexes and oxidize by attaching 1 (dangling O⁻) more oxygen (O^{2-}) to the stated initially stoichiometric formulae. This phenomenon, alongside some shallowly present 4f/5d RE charge-transfer states, could contribute to overall photobleaching phenomena transforming all constituents & dopants from RE^{3+} species to photo-inactive RE^{4+} ions rendering these already-weak phosphors as bad emitters.
- At high temperatures, when overheated (intentionally) in the air-filled furnace or by a laser, a similar process (to this stated in the previous bullet point) would be responsible for the same oxidation effect. It concerns, however, throughout decomposition of the structures into separate products like RE oxides (substrates) and AWO_4 tungstates. When overheated above a specific limit, as stated by DSC & TG, specific $RE^{(3+)}$ or $W^{(4/5+)}$ species will oxidize, changing the band gap barrier by utilizing the materials' current structure in the ambient environment. This was confirmed while comparing the behavior against a sealed or protective inert gas environment media. In those gases, the process was still present, however severely limited. Unintentional O_i/O^- storage might be the culprit behind all of this. It is difficult, if not impossible, to monitor & control the content of all O species even during the synthesis since oxygen substrates are also the actual donors. However, some signs of unit cells'

amorphization or (octahedral) tilting are quite prominent. That statement alone might create a firm basis for all material's time-resolved sensitivity to photoionization.

- Polymorphism investigation in high pressure concluded that CCWO and SCWO derivatives could encounter some type of (theoretically-supported)^{71,72} reversible phase transition above 11 and 15 GPa, respectively. Most likely to $C2/c$ SG of lower, monoclinic symmetry (possible for both CCWO & SCWO) than $C2/m$ (exclusive solely for SCWO). However, this statement needs to be verified by HP XRD since the change was observed only on the Raman spectrometer. It all seems to bear second-order transition characteristics - stretched out, continuous yet subtle phonon energy fluctuations, although with noticeable yet slight hysteresis during decompression (first order?). Barium double perovskites, contrary, stay the same up to 20 GPa.
- DFT calculations provided by *Howard et al.* also support low-temperature research by reporting another set of second-order phase transitions - this time related to barium DPs. In them, the $I2/m$ phase probably shrinks and tilts more towards the monoclinic $P2_1/n$ SG. However, this could also be a reversible transition from coexisting $R-3 \div I2/m$ phases towards a more RT-stable (latter = $I2/m$) option.^{71,72} On the other hand, that last presumption would not support spectral group theory calculations with an additional number of Raman-active modes slowly appearing just below 200K (\leftrightarrow 100K). Just as before, further investigation through LT powder XRD or NPD would be required since this might be just some sort of decomposition (not likely since irreversible) or rather $RE^{4+} \leftrightarrow RE^{3+}$ & $W^{4/5+} \leftrightarrow W^{6+}$ disproportion including O_i ionization as further XAS & EPR measurements will show.
- Low-temperature XAS was done to check whether or not changing conditions & probable phase transitions (for BPW & BCW) will influence the ionic, B-site structures of our materials during the unit cell's shrinkage. The charge-transfer phenomenon would surely upset the discrete balance of B-site ions, influencing future down-converters' PL & energy-transfer properties. Low-eV radiation (during PL experiments) does not seem to influence barium RE-tungstates' unpredictable behavior. LT phase transition that transpires in them seems to have something more lurking beneath their purely structural nature. In LT, the SDD acquisition mode of XAS shows signals from derivatives (SCWO & CCWO) with similar ratios regarding Pr dopants, although relatively weak and noisy. The error bars are huge enough to set all values in common grounds to RT with no proper conclusions derived. However, things are different regarding barium samples & their constituents. RE^{4+} seems to change entirely to RE^{3+} - that is probably the reason why EPR does not see any Pr^{4+} . Moreover, TEY results show nothing at all (or only a few residual peaks from previously main features), promptly hinting at a massive drop in the conductivity of pellets towards more insulating values - as confirmed later by EPR's microwaves. Such behavior would imply temporal freezing of the carriers below a specific temperature (presumably supported by obstructed e^- propagation through abundant-in-V_o/defects matrix - mainly in derivatives) or charge transfer from W to RE in such a way that it would expand the band gap - that last scenario seems reasonable for barium samples). The last process would expand the bandgap (i.e., lone $W^{4/5+} + O_i \rightarrow W^{6+}$ or $2 RE^{4+} + W^{4+} \rightarrow 2 RE^{3+} + W^{6+}$; $RE^{4+} + W^{5+} \rightarrow 2 RE^{3+} + W^{6+}$) & probably change stoichiometry for both barium DPs due to prominent tilt from changing ionic radii. O_i/O^- is probably involved too, but not much since EPR does see considerable amounts of it in BCW. The interference might be a result of changing magnetic properties.
- While the postulated conductivity changes for all the materials, the phase transition seems to proceed reversibly & only within BCW & BPW - somewhere below 150K (according to Raman spectra, the point is not sharp). There is also a hint about another antiferromagnetic change - a phase that would appear around 120-90K solely for BPW, which clearly has more O_i content than the other materials and does not possess any Pr^{4+} & W^{5+} ions on LT EPR (20 \div 3K). Below (typical, second-order) transition temperature, PL Pr^{3+} peaks also seem to recover quite nicely during (3) heating-cooling-heating cycles in the cryostat, so radiative processes there seem to be at least partially recoverable.

- All gradual changes in PL seem to be induced by time-resolved exposure to high-energy radiation or phase transitions mentioned above (structural or magnetic still awaiting further confirmation by more elaborate techniques). Both changes could carry a bit different stoichiometric outcome – simple cryogenic cooling seems to spare RE³⁺ ions but sacrifices all RE⁴⁺ & most importantly, W^{4/5+} species towards W⁶⁺. The assumed outcome would influence the structure by changing ionic radii or B-site balance, introducing a total 8→9+ charge disproportion compensated by reintegrated O_i→O²⁻. Barium samples would be affected by the creation of Ba_{2+x}RE³⁺□W⁶⁺O*_{6+O_i} complexes where ‘□’ is a ‘defect’ or rather a vacancy. Sever outcome would presumably be Ba₃RE³⁺₂W⁶⁺₂O₁₂, similar to Ca₃Ce₂W₂O₁₂. Both (BPW & BCW) LT transitions, however, seem to be reversible (as proven by Raman spectra), so the Ba_{2+x}Ce³⁺□W⁶⁺O*_{6+O_i} structure would be a more adequate approximation considering that Ba₃RE³⁺₂W⁶⁺₂O₁₂ would completely (irreversibly) change the internal packing order. The band gap modeling, absorption, and widespread RT & LT PL measurements commenced shortly after since radiative mechanisms of shallowly placed RE^{3/4+} charge-transfer states (under the 5d conduction band) were at risk. Their response would be pivotal in the energy transfer phenomena we dearly searched for based on introduced RE³⁺ ions.
- Compounds naturally containing W⁶⁺ indeed fit experimentally estimated values of absorption spectra & theoretically established tight-binding E_g - SCWO & CCWO are presumably semiconducting. However, band gap modeling in VASP software showed that BCW & BPW should be semi-metallic with ions such as W^{4/5+} (respectively E_F shallowly buried or just below the CB). That mainly explains no PL from ~70% of the grains. Both materials transform during intense illumination (NUV or deeper) due to charge transfer, defects, or O_i photoionization. Time-resolved acquisition with a weaker Xe lamp gives room to observe a slow, progressive evolution (allowing only for a short, brief exchange of 1 e⁻ from RE³⁺ to W⁵⁺), making PL spectra completely flat. Later stages of the experiment, involving potent laser power, instant feature integration of O_i/O⁻ into the matrix (Ba₂REWO₆₊₁), promptly changing the W^{4/5+} mix towards W^{5/6+} and eventually, W⁶⁺ expanding the band gap & lowering the Fermi level from almost the bottom of the conduction (semi-metallic) to nearly half the way towards the valence band making them; as a result, deeply-insulating. All aforementioned discussion refers to theoretical values at Γ fitting and the chemical mechanism depicted in *Figure 56*. RE would also be fully oxidized at the end (to 4+ species), creating Ba₂RE⁴⁺W⁶⁺O₇ by that point – they seem to respond first to the stimuli towards W, but its sole oxidation (to 6+ charge, whether from RE or O⁻ electrons) is what actually steers the band gap and whole observable PL effects.
- Based on modified Arrhenius law (*Equation 6*), activation energies (E_a) of all Pr-doped compounds and pure barium DP matrices were calculated. Thanks to that procedure, the depth of shallowly placed RE sublevels (8-19 meV) and their common interactions with the conduction band could be addressed (E_a rises with Ba→Sr→Ca substitution contrary to the E_g). Of course, that does not change the fact that both tight-binding band gap approximations made in VASP software and reflectance measurements grossly underestimated the width of the E_g - by at least ½ of its value. Given the emission seen from Pr, the band gap is undoubtedly bigger, and part of it corresponds to radiative transitions from ¹S₀ states placed somewhere close by the CB (E_a values would correspond to that). Since the influence of possibly coexisting alternate (*R-3*) SG was rejected (considering it to be an intrusion having a more significant E_g), there could be only two other sets of solutions to answer this ‘bigger-than-usual’ band gap dilemma. First are independent inclusions since most (~70%) non-emitting grains presumably have almost-closed, direct band gaps. The other is a nonuniform distribution of dopants/RE ions. Both cases could explain such an erratic response. The origin and effects of the first scenario are quite obvious. Structural fluctuations in non-trivial LT/HT conditions would influence the second outcome during low/high laser-powered experiments. The electronic structure of DPs could be manipulated during synthesis by just slight changes in delicate p(T)/media conditions influencing final B-site ratios. The interchanging RE/W charge states with varying ionic radii (especially towards W⁶⁺), artificially injected O²⁻ from O_i/O⁻ intrusions mangled within the lattice, and structural or possible LT antiferromagnetic

transitions (the last one very typical for oxides) would also unpredictably modulate the band gap by blurring its width with many intermittent levels – that unintentional meddling with the band gap seems like engineering, but in our case, it could cause electrons to weakly emit or slide non-radiatively back to their valence bands what is not appreciated.

- All the main Pr 1S_0 and 3P bands (from dopants & constituents) with their weaker Ce+WO_x contribution at the base of the PL were accounted for. There was only one particular ion missing. The Yb³⁺ emission was searched to find any signs of energy transfer phenomena. However, unfortunately, only direct illumination (typical for NIR Yb excitation) gave some weak to almost no response despite interesting Pr emission in the form of quantum cutting just beneath the blue region (- we were expecting red). Yb³⁺ response was later measured at LT, hosting an interesting Stark splitting effect – some of the Yb lines that emerged might be useful for future applications. However, one, denoted as 6>4 transition, required special attention regarding selection rules since its odd appearance depends on 917, 978 nm laser illumination.
- Without any direct energy transfer/UV excitation from Pr³⁺ sensitizers or WO_x mediators, the whole energy transfer ordeal seems rather futile. Derivatives were slightly better in exposing Yb, though, maybe because they are more stable than barium DPs regarding PL. As it turned out later on XAS, a part of Yb³⁺ was reduced during the synthesis - the expected/applied dopant concentration dropped towards Yb²⁺ probably while applying Ar:H₂ atmosphere during SSR or due to reversed charge-transfer phenomena (a charge transfer side-effect victim). O_i interference could also meddle in that matter; however, in terms of Yb, it helps by oxidizing their ions fully back to the 3+ state. O_i/RE interference seems not selective and mainly worsens the PL effects by changing Ce/Pr³⁺ to 4+. In the case of Yb, the 3+ charge is the highest PL-active oxidation state, so the whole system benefits from this process.
- Materials (primarily referring to Ba DPs since derivatives are slightly more stable) are weak phosphors and are only useful (PL-wise) when RE³⁺ ions exist within a limited time frame of coherent-light exposure –at least from HT, RT, down to 150K where changes are rapid and irreversible. That is probably due to the proximity of several states & thermal kT effect dictated by E_a - 10-15 meV gaps of states placed maybe not coincidentally between 115-170K where transitions mentioned above occur). Below 90K, there is a chance to recover them but not entirely - to the same value as before but up to approximately three times). The general implication is that they could be, unfortunately, used only as a one-time, irreversible sensor, best without any air around & when cooled down under respectable and coherent NUV excitation light.
- As it can be necessary for any other purpose, it is worth stating that, during charge transfer (RE-W/O_i RT/HT evolution involving NUV lasers), most RE modes broaden, red-shift, and diminish, giving place to much stronger W⁶⁺O₄ bands - typical for Ca/Sr/BaWO₄ scintillators. That means the (irreversible but surely noticeable) changes disrupt the balance of differently charged RE/W ions. A brief moment of exposure (around 20-40 min in RT, NUV light shrinking to mere seconds in LT, deep X-rays) could be a trigger signal of changing B-site ratios outside their delicate equilibria, swiftly diminishing white light from RE→WO_x energy transfer. Any external, highly energetic stimuli in RT or HT could be easily distinguished if exposed to UV by rapidly growing yellowish-green emission. In LT, the cold-white light would partially decay – just like under the weak Xe lamp.
- Judging by LT experiments on Pr-doped samples, ionized electrons can be retrapped on Pr³⁺. They can also land on deeper electron traps, which could be many (like O_i/O⁻, W⁵⁺ reported by XPS & EPR). Slightly heated-up samples at low temperatures can ionize the shallow traps close to praseodymium levels and jump to the conduction band. Some electrons may recombine, and others may go toward other shallow traps. However, at RT room temperature, kT would be ~25 meV, so only deep traps would remain with increased ionization probability towards emission. As a result, after prolonged exposure, we should have reloaded all electrons from Pr³⁺ towards the deeper potential well. This phenomenon could be useful in dosimetry, especially if forged into more efficient crystals.

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Appendix

Warszawa, 08.06.2023

(miejsowość, data)

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OŚWIADCZENIE

Autora rozprawy doktorskiej

Jako autor rozprawy doktorskiej pod tytułem:

[PL]: *Synteza i właściwości fizykochemiczne nowych wolframianowych podwójnych perowskitów oraz ich pochodnych domieszkowanych wybranymi jonami ziem rzadkich.*

[ENG]: *Synthesis and physicochemical properties of novel, tungsten-oxide double perovskites and their derivatives doped with selected rare-earth ions.*

przygotowanej pod kierunkiem Prof. Dr hab. Andrzeja Suchockiego,

przedłożonej Radzie Naukowej Instytutu Fizyki Polskiej Akademii Nauk oświadczam, że:

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- 2) rozprawa nie była wcześniej przedmiotem procedur związanych z uzyskaniem stopnia naukowego w innym przewodzie doktorskim,
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Damian Włodarczyk
(podpis)

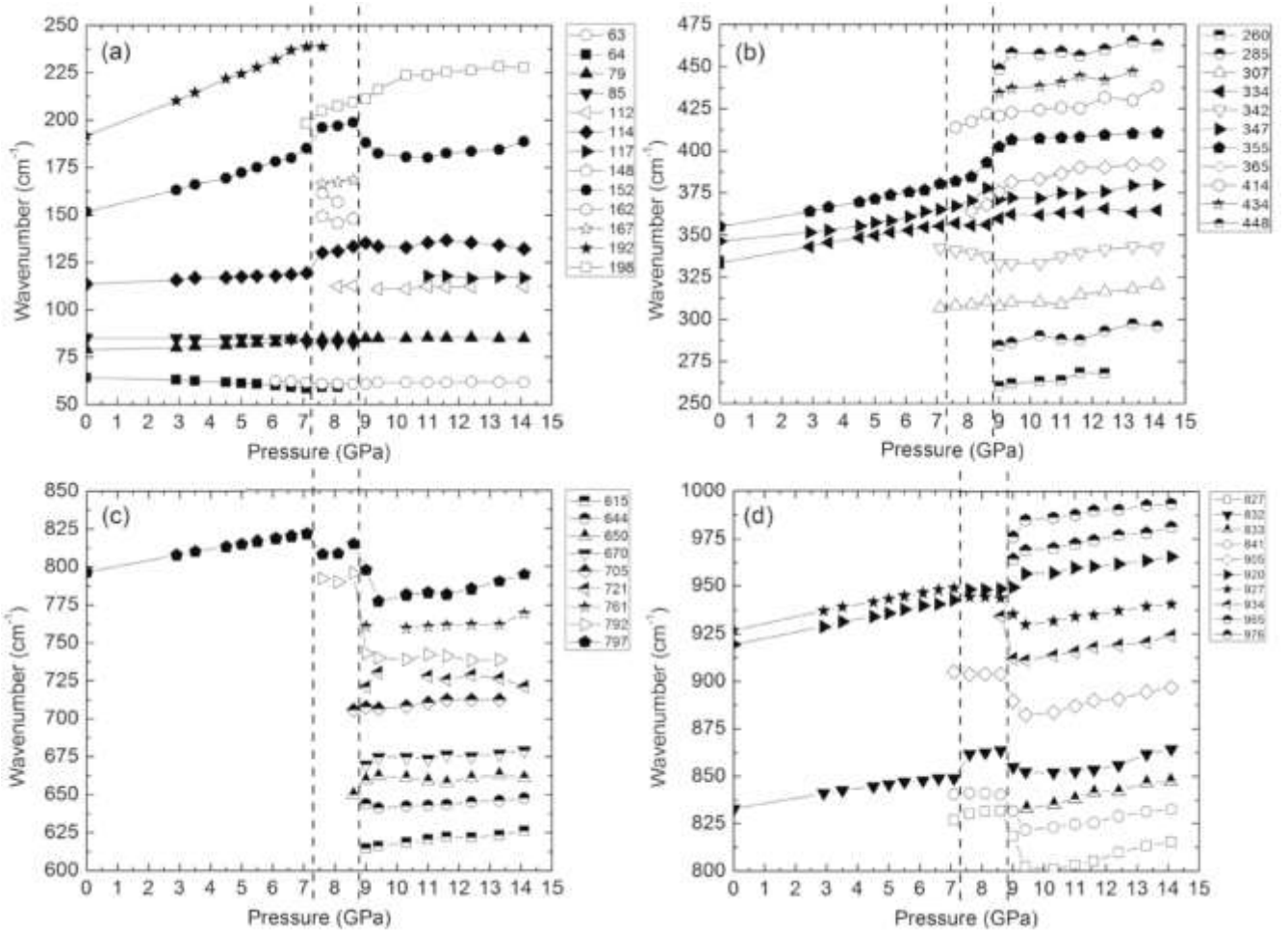
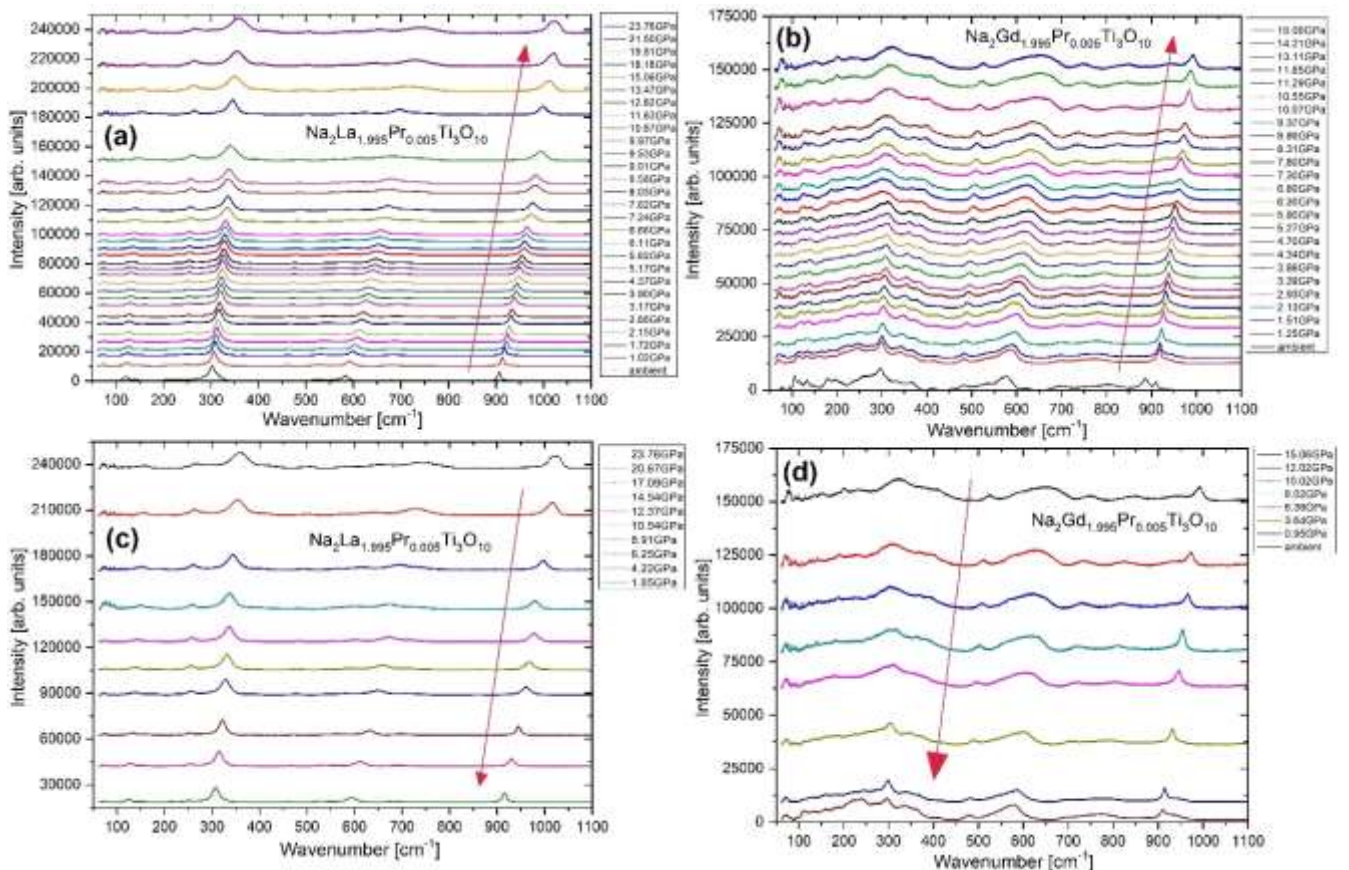


Figure A1 Pressure dependence of Raman scheelite modes (solid symbols), emerging fergusonite (open symbols), and the HP BaWO₄-II bands (semisolid symbols) up to 14.1 GPa. Dotted lines above 7 GPa show the onset of the scheelite-to-fergusonite phase transition, while dashed lines below 9 GPa the other related to the fergusonite - monoclinic BaWO₄-II change.



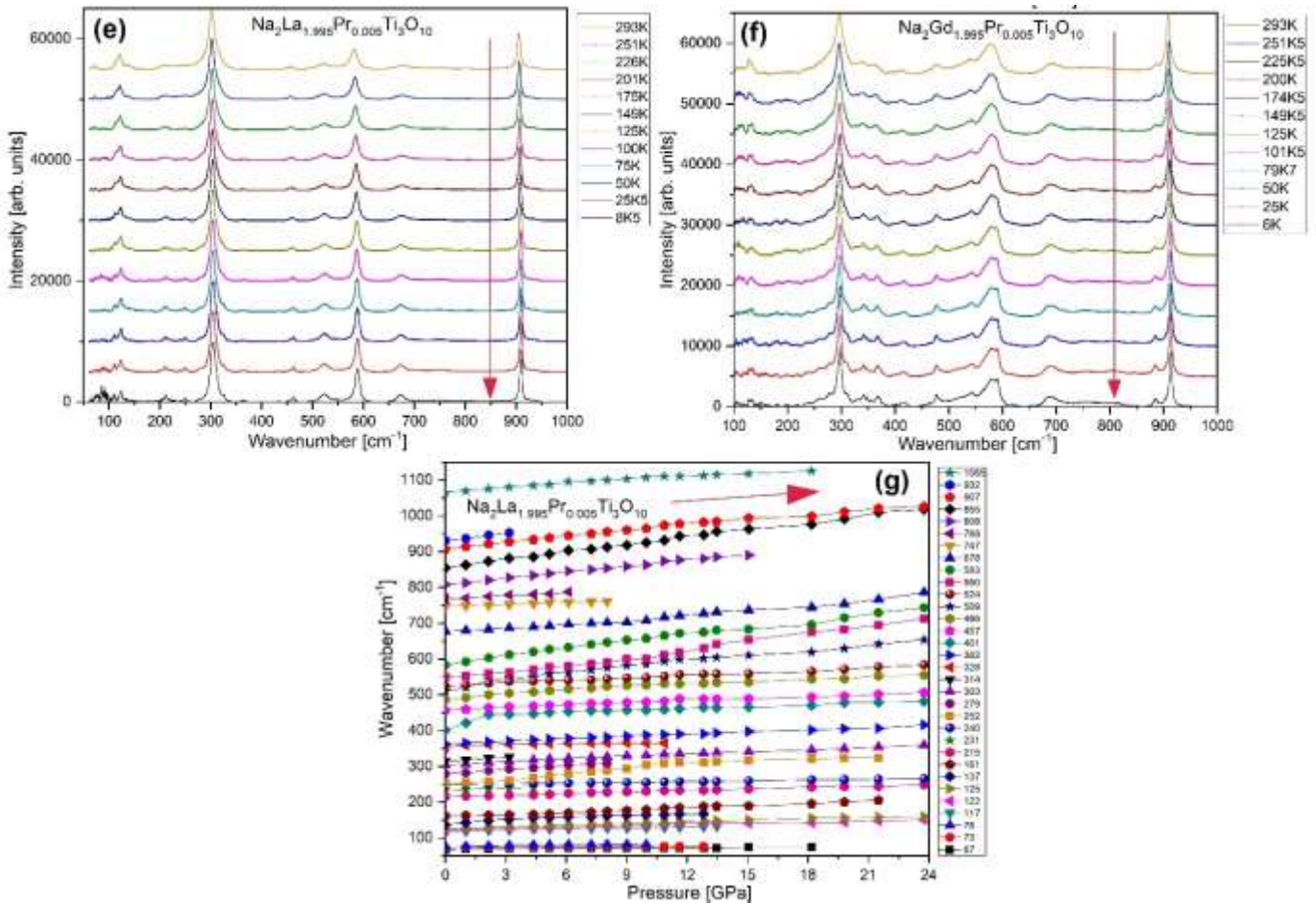


Figure A2 Tripple-decker, Ruddlesden-Popper sodium-titanate perovskites doped with 2.5% Pr^{3+} and their polymorphism viewed via Raman spectroscopy. The left figures relate to the $\text{Na}_2\text{La}_2\text{Ti}_3\text{O}_{10}$ matrix, while the right to $\text{Na}_2\text{Gd}_2\text{Ti}_3\text{O}_{10}$ in respective order: (a,b) compression, (c,d) decompression, (e,f) cooling from RT to cryogenic conditions. The bottom last graph (g) showcases peak versus pressure relations for lanthanide derivative solely to establish unique coefficients and calculate much-needed Grüneisen parameters.

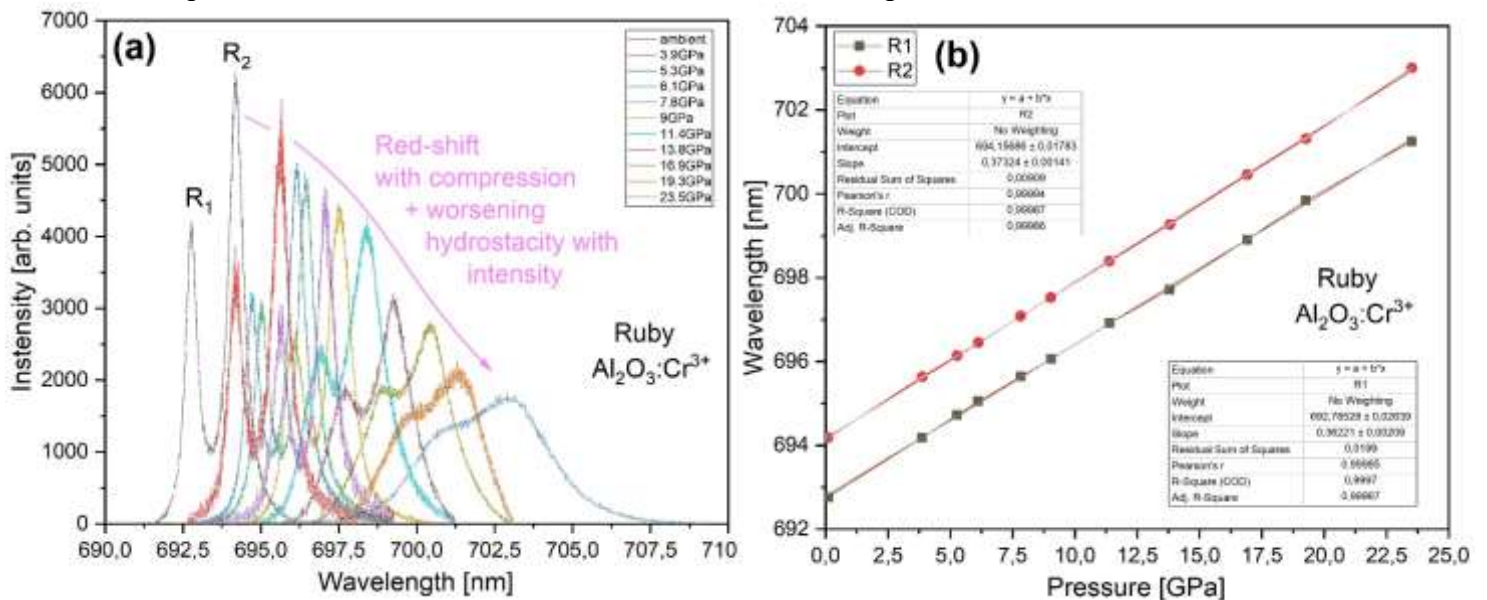


Figure A3 (a) An example of ruby fluorescence showcasing typical Cr^{3+} bands slowly red-shifting during compression. Pressure quenching and amorphization (loss of hydrostatic capabilities above 15 GPa) are also visible. Graph (b) depicts the position dependence of the exemplary R1 line and, more importantly, the R2 mode used as a reference for DAC interior pressure calculations provided by Syassen's and Mao's work.^{149,150}

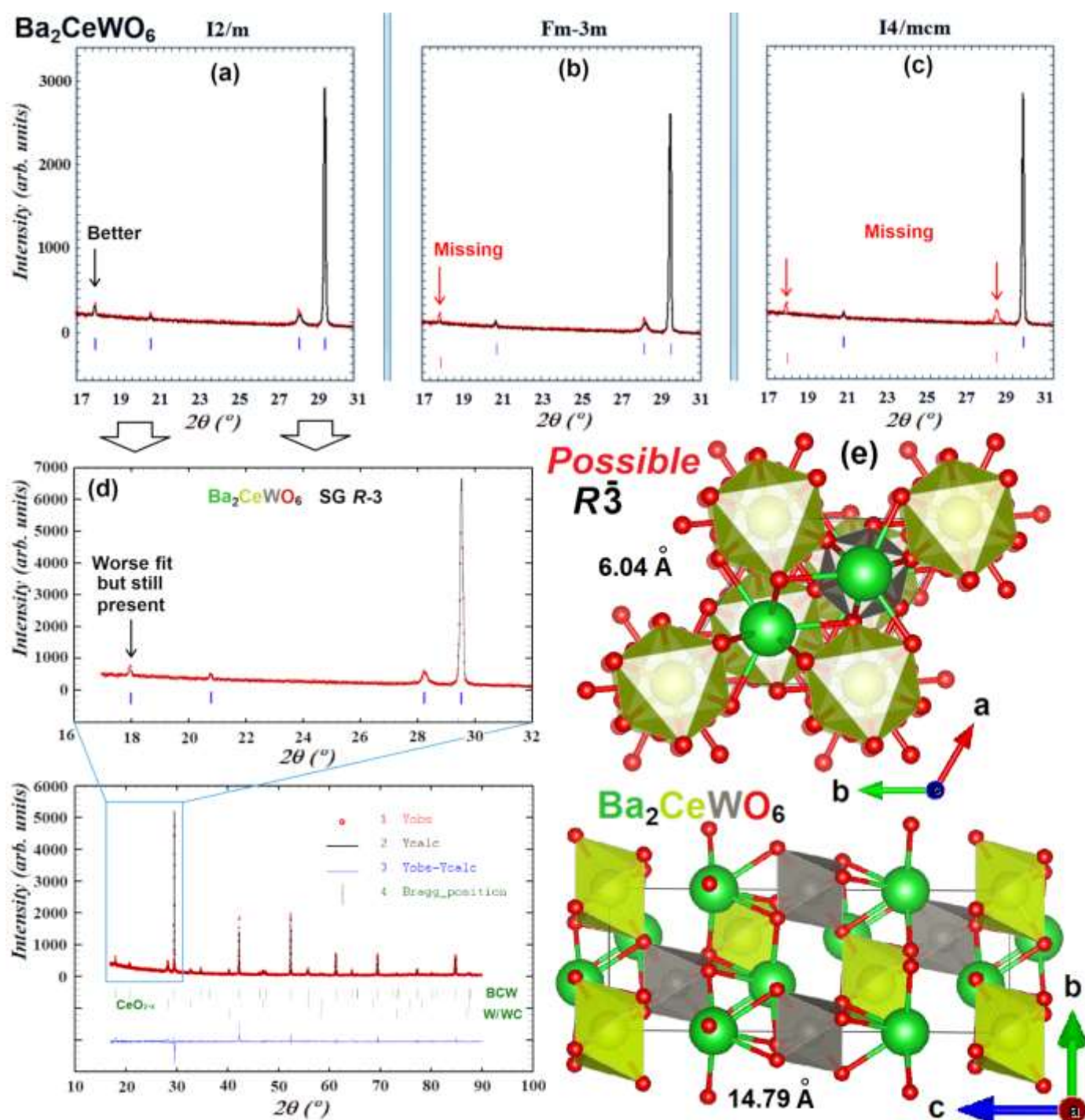


Figure A4 Case-specific powder XRD diffractograms showcasing a meticulous fitting process for the most possible SGs hosting BCW in a narrow 2θ region of interest. Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit. Blue and red vertical lines depict present and missing Bragg positions (green lines) compared to most ideal $I2/m$ SG. As one can see, only pattern (a) for $I2/m$ contains all four peaks, while cubic $Fm-3m$ at (b) and $I4/mcm$ in (c) came short of 1 or 2 positions, respectively. Figure (d) also matches all of them, however, a hkl at 18° (2θ) lacks in matters of intensity – closely associated with it, alternative $R-3$ unit cell is presented in mixed polyhedral/ball-stick convention to the right – Figure (e).

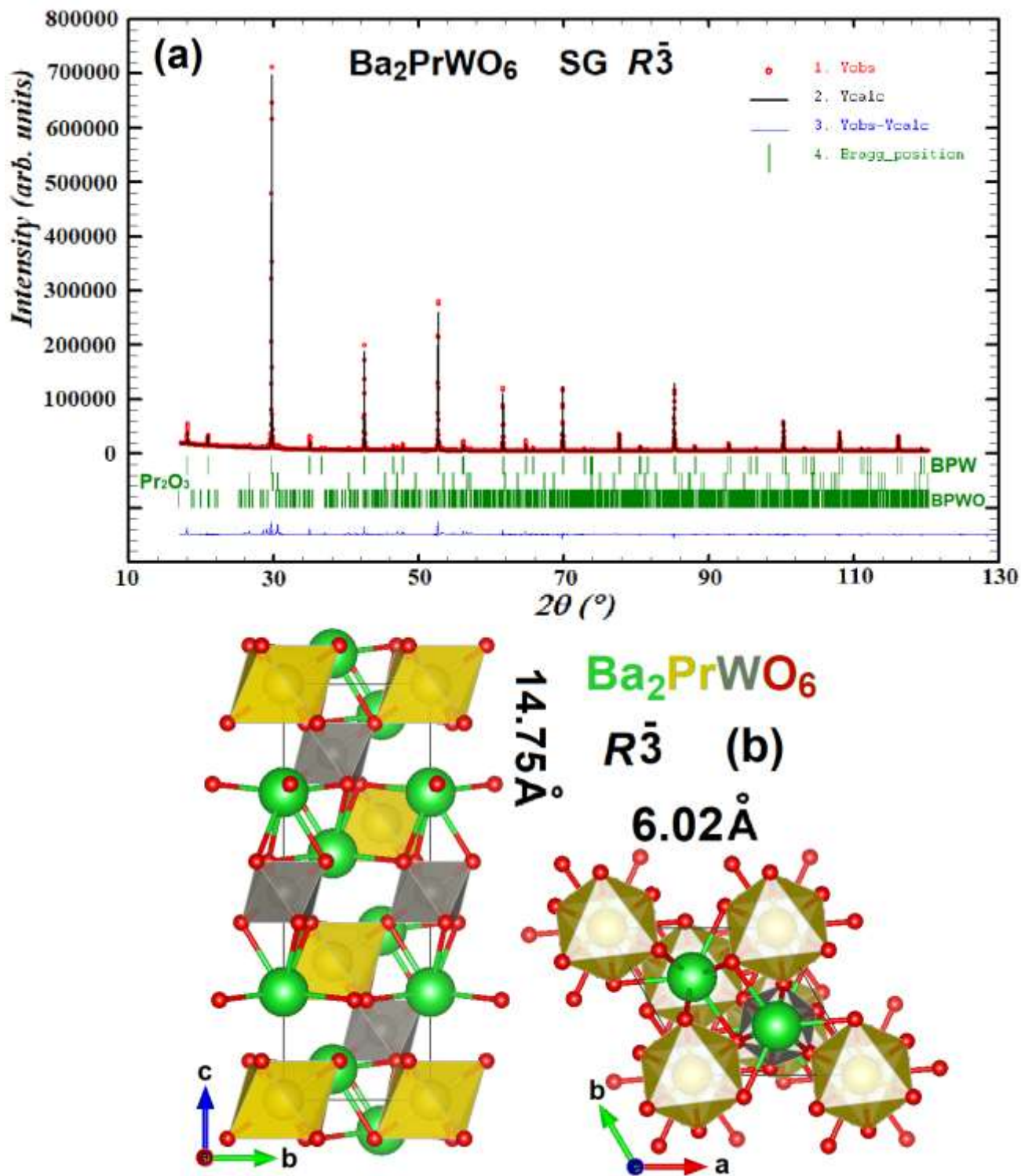


Figure A5 Alternate, almost identical XRD pattern fitting acquired post-Rietveld refinement to (a) $R\bar{3}$ phase for purest ($\sim 96\%$) Ba_2PrWO_6 (BPW) batch still containing traces of BaPr_2WO_7 and distorted Pr_2O_3 oxide impurities. Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown in partially ball-stick, partially polyhedral conventions in (b) with some additional structural parameters alongside the c (right) and a (left) axes. Yellow spheres and polygons are Pr ions. Red is oxygen, and grey is assigned to tungsten space filling. Barium is depicted solely as green spheres to improve overall visibility since its dodecahedra would cover most of the view.

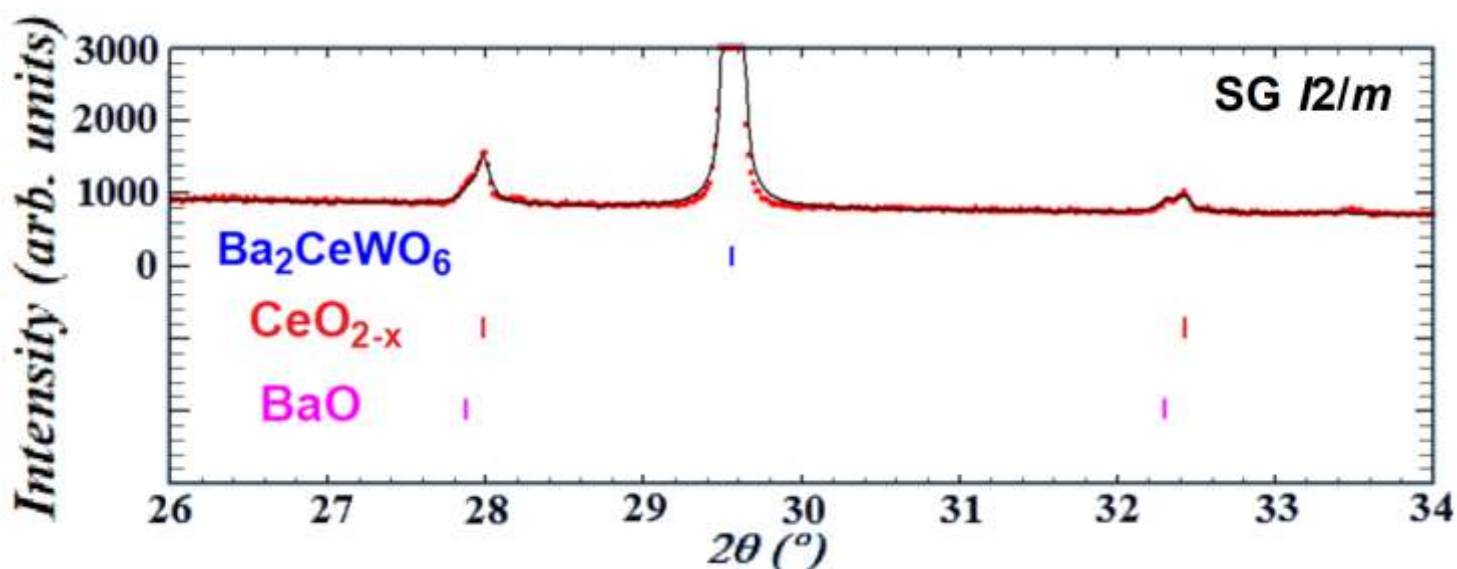
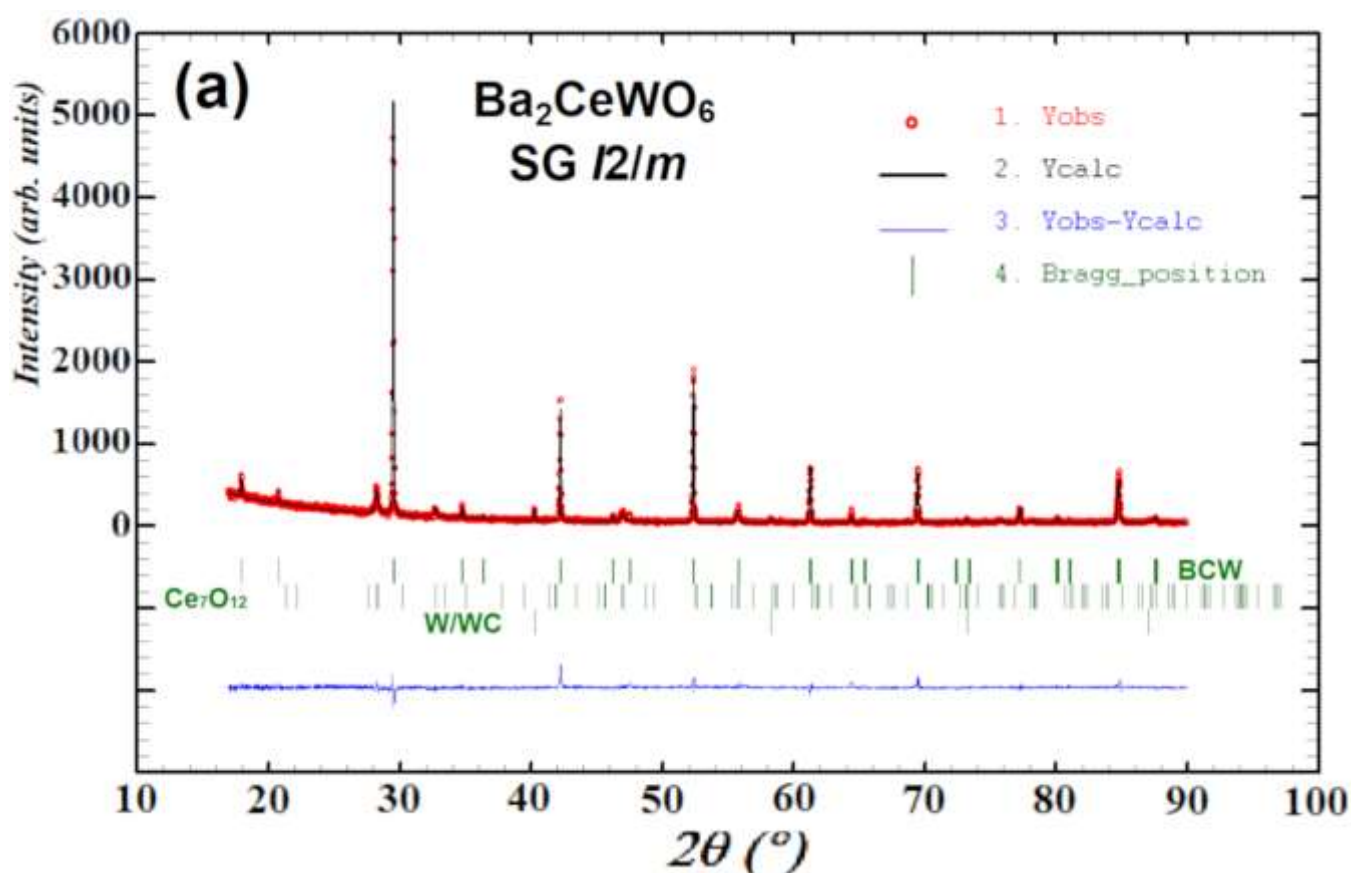


Figure A6 Powder XRD diffractogram showcasing a typical $I2/m$ fit for one of the worse BCW batches. As seen, quite a few minor phases could form near crucial Bragg positions, such as here - highly-distorted CeO_{2-x} and BaO oxides are the culprits. Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit. The blue, red, and pink vertical lines relate to specific compounds and their Bragg positions – BCW, CeO_{2-x} , and BaO , respectively. Those impurities form due to the evaporation of 1 necessary reactant - WO_2 .



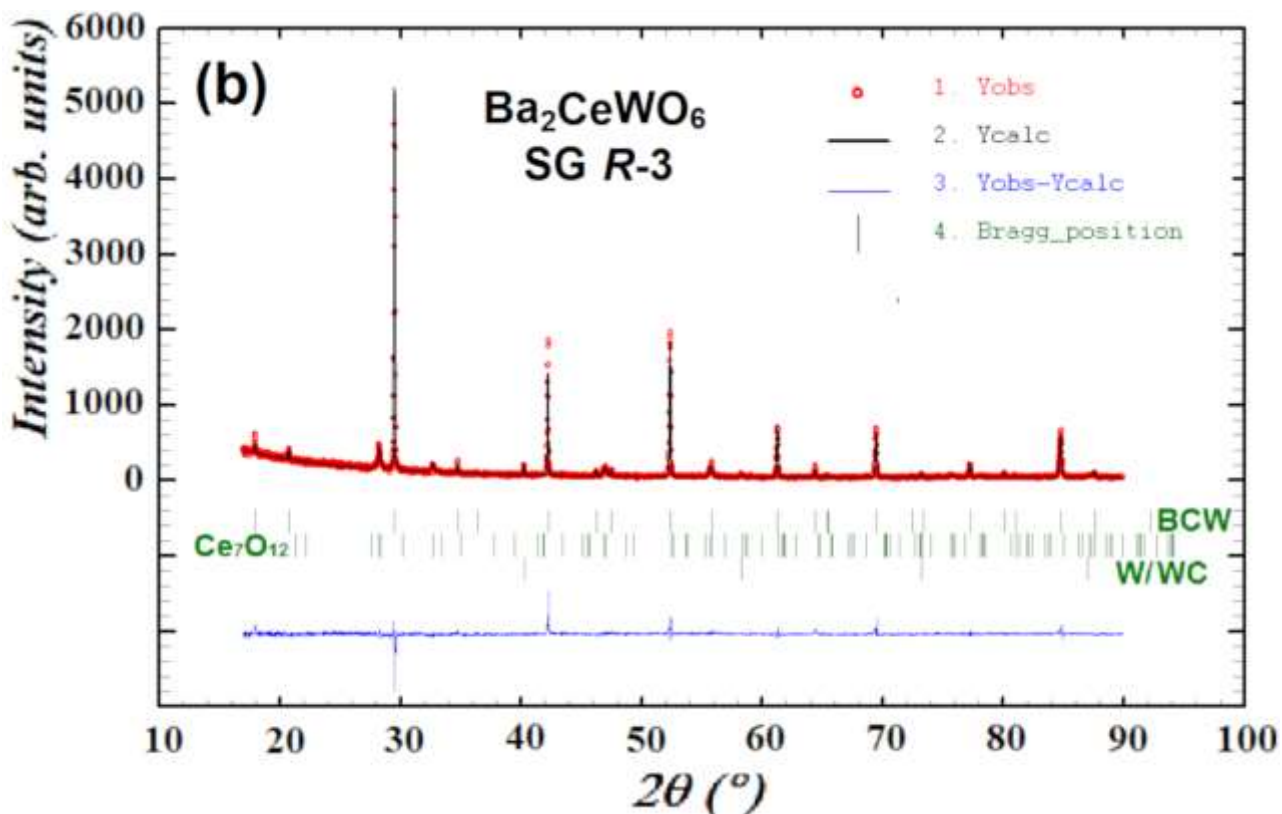


Figure A7 Powder XRD diffractogram fitting for one of the better BCW batches shows that in a few particular cases, a minor phase, such as highly-distorted CeO_{2-x} , could form an almost perfect Ce_7O_{12} pattern. It is probably a targeted form of this undesired impurity concerning almost identical: (a) $I2/m$ and (b) $R-3$ SG patterns. Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit. The green vertical lines relate to specific compounds and their Bragg positions.

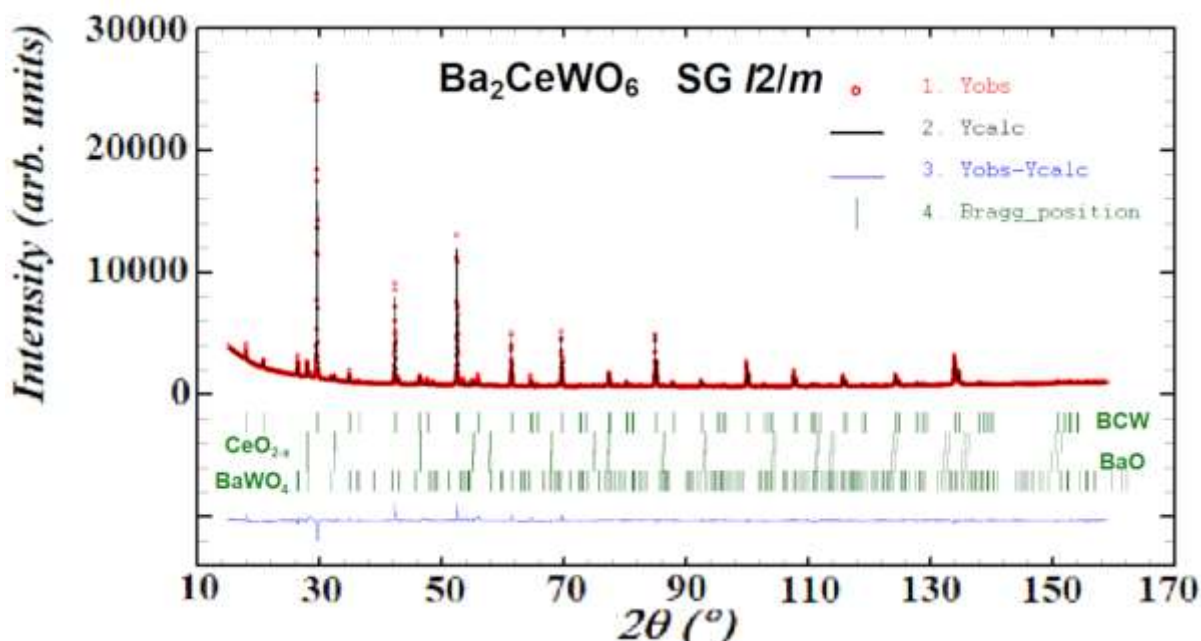


Figure A8 Powder XRD diffractogram fitting for one of the worst BCW batches showing that quite a few minor phases could form: not only highly-distorted CeO_{2-x} and BaO present due to WO_2 evaporation but also quite popular BaWO_4 created when an excessive amount of air is not removed from the furnace environment. Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit, and the green vertical lines relate to specific compounds and their Bragg positions – BCW , CeO_{2-x} , BaO , and BaWO_4 .

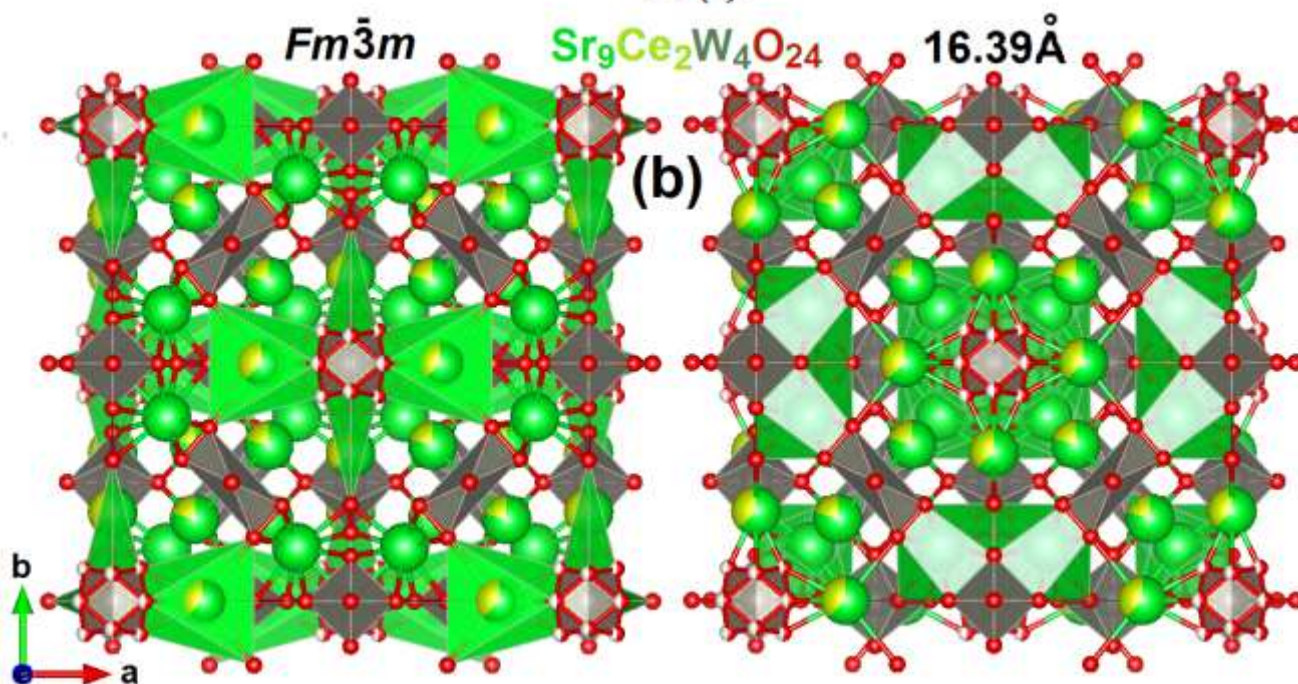
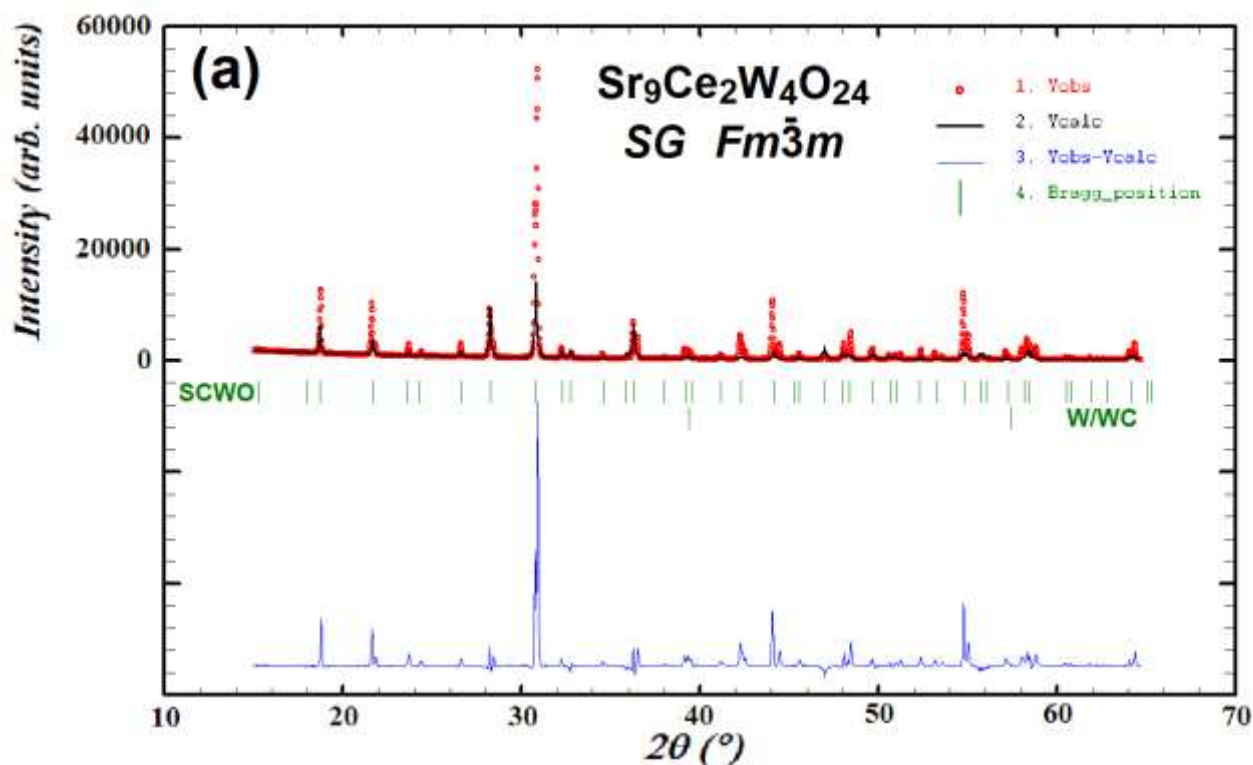


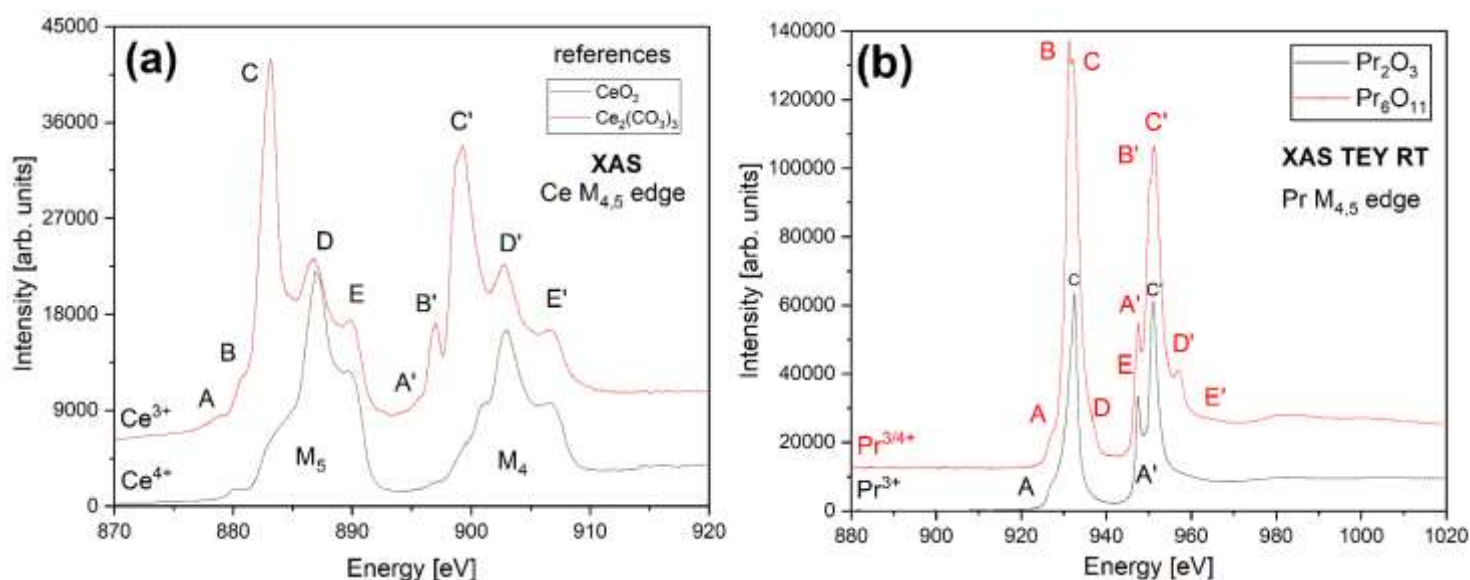
Figure A9 Alternate, closest XRD pattern fitting acquired post-Rietveld refinement for the most probable (a) $Fm\bar{3}m$ phase of $Sr_9Ce_2W_4O_{24}$ (SCWO) still containing traces of W/WC ceramics as impurities (~99% pure batch). Red dots are the manifestation of an actual, observed pattern. A straight black line is a calculated fit. The blue line depicts the difference between them. Green vertical lines are specific Bragg positions assigned to specific phases. The unit cell is shown using partially ball-stick, partially polyhedral convention in figure (b), with additional structural parameters alongside the main axis. The right projection shows shared Sr/Ce polyhedral sites, while the left depicts only those specific for Sr. Green spheres and polygons are strontium atoms. Yellow is cerium. Red is oxygen, and grey is assigned to tungsten.

Table A1 Conventional Rietveld reliability factors for all the alternative space groups investigated, undoped compounds and their minor phases present in the best possible batch. Data refers to diffractograms hosting Bragg contribution patterns corrected with the background concerning SG that will be disputed and deemed less plausible (but not entirely discarded) in the conclusions.

Formula	SG	Z	V [Å ³]	d _{cal} [g/cm ³]	R _B	R _P	R _{WP}	R _{EXP}	N _σ GoF	χ ²	Fract [%]
Ba ₂ CeWO ₆	<i>R</i> -3	3	467.009	7.410	13.6						94.73
CeO _{2-x}	<i>Fm</i> -3 <i>m</i>	4	164.207	6.962	42.3	23.7	20.9	4.41	1295.277	23.0	2.41
BaO	<i>Fm</i> -3 <i>m</i>	4	168.007	6.062	37.0						2.55
W(/WC)	<i>Im</i> -3 <i>m</i>	2	31.698	19.263	93.8						0.31
Ba ₂ PrWO ₆	<i>R</i> -3	3	463.261	7.478	6.02						95.62
BaPr ₂ WO ₇	<i>P</i> 2 ₁ / <i>b</i>	4	623.894	7.612	28.7	18.5	17.5	2.66	2736.675	43.1	3.98
Pr ₂ O ₃	<i>P</i> -3 <i>m</i>	1	77.682	7.050	57.9						0.40
Sr ₉ Ce ₂ W ₄ O ₂₄	<i>Fm</i> -3 <i>m</i>	8	4407.122	6.604	15.2	76.9	80.2	3.77	14242.9	45.2	99.34
W(/WC)	<i>Im</i> -3 <i>m</i>	2	31.621	19.260	75.3						0.66
Ca ₃ Ce ₂ W ₂ O ₁₂	<i>R</i> 3 <i>c</i>	18	4528.403	6.339	4.95	8.89	10.4	2.70	900.651	14.8	99.61
W(/WC)	<i>Im</i> -3 <i>m</i>	2	31.678	19.257	70.8						0.39

Table A2 Crystallographic parameters obtained for all relevant, alternate, and undoped structures stored in the CIF CCDC database. Though R_{wp} suggests all mentioned rhombohedral SG are indistinguishable projections, they were moved to the appendix because the discussion renders them less plausible at the end, although not fully rejectable.

Formula	Main SG	Unit cell axis length [Å]			Angles [°]			Volume [Å ³]
		a [100]	b [010]	c [001]	α	β	γ	
Ba ₂ CeWO ₆	<i>R</i> -3	6.03818(8)		14.7905(4)	90.000		120.000	467.00(9)
Ba ₂ PrWO ₆	<i>R</i> -3	6.02194(4)		14.75098(2)	90.000		120.000	463.26(1)
Sr ₉ Ce ₂ W ₄ O ₂₄	<i>Fm</i> -3 <i>m</i>		16.395(3)			90.000		4407.12(2)
Ca ₃ Ce ₂ W ₂ O ₁₂	<i>R</i> 3 <i>c</i>	9.72582(3)		55.2791(2)	90.000		120.000	4528.40(3)



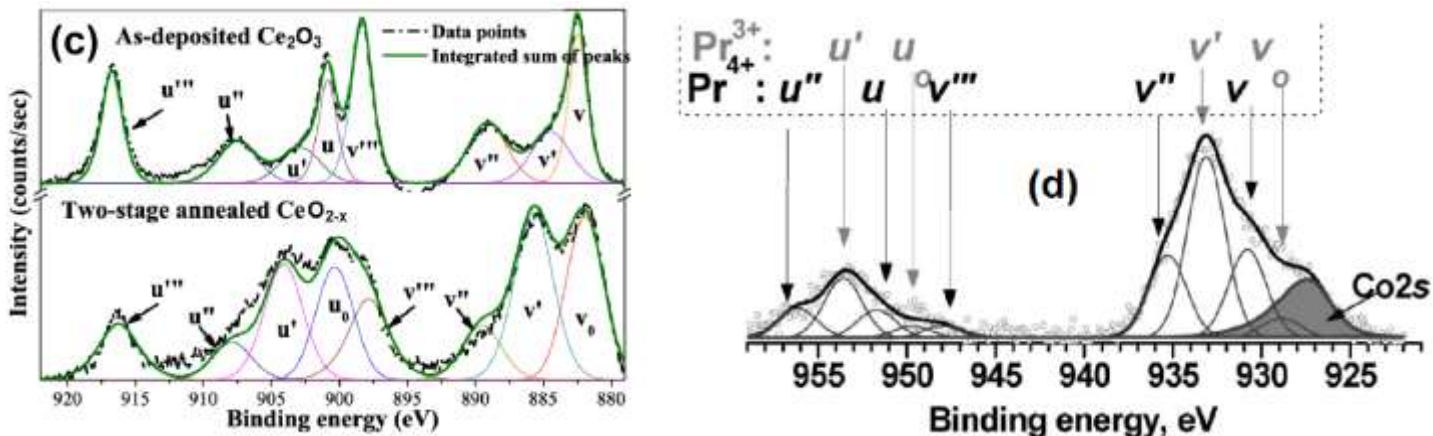


Figure A10 Reference XAS spectra taken at RT and ultra-high vacuum off 99.9%+ pure substrates used for synthesis: (a) cerium and (b) praseodymium oxides confronted with respective literature data shown in graphs (c) and (d) below.^{219,220}

Figure (c) was obtained from R. K. Nachimuthu et al. research entitled ‘*Investigation of Cerium-Substituted Europium Iron garnets Deposited by Biased Target Ion Beam Deposition*’, IEEE Transactions on Magnetics, 2014, **50** (12), pp. 1-7 copyrighted and licensed by IEEE. Figure (d) is borrowed from article E. Y. Konyshva and M. V. Kuznetsov ‘*Fluctuation of surface composition and chemical states at the hetero-interface in composites comprised of a phase with perovskite structure and a phase related to the Ruddlesden–Popper family of compounds*’, RSC Adv. 2013, **3** (33), pp. 14114-14122, copyrighted and licensed by the Royal Society of Chemistry.

Table A3 Factor group analysis for all alternative-yet-relevant space groups assigned to all discussed tungstates.

Material (SG)	Ion	Wyckoff Positions	Point Symmetry	Red. Representation
BCW	A ²⁺ (Ba)	2c	C ₃	A _g +A _u + ¹ E _u + ² E _u + ¹ E _g + ² E _g
BPW	B ^{3+/4+} (Ce)	1a	C _{3i}	A _u + ¹ E _u + ² E _u
R-3 (C _{3i})	B ^{5+/4+} (W)	1b	C _{3i}	A _u + ¹ E _u + ² E _u
	O ²⁻	6f	C ₁	3A _g +3A _u +3 ¹ E _u +3 ² E _u +3 ¹ E _g +3 ² E _g
$\Gamma_{\text{TOTAL}} = 4A_g + 6A_u + 6^1E_u + 6^2E_u + 4^1E_g + 4^2E_g$; $\Gamma_{\text{ACUSTIC}} = A_u + ^1E_u + ^2E_u$; $\Gamma_{\text{IR}} = 5A_u + 5^1E_u + 5^2E_u$; $\Gamma_{\text{Raman}} = 4A_g + 4^1E_g + 4^2E_g$				
SCWO	B ^{6/5+} (W)	4a, 4b, 24d	D _{3d} , C _{2h}	5T _{1u}
	B ^{3/4+} (Ce)	24e, 32f	C _{2v} , C ₃	2A _{1g} +2E _g +3T _{2g} +4T _{1u}
Fm-3m (O _h)	A ²⁺ (Sr)	24e, 32f	C _{2v} , C ₃	2A _{1g} +2E _g +3T _{2g} +4T _{1u}
	O ²⁻	24e, 48g, 48h	C _{2v} , C ₂ , C _{1h}	3A _{1g} +4E _g +6T _{2g} +8T _{1u}
$\Gamma_{\text{TOTAL}} = 5A_{1g} + 8E_g + 12T_{2g} + 21T_{1u}$; $\Gamma_{\text{ACUSTIC}} = T_{1u}$; $\Gamma_{\text{IR}} = 20T_{1u}$; $\Gamma_{\text{Raman}} = 5A_{1g} + 8E_g + 12T_{2g}$				
CCWO	B ^{6/5+} (W)	2a, 6b	C ₁ , C ₃	4A ₁ +8E
	B ^{3/4+} (Ce)	2a, 6b	C ₁ , C ₃	4A ₁ +8E
R3c (C _{3v})	A ²⁺ (Ca)	2a, 6b	C ₁ , C ₃	4A ₁ +8E
	O ²⁻	6b	C ₁	3A ₁ +5E
$\Gamma_{\text{TOTAL}} = 15A_1 + 29E$; $\Gamma_{\text{ACUSTIC}} = A_1 + E$; $\Gamma_{\text{IR}} = 7A_1 + 14E$; $\Gamma_{\text{Raman}} = 7A_1 + 14E$				

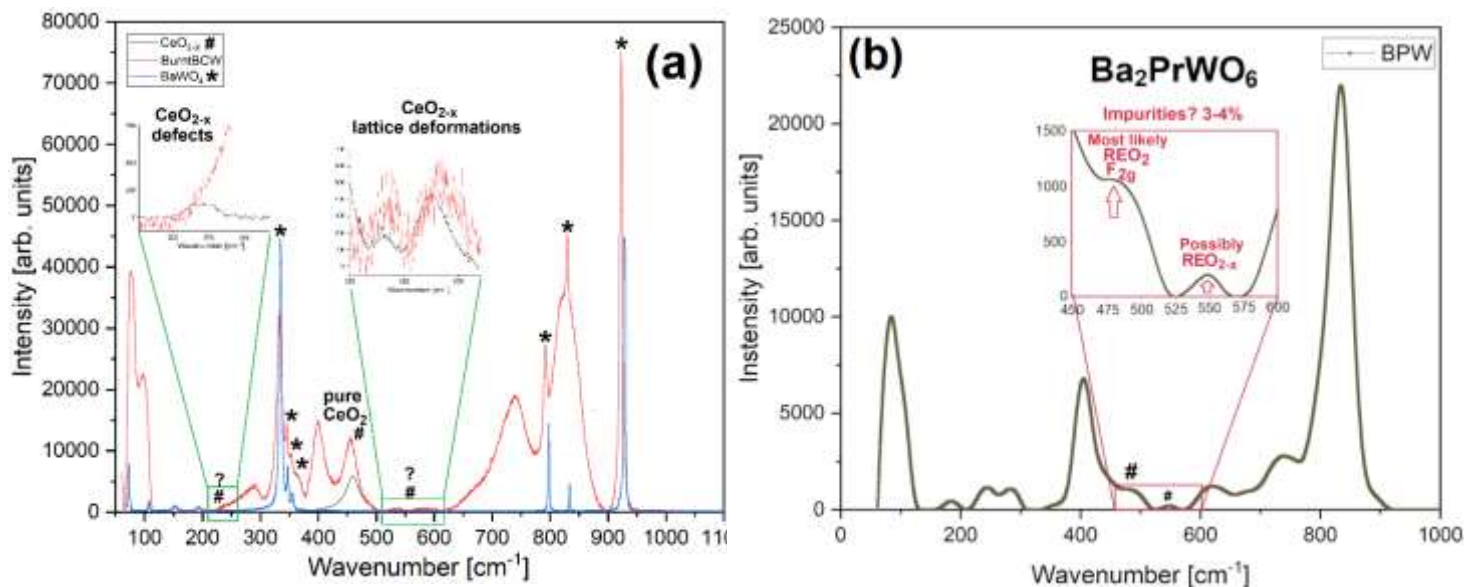


Figure A11 Raman spectra showcasing popular impurity referrals such as CeO_{2-x} (black) or BaWO_4 (blue lines) that can also transpire in worse or laser-burnt (a) BCW (red line) and (b) BPW batches analogically showcasing PrO_{2-x} . Visible peaks are pinpointed via hashtags (REO_{2-x}) and asterisks (BaWO_4) to guide the eye. One can also associate the former group with defective transverse $\text{RE}^{3/4+}\text{O}_x$ modes, which have similar scattering energy.

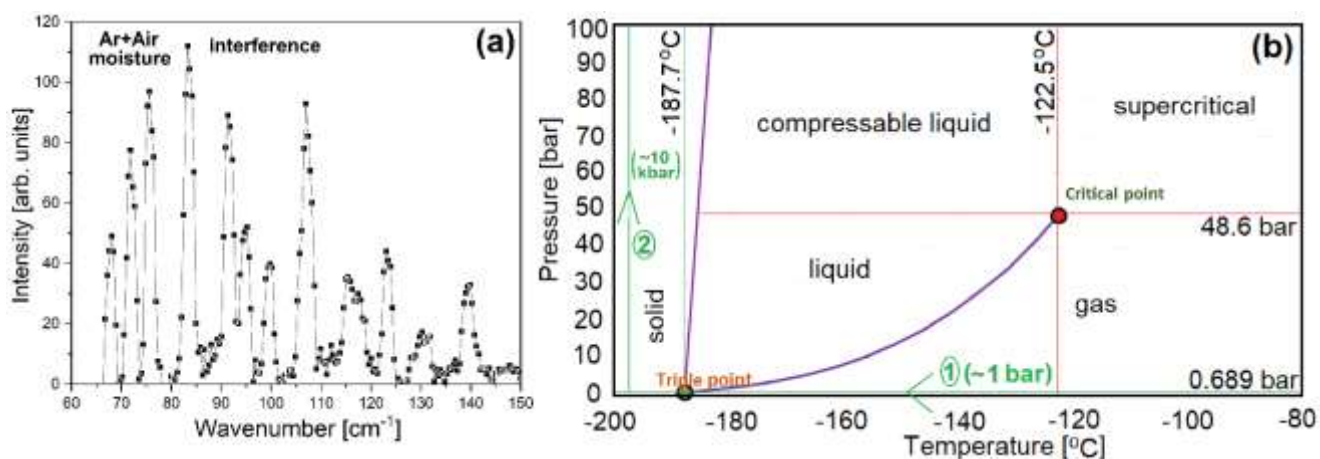


Figure A12 (a) Minor, static scattering interference registered at the beginning of each HP Raman scattering spectra with and without the sample in Ar-filled DAC and (b) a p(T) phase diagram of Argon featuring a way of preparing solid PTM after 1) DAC initial cooling and then 2) compressing it via hydraulic drive.

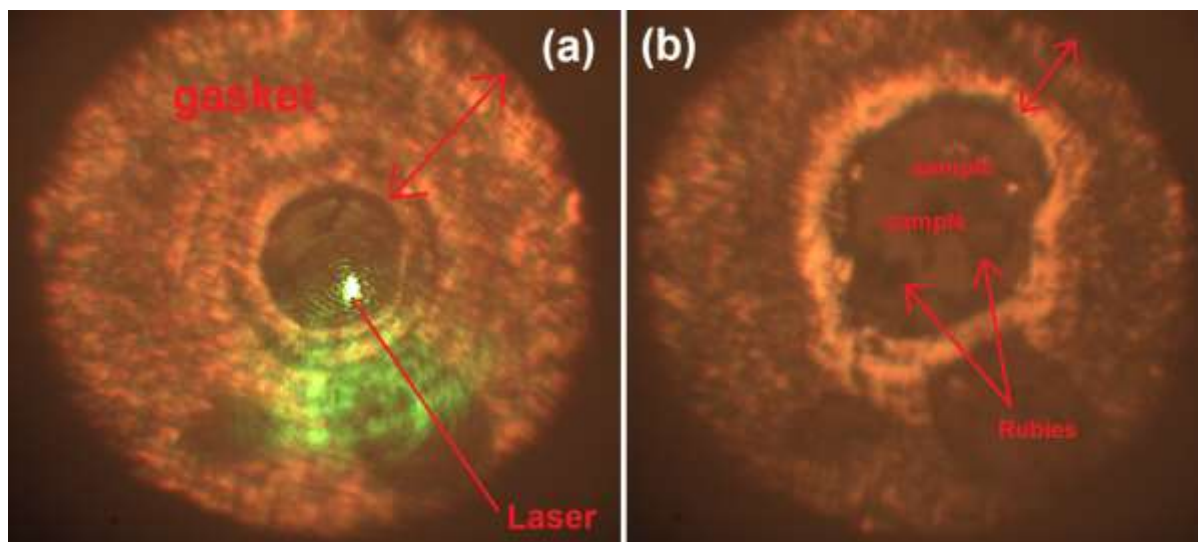


Figure A13 Digital microscopic photographs of DAC gasket filled with Ar PTM, referential (PL) pressure-gauging ruby, and powder from BCW at (a) the beginning and (b) end of the HP experiment (near its breaking point) featuring deformations and typically encountered laser size spot.

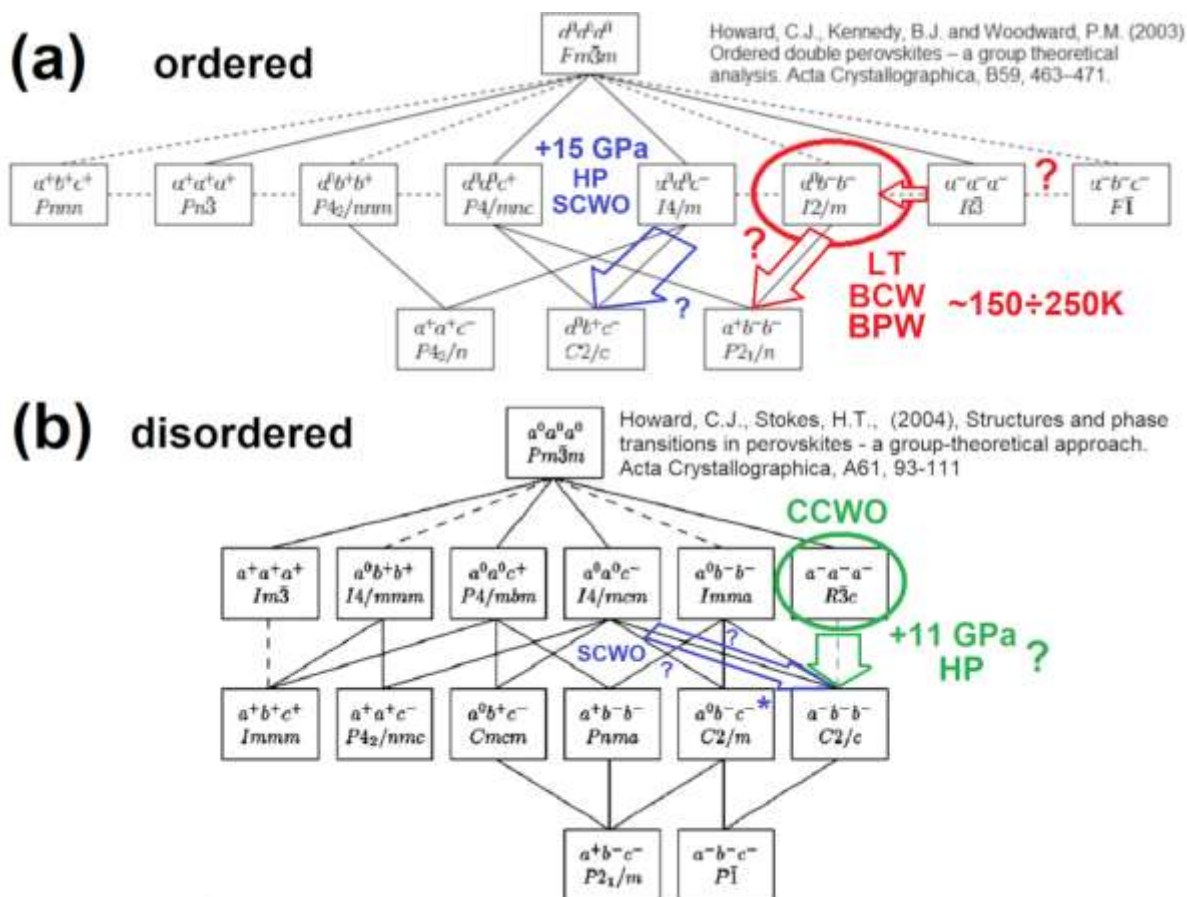


Figure A14 Particular group–subgroup relations associated with interchangeable $A_2BB'X_6$ tilt systems in (a) ordered (like BCW or BPW - red) and (b) disordered, perovskites-like materials (ilmenite CCWO - green) evaluated by Howard et al. between 1998-2005 using Landau DFT calculations.^{71,72} Continuous black lines are second- and dashed first-order (Landau-Ginzburg) phase transitions. Possible sample placement has also been established artificially amongst possible SG. SCWO, as a disputable intermediate compound bearing similarities of both groups, is placed on both graphs (blue) despite not having any of the precisely mentioned SG ($I4_1/a$). Both graphics were partially adapted and reprinted with permissions from Howard C. J.; Kennedy B. J.; Woodward P. M. *Ordered double perovskites – a group-theoretical analysis*, *Acta Crystallogr.* **2003**, B59, pp. 463-471, and Howard C. J.; Stokes H. T. *Structures and phase transitions in perovskites – a group-theoretical approach*, *Acta Crystallogr.* **2004**, A61, pp. 93-111. They are copyrighted and licensed by IUCr Journals Publisher via *Acta Crystallographica B* journal after 2003 & 2004, respectively.

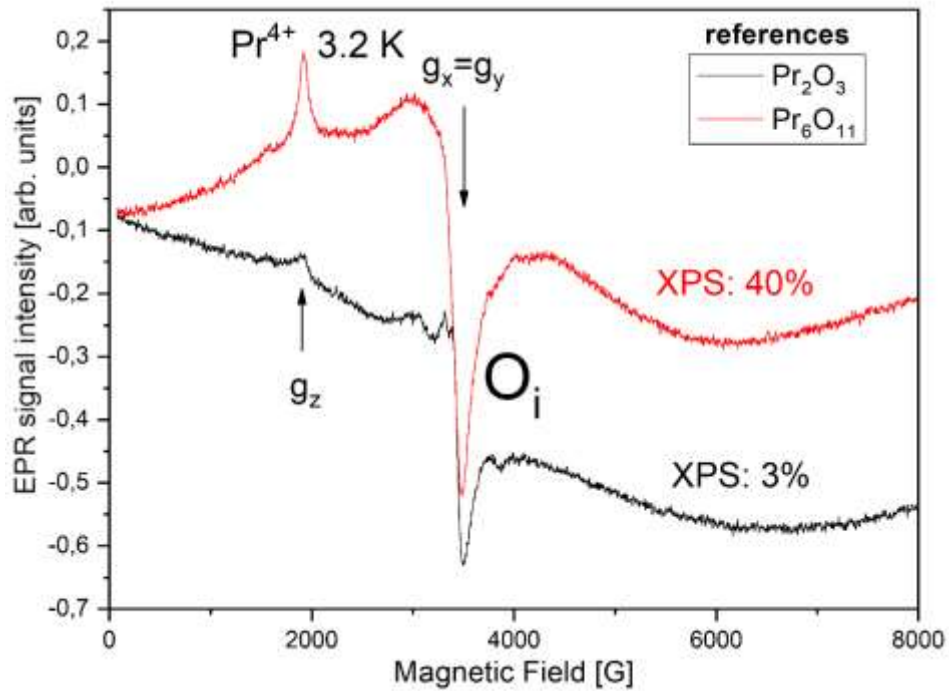


Figure A15 LT EPR spectra of Pr_6O_{11} & Pr_2O_3 substrates used as references for SSR synthesis showcasing the nature of both applied oxides. A weak Pr^{4+} response joins the place where O_i intrusions appear. It should emerge much more potent in perpendicular magnetic fields if there were no antiferromagnetic transitions and high enough concentration.

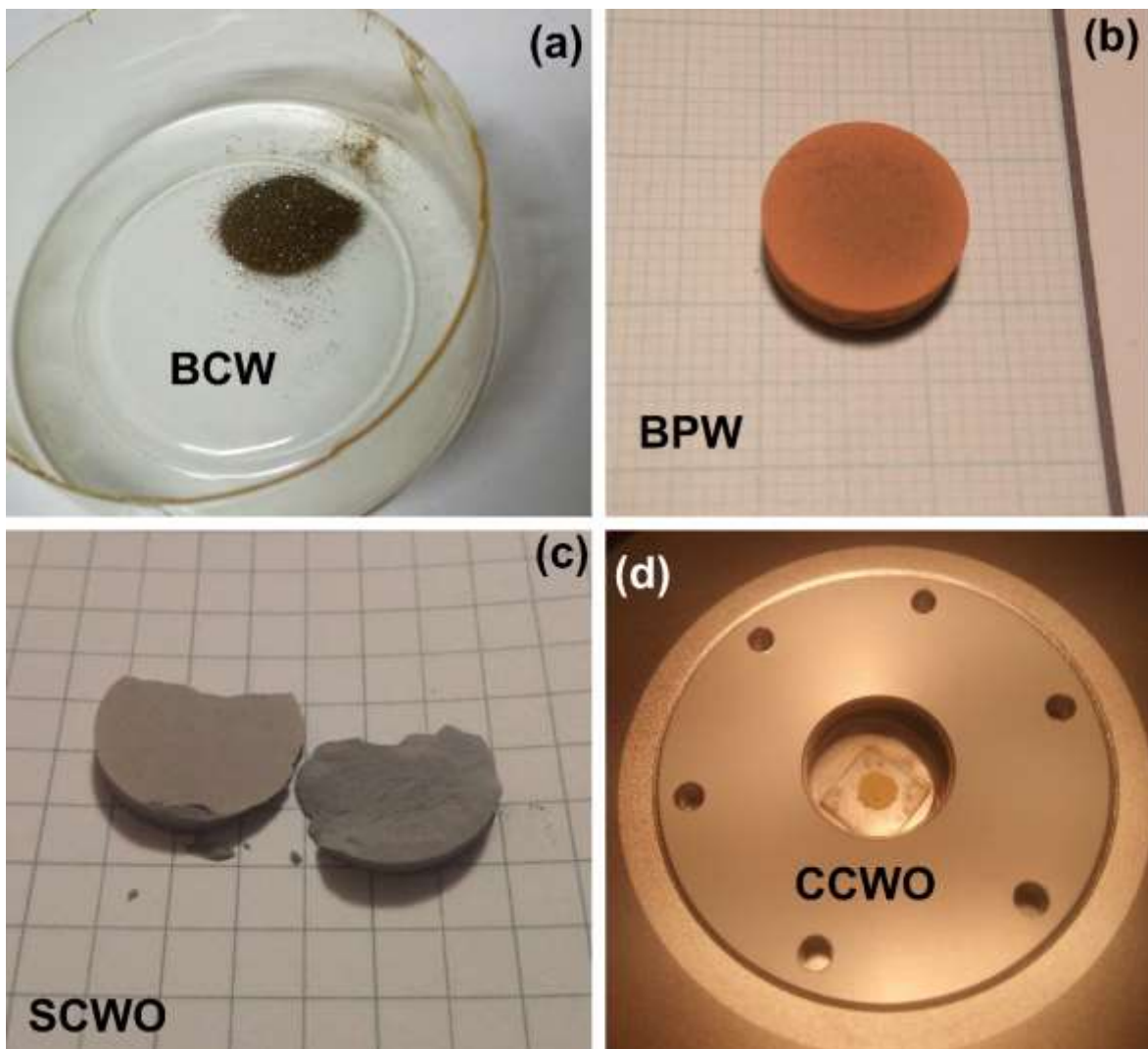


Figure A16 Macro photographs of all four investigated substances: (a) Ba_2CeWO_6 , (b) Ba_2PrWO_6 double perovskites, and (c) $\text{Sr}_9\text{Ce}_2\text{W}_4\text{O}_{24}$ with (d) $\text{Ca}_3\text{Ce}_2\text{W}_2\text{O}_{12}$ derivatives after decomposition in the air at HT (above II stage of decomposition mentioned in *Figure 35*).

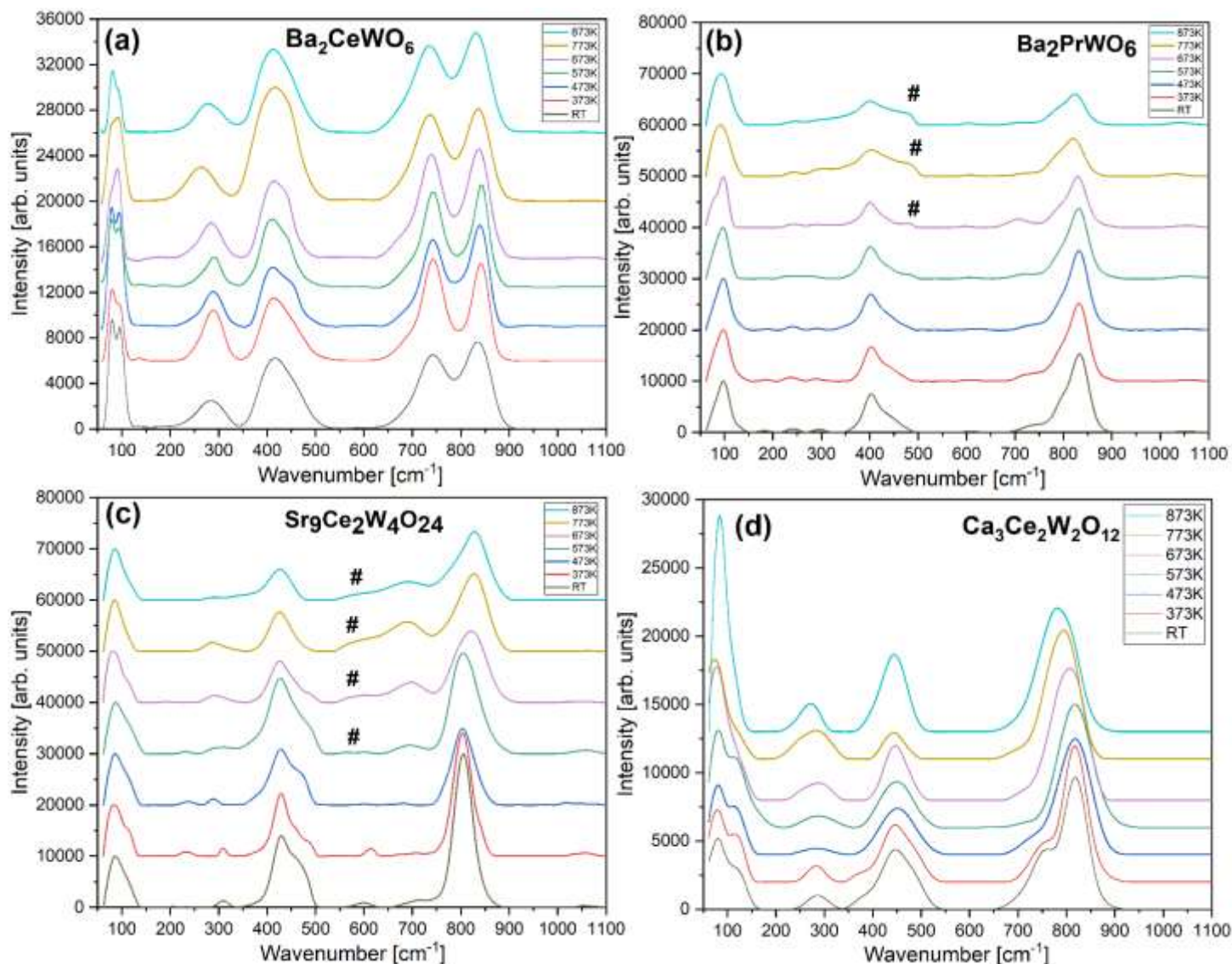


Figure A17 Raman spectra for all investigated compounds collected in the function of heating from room temperature up to 873K in a protective, inert N_2 atmosphere showing little to no change while subtle blue-shifts occur. Hashtags mark noticeable intensity changes in already denoted peaks within mostly defective materials, probably related to the internally stored O_i slowly creeping/integrating into the structures, creating many RE-O defects.

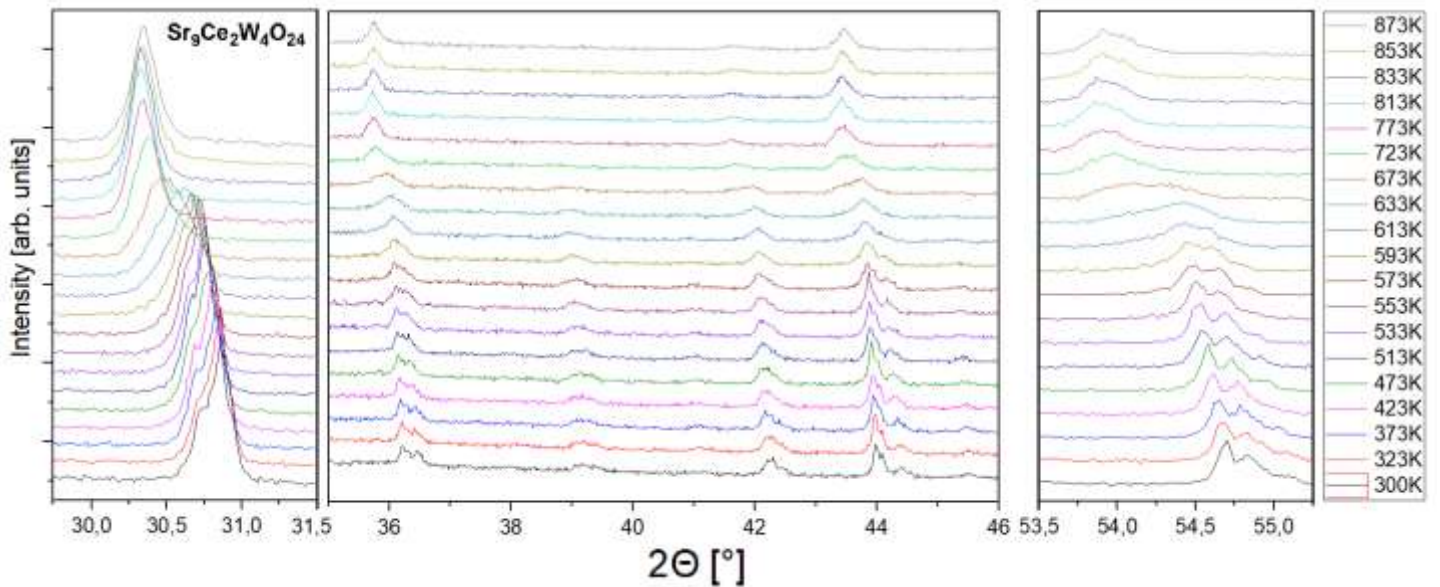


Figure A18 Close-ups on HT XRD powder diffractograms of SCWO quaternary perovskite done in the air. They feature some of the most prominent (decomposition) changes highlighted in densely packed *Figure 39d*.

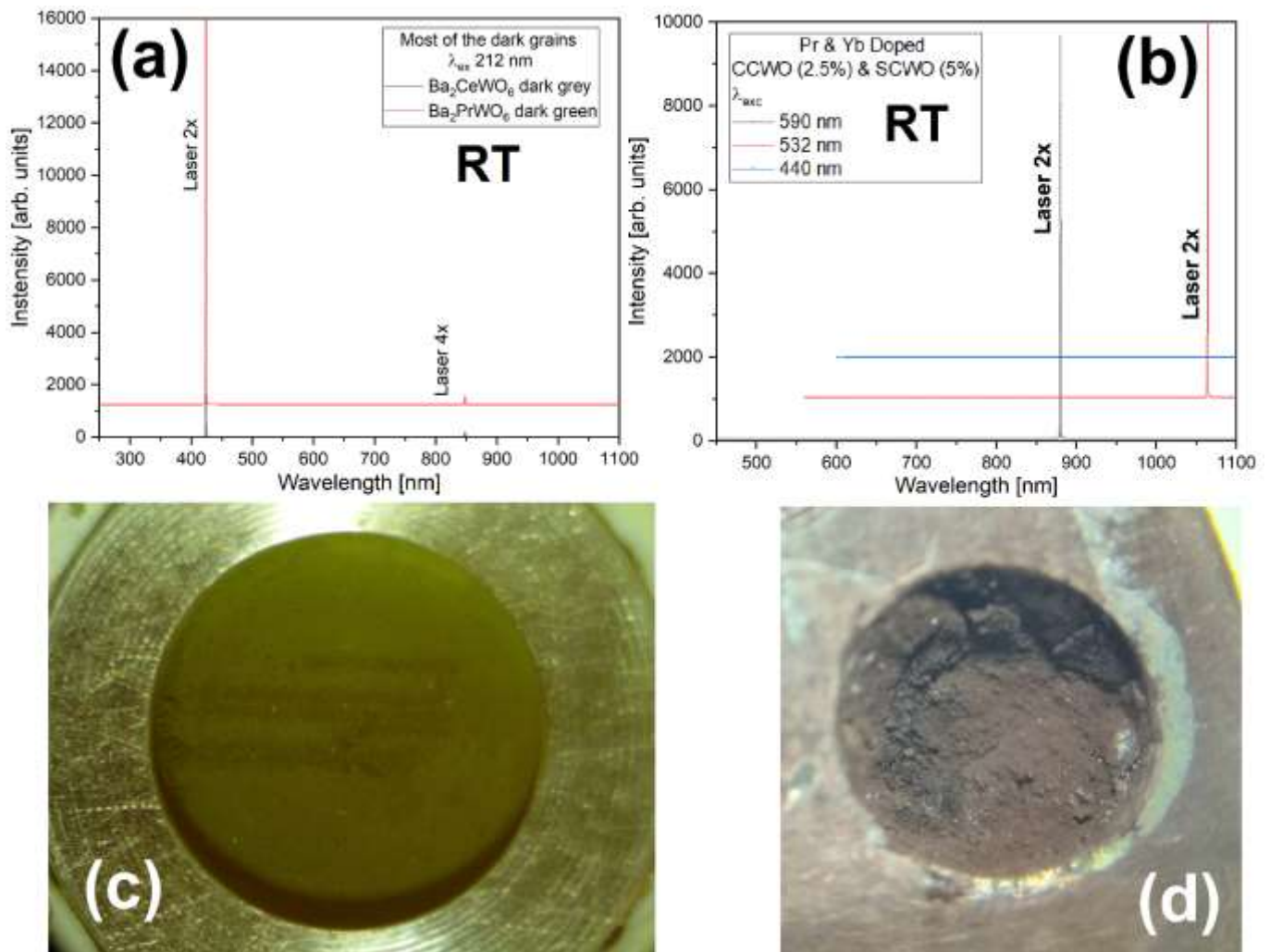


Figure A19 Inconsistent luminescence behavior of both: (a) undoped barium double perovskite materials, (b) doped tungstate derivatives given various excitation energies around possible absorption lines for typical RE^{3+} ions. Most samples tend to be physically drilled if exposed briefly to laser illumination (c) or respectively burned (d) due to disadvantageous dark coloring when exposed for a long time without any inert, protective atmospheres.

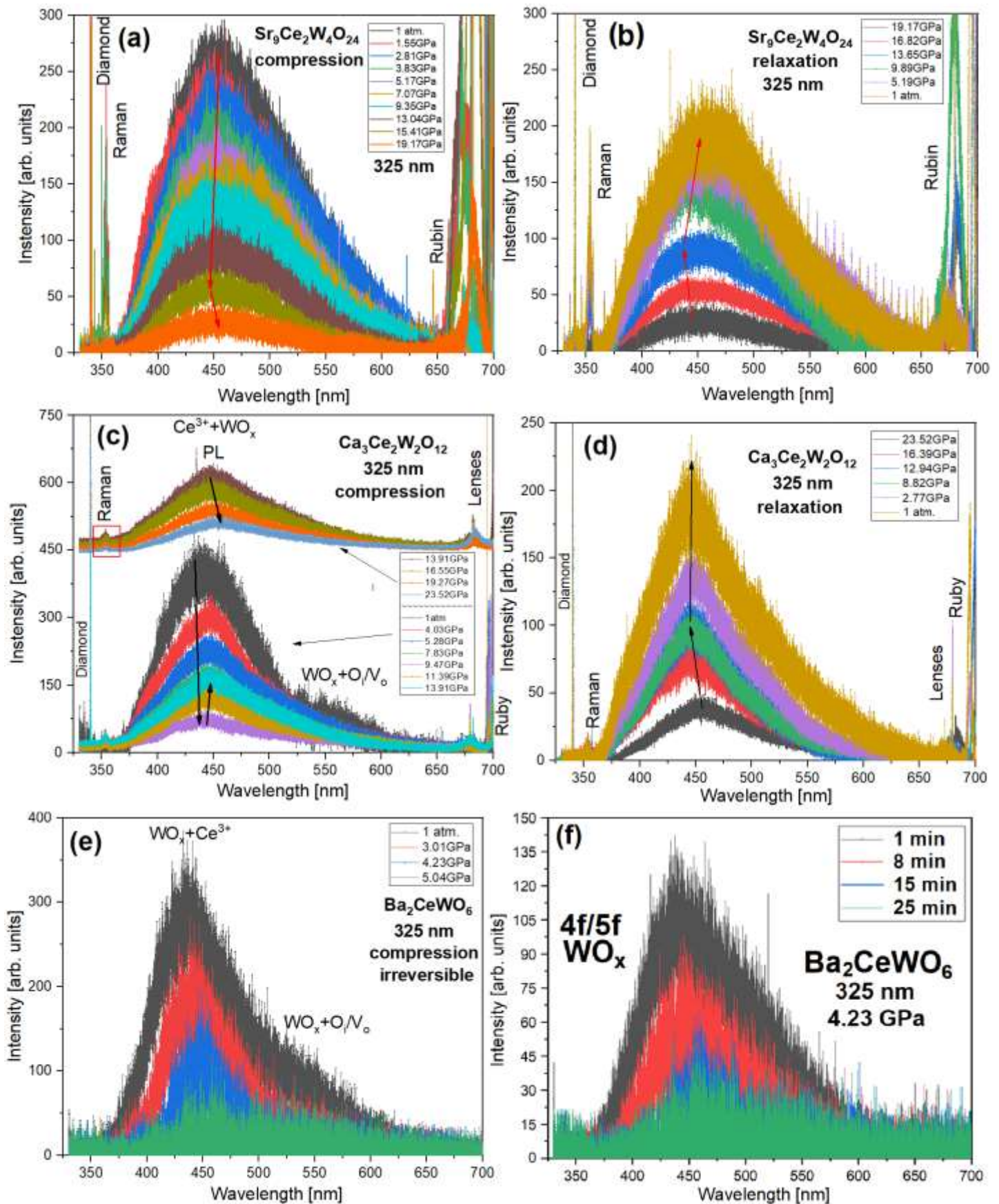


Figure A20 Alternating HP PL spectra of all undoped cerium tungstates presented in compression-decompression cycles respectively under NUV, 325 nm excitation: (a) & (b) for SCWO, (c) & (d) regarding CCWO, and irreversibly diminishing BCW light in (e) where it seems to be rather time- than pressure-resolved (f).

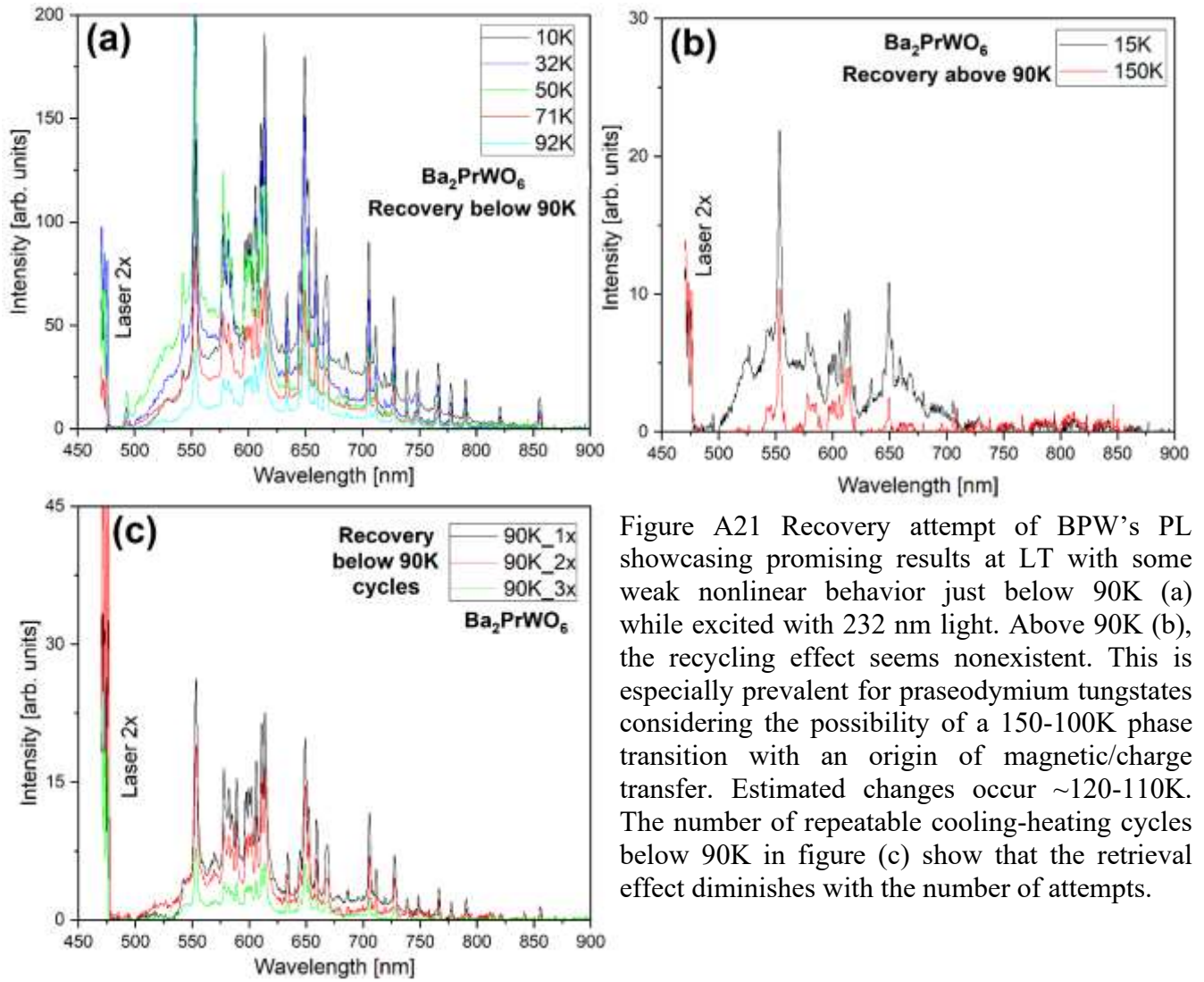


Figure A21 Recovery attempt of BPW's PL showcasing promising results at LT with some weak nonlinear behavior just below 90K (a) while excited with 232 nm light. Above 90K (b), the recycling effect seems nonexistent. This is especially prevalent for praseodymium tungstates considering the possibility of a 150-100K phase transition with an origin of magnetic/charge transfer. Estimated changes occur ~ 120 - 110 K. The number of repeatable cooling-heating cycles below 90K in figure (c) show that the retrieval effect diminishes with the number of attempts.

Equations A1:

$$\begin{aligned} \frac{dN^*}{dt} &= f(N - N^*) - (p + r(T))N^* \\ N &= \text{const.} \\ \frac{dN}{dt} &= -f(N - N^*) + (p + r(T))N^* \\ N^*(T) &= \frac{fN}{(f + p + r(T))} \quad N^*(T = 0) = \frac{fN}{(f + p)} = \frac{I_0}{p} \\ \frac{I_0}{I} - 1 &= r(T)/(f + p) \end{aligned}$$

where N - is the concentration of (e.g., Pr³⁺); N^* - occupation of the excited state (e.g., ¹S₀); f - is excitation density; 'p' and 'r' are the probabilities of radiative and non-radiative recombination, respectively; In this case, dynamic equilibrium is established quickly (in the order of milliseconds or less), which means that with constant excitation $\frac{dN^*}{dt} = \frac{dN}{dt} = 0$

Equations A2:

$$\begin{aligned} \frac{dN^*}{dt} &= f(N_0 - N^* - n) - (p + r)N^* - \beta N^* \exp\left(-\frac{E}{kT}\right) \\ \frac{dn}{dt} &= \beta N^* \exp\left(-\frac{E}{kT}\right) - wn^2 \end{aligned}$$

$$n = \sqrt{\frac{\beta}{w} N^* \exp\left(-\frac{E}{kT}\right)} \quad f(N_0 - n) = N^* \left[f + p + r + \beta \exp\left(-\frac{E}{kT}\right) \right]$$

$$N^*(T) = \frac{N_0}{a} - \frac{\sqrt{(\beta \exp(-E/kT)/w)^2 + 4N_0 a \beta \exp(-E/kT)/w}}{2a^2}$$

$$a = \left(1 + \frac{p + r + \beta \exp(-E/kT)}{f} \right)$$

$$N^*(T = 0) = \frac{N_0}{a(T = 0)}$$

N^* - concentration of ionized dopant, always equal to the concentration of electrons in the n band; N_0 - total dopant concentration (= const.); $N_{(t,T)} + n_{(t,T)} = N_0$; β - the probability of ionization (contains the density of states in the conduction band); w - a cross-section of e^- capture.

Equations A3:

$$\frac{dN^*}{dt} = f(N - N^*) - (p + r)N^* - \beta N^* \exp\left(-\frac{E}{kT}\right)$$

$$\frac{dn}{dt} = \beta N^* \exp\left(-\frac{E}{kT}\right) - w_M n(M_0 - M^-) - wn(N_0 - N)$$

$$\frac{dN}{dt} = -f(N - N^*) + (p + r)N^* + wn(N_0 - N)$$

$$\frac{dM^-}{dt} = w_M n(M_0 - M^-)$$

According to these equations, equilibrium will be reached after time 't' when $N(t) = n(t) = 0$ and $M^-(t) = N(t=0)$. This is a consequence of no trap ionization. It takes roughly 15 minutes at room temperature, as showcased in *Figures 50 & 55*. At low temperatures, it would be much longer. By measuring the luminescence, we measure the kinetics of Pr^{3+} decay, or rather its beginning, because the measurement lasts only about a minute.

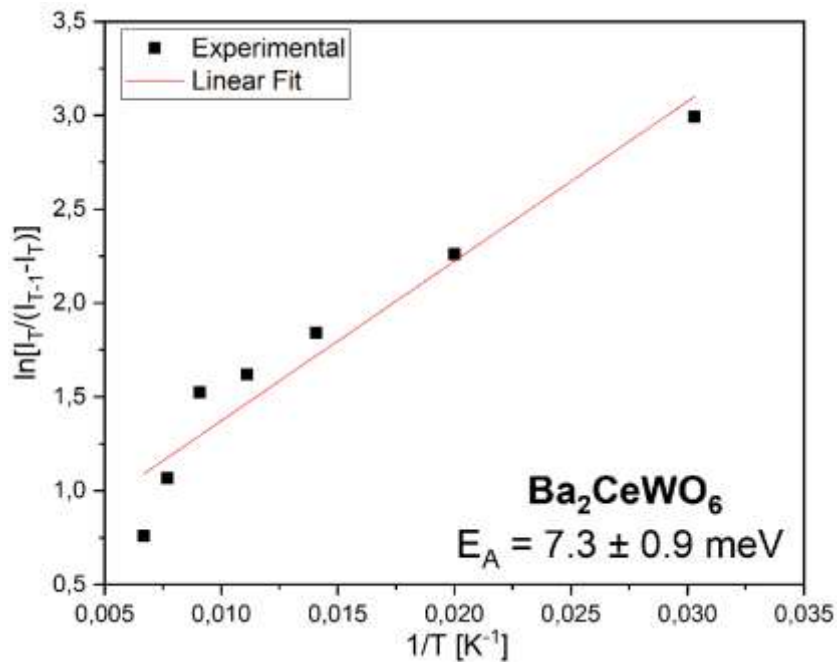


Figure A22 Activation energy estimate done for undoped BCW based on WO_x bands & modified Arrhenius equation. Measurements were done in the same pumped, helium-filled cryostat where no noticeable PL-related decomposition changes were observed under $\lambda_{\text{exc}} = 232$ nm.

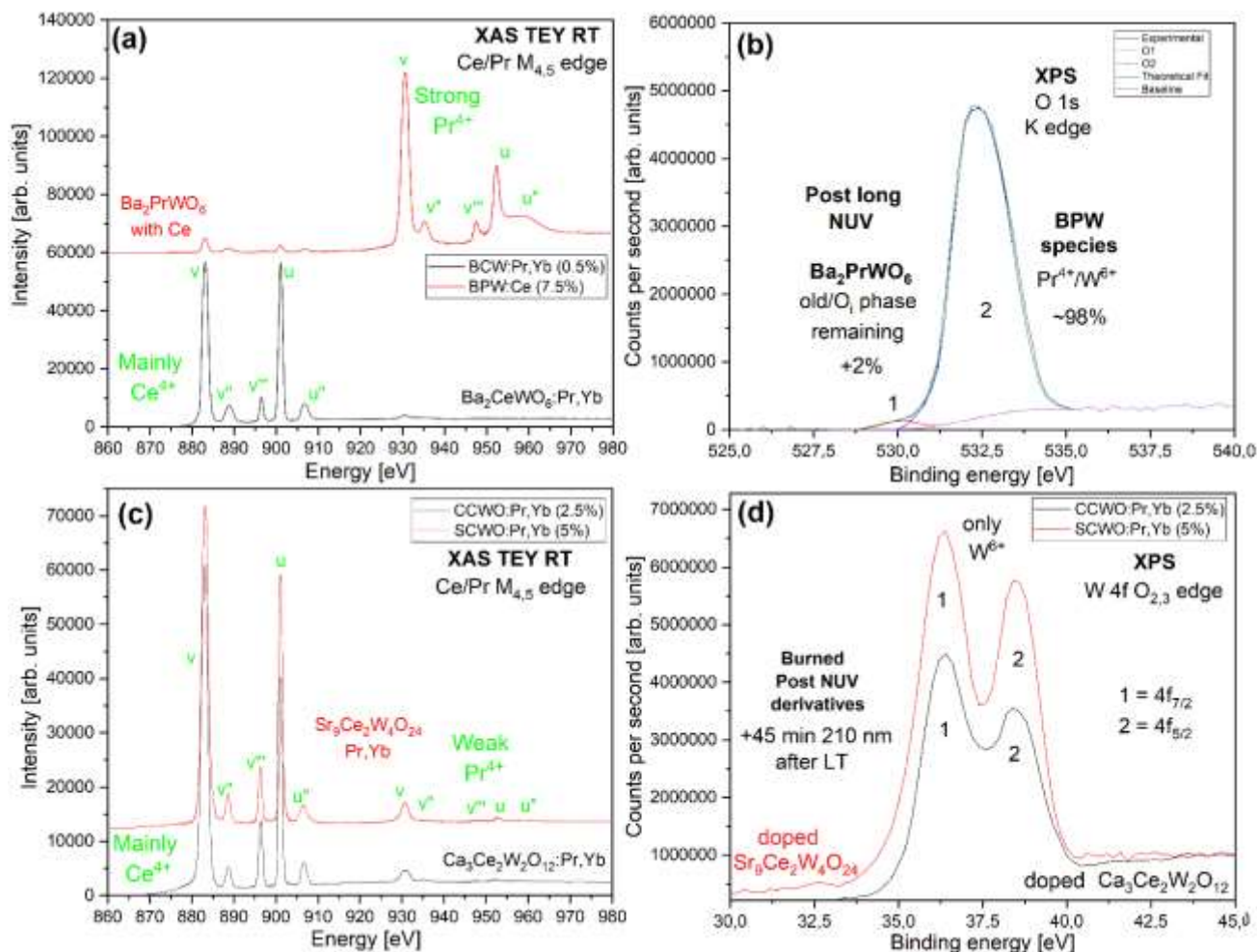
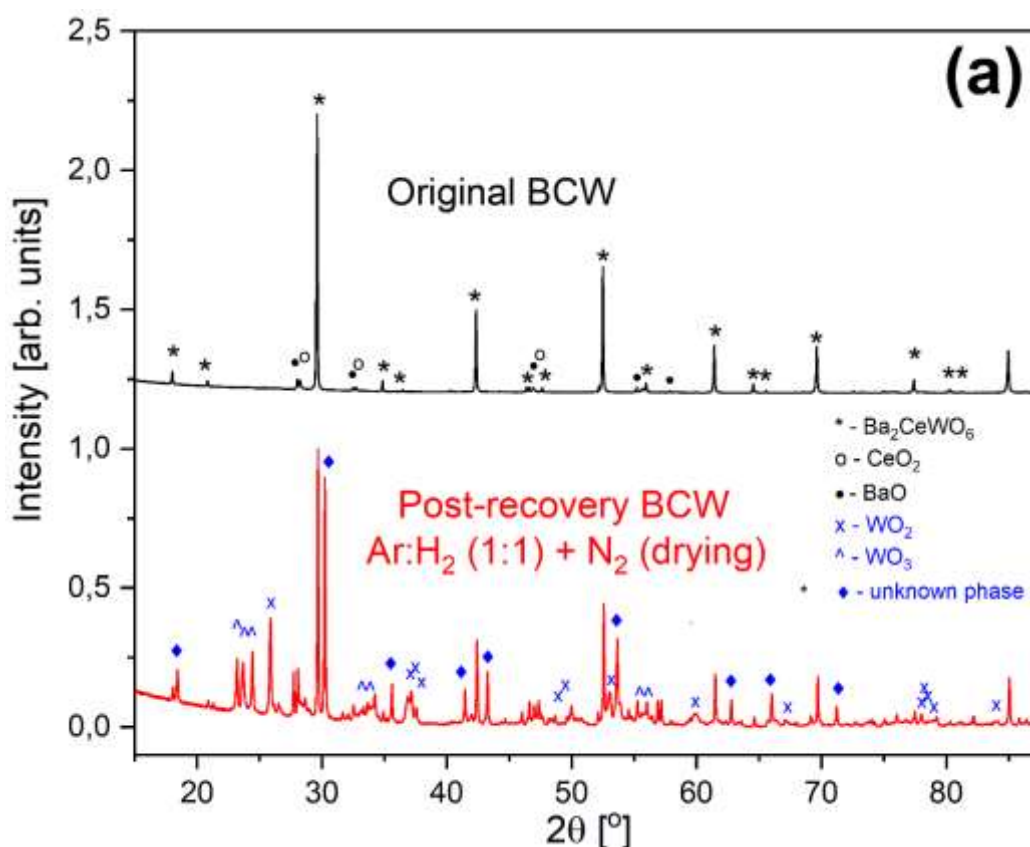


Figure A23 Supplementary, structural X-ray investigation of crucial atom sites after intense UV exposure featuring irreversible decomposition/oxidation signs in both doped double perovskites via complementary (a) Cerium, and (b) BCW's oxygen edges. Implanted derivatives also change with time after prolonged illumination (but through visual scorching), which fatally influences their PL capabilities and composition as seen on (c) RE & (d) W B-sites sells in comparison to Figures 19, 20 & 25. Intense excitation is needed to receive any response but with the most cautious approach.



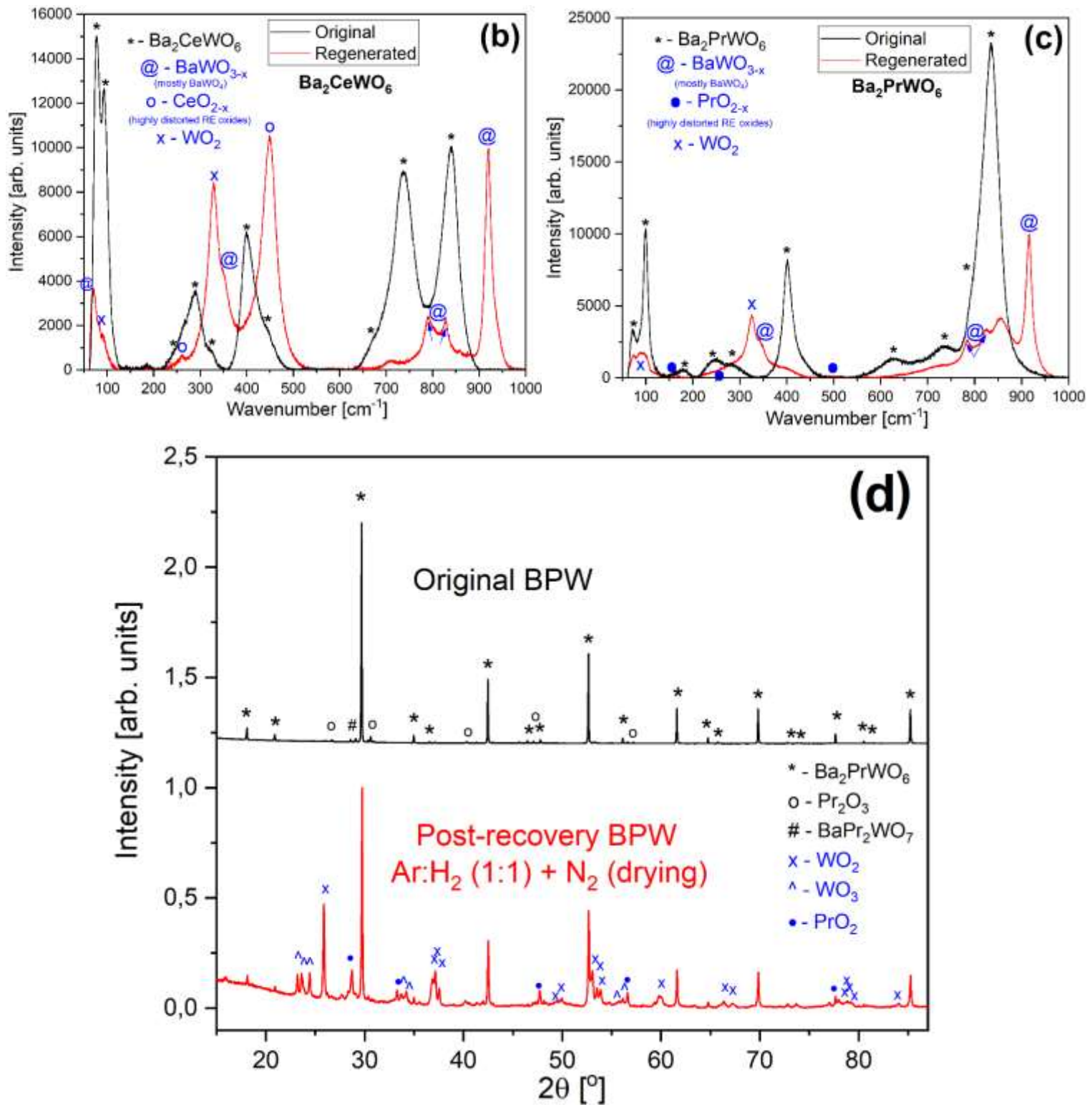


Figure A24 Ambient powder XRD diffraction patterns and complimentary Raman spectra of original (black) and theoretically regenerated (red) BCW (a,b) & BPW (c,d) barium double perovskites presented consecutively after HT SSR treatment in Ar:H₂ 1:1 and N₂ drying. Latter samples feature partial decomposition of the main matrix and separation of the primary B/B' (RE/W)O_x components with eventual oxidation (highlighted in insets as blue), deeming the materials unrecoverable post deep-UV exposure. The unknown phase marked with a diamond (♦) in BCW figure (a) seems to be the main phase with smaller lattice constants (presumably shrunk due to missing, crucial, B-site components). No other reliable match regarding that last issue could be found in the crystallographic database.

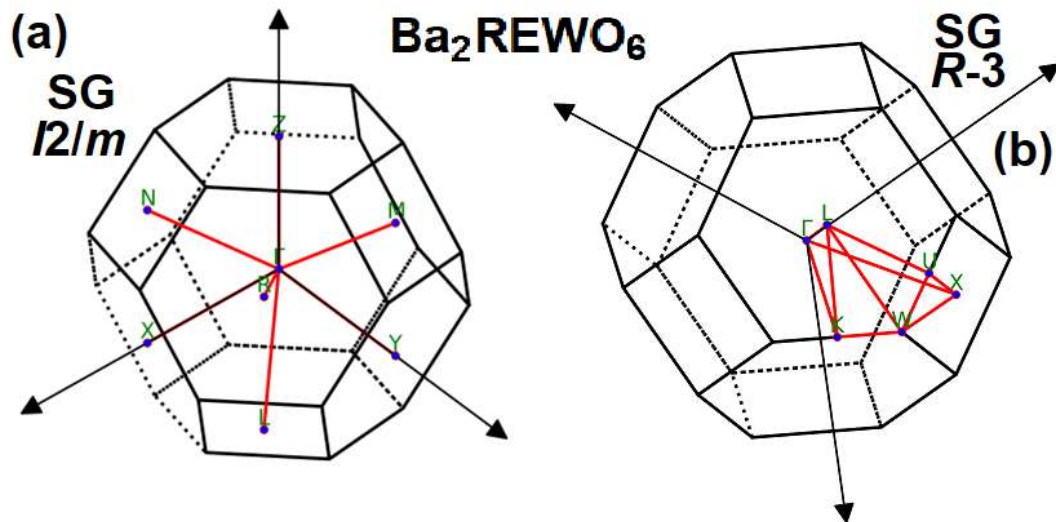


Figure A25 Depiction of Brillouin zone's high symmetry points referenced in Wannier/Tight binding equations for band gap calculations provided for barium RE tungstates' small unit cells in Figures 59 and A26. Figure (a) showcases $I2/m$ SG, and (b) alternate $R-3$.

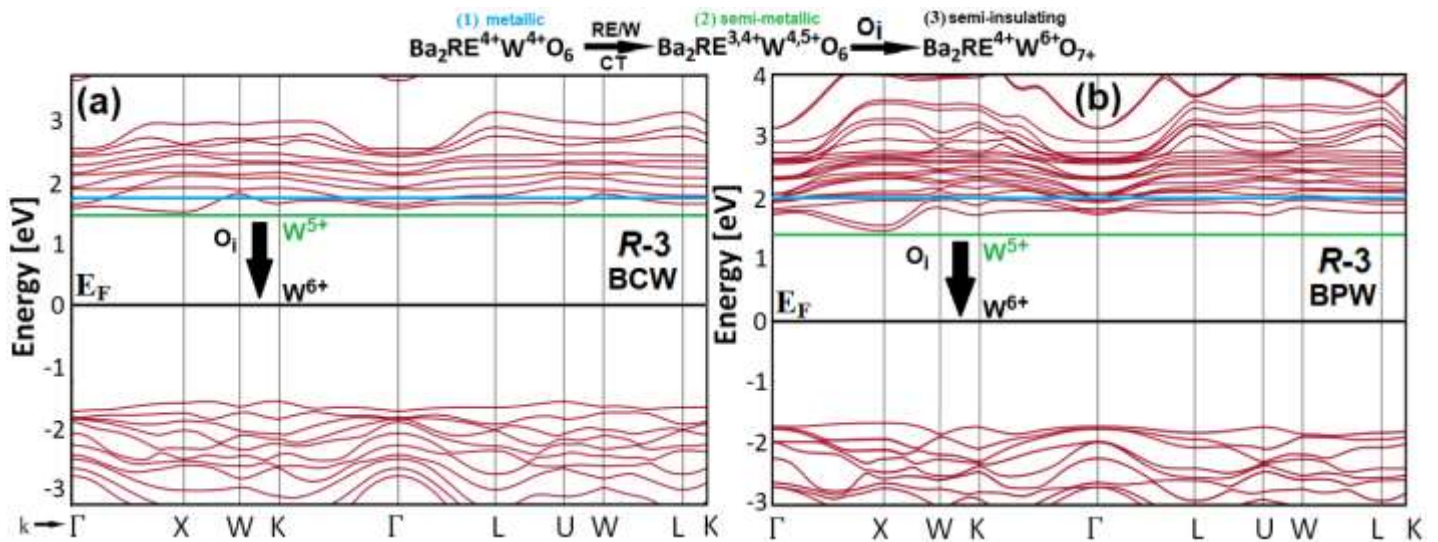


Figure A26 Tight-binding bandwidth estimations provided by VASP software using Wannier equations for both barium double perovskites (a) BCW & (b) BPW hosting alternate $R-3$ unit cells influenced by the O_i presence. As one can see, the evolving charge of W ions still does not account for such a broad emission which would require at least a $2\times$ bigger gap.

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### ◆ Education:

2016 – 2023 – Institute of Physics PAS – PhD in Solid State Physics,

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2013 – 2014r.– Gdansk University of Technology, Master's degree in Materials Engineering - anticorrosion & individual course studies chosen from Chemical Technology (organics),

2009 – 2013r.- Gdansk University of Technology, an Engineering degree in Chemical Technology – anticorrosion specialty (inorganic & physical chemistry).

### ◆ Employment: (including internships)

04.2022 – 02.2023 – Institute of Physics PAS – Technician: Physicist & Chemist,

02.2020 – 02.2023 – Institute of Physics PAS – Head of Preludium 17 ST5 grant,

08.09 – 08.10.2014 - Jotun Polska S.A – Quality control department – inspector anticorrosion assistant,

18.07 – 22.07.2014 - Polfarma S.A.– R&D – Statistical, pharmaceutical analysis,

01.04.2014 – 26.06.2014 & 1.08.2015 – 1.10.2015 – PPH Kamix Sp. z o.o. Sp. k. – Main Technologist assistant – Department of Production & quality control,

05.08.2013 – 18.08.2013 - FUU Rumia Sp. z o.o.– – Main Technologist assistant – Department of production & quality control,

06.08 – 20.08.2012 – Soda Polska Ciech Sp. z o.o. – Main Technologist assistant – 3-person team – Department of water treatment, carbonization & calcination.

### ◆ Responsibility and current scientific duties:

- Minor repairs, maintenance work, and calibration of spectroscopic devices for pressure and low-temperature luminescence, Raman, and FTIR measurements. Assistance in servicing optics and electronics (software-wise),

- Waste disposal and ordering chemicals experimental instruments in the docusafe invoice system with offer searching for grant purposes,
- Synthesizing, shaping, cutting, and polishing of metallic samples immersed in epoxy resins & those made into single crystals. Distillation & demineralization of water,
- High-temperature Solid State synthesis of inorganic compounds in inert and air conditions (mainly powders). Synthesis of organic compounds in enclosed glass systems involving tube furnaces, transformers, heating mantles, and other typical laboratory glassware,
- Research on numerous materials such as: clinopyroxenes doped with Eu and Mn; YAM, YAP, and LuAP oxides; BaWO<sub>4</sub> single crystals; Spectral mapping of graphene and black phosphorus structures; the topological character of rhenium sulfide; Research on GaN crystals doped with C, Mn, and Fe; Study of amorphization in vanadium and phosphorus whitlockites; Low-temperature measurements of gadolinium gallium garnets with cerium as a low-temperature sensor; Thermal diffusion studies of elements such as Ti, Cu, Cr, Co, Fe in lithium niobate and YIG; Spectral analysis of AgGaGeSe<sub>8</sub> used for electronic purposes; Studies of unique structures and their subtle phase transitions of first or second order like martensitic type Y<sub>4</sub>Al<sub>2</sub>O<sub>9</sub>:Ce; Specialty in researching perovskites and double perovskites – created four novel ones: cerium (3) and praseodymium (1) tungstates hosting Ba, Sr, and Ca.

◆ Hobbies & additional appliances:

- Courses and training as an internal auditor of the management systems: in the PN-EN ISO 17025:2005 laboratory; OHSAS 18001:2007; and environmental systems 14001:2004. Training from the State's Pharmaceutical Inspectorate regarding supervision of psychoactive, psychotropic, and addictive substances,
- Participant in projects: (subcontractor) Baltic TRAM (Leader: Prof. Jabłońska, IPPAS); NCN OPUS 9 (Leader: Prof. Suchocki, IPPAS), PUR-GRAF NCBiR (Leader: Prof. Strankowski, Gdańsk University of Technology); Project Manager of Preludium 17 ST5 panel (IPPAS),
- Operator of numerous spectrometers as part of MSc/Ph.D. work at the Gdańsk University of Technology and IMP PAN (Horiba Yobin Yvon and Renishaw apparatus) & visits at the Innovation Research Center of the State Higher School in Biała Podlaska - Ntegra Spectra C.

◆ Languages:

English – C1;  
 German – B1;  
 Swedish – A2

◆ Software skills:

- Microsoft Office, ISIS Draw, AutoCad, Mendelej, Origin, LaTeX, Grapher 9; Nanotechnologia: HyperChem Professional; ToupView. Gamry Framework, ZsimpWin, Labspec; VistaControl, GRAMS, Fityk & Wire.

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- Cat. B driving license (cars)
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## List of Publications

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3. **D. Włodarczyk**, M. Urban, M. Strankowski, *Chemical modifications of graphene and their influence on properties of polyurethane composites: a review*, Physica Scripta, 09.2016, Vol. **91** (10), Article ID 104003, doi:10.1088/0031-8949/91/10/104003 – [Corresponding author](#)
4. M. Strankowski, **D. Włodarczyk**, Ł. Piszczyk, J. Strankowska, *Thermal and Mechanical Properties of Microporous Polyurethanes Modified with Reduced Graphene Oxide*, International Journal of Polymer Science, 09.2016, Vol. **2016**, Article ID 8070327, <http://dx.doi.org/10.1155/2016/8070327>

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6. J. Kucińska-Lipka, I. Gubanska, O. Korchynskiy, K. Malysheva, M. Kostrzewa, **D. Włodarczyk**, J. Karczewski, H. Janik, *The influence of Calcium Glycerophosphate (GPCa) modifier on the physicochemical, mechanical and biological performance of polyurethanes applicable as biomaterials for bone tissue scaffold fabrication*, Polymers, 08.2017, Vol. **9** (8), pp. 329-350, doi:10.3390/polym9080329
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22. **D. Włodarczyk**, M. Amilusik, K. M. Kosyl, M. Chronik, K. Lawniczak-Jablonska, M. Strankowski, M. Zajac, V. Tsiumra, A. Grochot, A. Reszka, A. Suchocki, T. Giela, P. Iwanowski, M. Bockowski, H. Przybylinska, *Synthesis attempt and structural studies of novel A<sub>2</sub>CeWO<sub>6</sub> double perovskites (A<sup>2+</sup> = Ba, Ca) in and outside of ambient conditions*, ACS Omega, 04.2022, Vol. **7** (22), pp. 18381-18408, doi: 10.1021/acsomega.2c00669 - [Corresponding author](#).

## List of conferences

### Oral presentations

#### International conferences:

1. **D. Włodarczyk**, K. M. Kosyl, W. Paszkowicz, J. Z. Domagala, O. Ermakova, R. Minikayev, A. Suchocki, A. Shekhovtsov, M. Kosmyna, C. Popescu, F. Fauth – EURODIM 18 – Bydgoszcz, 8-13.07.2018 – *Structural studies focused on Ca<sub>9</sub>R(VO<sub>4</sub>)<sub>7</sub> (R = La, Nd, Gd) whitlockites under elevated pressure;*
2. **D. Włodarczyk**, M. Berkowski, M. Głowacki, S. Kaczmarek, Z Kowalski, A. Wittlin, H Przybylinska, L-I Bulyk, Ya. Zhydachevskii, A. Suchocki – Phosphor Safari & VI IWASOM – Gdansk, 9-14.07.2017 - *Optical studies of BaWO<sub>4</sub>:Ce,Na;*
3. **D. Włodarczyk**, M. Amilusik, M. J. Chronik, K. M. Kosyl, V. Tsiumra, M . Strankowski, T. Giela, M. Zajac, A. Grochot, A. Reszka, P. Iwanowski, M. Boćkowski, K. Lawniczak-Jablonska, H. Przybylinska, A. Suchocki – LUMDETR XI – Bydgoszcz, 12-17.09.2021 – *Structural and optical studies of novel, cerium-tungstate double perovskites doped with rare-earth ions;*
4. **D. Włodarczyk**, M. Amilusik, R. Islam, L-I. Bulyk, V. Tsiumra, M. Zajac, A. Suchocki, M. Bockowski, K. Jablonska, H. Przybylinska - FM&NT NIBS – Ryga, 3-6.07.2022 – *Novel double perovskites and their derivatives hosting rare earth ions – a prelude to charge transfer phenomenon;*
5. **D. Włodarczyk**, M. Amilusik, R. Islam, L-I. Bulyk, V. Tsiumra, M. Zajac, A. Suchocki, M. Bockowski, K. Jablonska, H. Przybylinska - VII IWASOM – Gdansk, 1-15.07.2022 – *Spectral and X-ray comparison between novel double perovskites and their derivatives based on tungstates doped with rare earths;*

## Polish Conferences:

1. **D. Włodarczyk**, A. Feliniak, M. Lubinska, K. Kopczyńska, E. Nalborska, S. Krakowiak – IV Ogólnokrajowa Konferencja Naukowa Młodzi Naukowcy w Polsce, Badania i Rozwój – Gdansk, 27.11 2016 - *Korozja materiałów konstrukcyjnych w zbiornikach browarniczych*;
2. **D. Włodarczyk**, E. A. Miśta, I. Żmuda-Trzebiatowska, P. Kalbarczyk, M. Kolbadinejad – IX Międzyuczelniane Seminarium Kół Naukowych WAT – Warsaw, 19-20.05 2016 – *Analiza chemiczna dekoracyjnych kafelków ściennych z Iranu*;
3. **D. Włodarczyk**, L-I. Bulyk, M. Głowacki, M. Berkowski, A. Suchocki - IV Interdyscyplinarna Akademicka Konferencja Ochrony Środowiska - Gdansk 5-7.04.2019 - *High-pressure, and low-temperature optical studies of BaWO<sub>4</sub>:Ce,Na single crystals*;
4. **D. Włodarczyk**, M. Tomczyk, D. A. Pawlak, A. Suchocki - XII Międzyuczelniane Seminarium Kół Naukowych WAT - Warsaw, 4-5.06.2019 - *Comparison of optical and structural properties between Ba<sup>2+</sup> and Zn<sup>2+</sup> tungstates*;
5. **D. Włodarczyk**, M. Amilusik, R. Islam, L-I. Bulyk, V. Tsiumra, M. Zajac, A. Suchocki, M. Bockowski, K. Jablonska, H. Przybylinska – VII Interdyscyplinarna Akademicka Konferencja Ochrony Środowiska – Gdansk, 21-23.09.2022 - *Synthesis and comparison of structural properties between tungstate double perovskites and their derivatives doped with rare earth ions*;

## Posters

1. **D. Włodarczyk**, K. M. Kosyl, K. Izdebska, W. Paszkowicz, M. Kosmyna, A. Shekhovtsov, A. Suchocki – OMEE 2017 – Lviv, 28.05-2.06 2017 - *High-pressure luminescence and Raman investigation of Ca<sub>9</sub>Nd(VO<sub>4</sub>)<sub>7</sub> whitlockite-like crystals*;
2. **D. Włodarczyk**, K. M. Kosyl, V. Tsiumra, M. Amilusik, M. Chronik, M. Zajac, R. Islam, M. Bockowski, H. Przybylinska, A. Suchocki - VII Interdyscyplinarna Akademicka Konferencja Ochrony Środowiska – Gdansk, 21-23.09.2022 - *Rare earth doped double perovskites & their tungstate derivatives created for UV detection*;
3. **D. Włodarczyk**, K. M. Kosyl, M. Strankowski, M. Chronik, M. Zajac, A. Suchocki - XVI Ogólnokrajowa Konferencja Naukowa Młodzi Naukowcy w Polsce - Badania i Rozwój – Online conference, 28.11.2022 - *Polymorphism analysis of novel tungstate double perovskites and their derivatives doped with chosen rare earth ions*;
4. **D. Włodarczyk**, M. Amilusik, V. Tsiumra, L-I. Bulyk, R. Islam, A. Suchocki - XVI Ogólnokrajowa Konferencja Naukowa Młodzi Naukowcy w Polsce - Badania i Rozwój – Online conference, 28.11.2022 - *Anomalous luminescent behavior of novel double perovskites and their rare-earth doped tungstate derivatives*.

## Seminars and Workshops

1. ON4 seminar – IP PAS - 14.06.2016 - *Od pomysłu do przemysłu, czyli zastosowanie spektroskopii ramanowskiej w praktyce*;
2. ON 4 seminar - IP PAS – 17.10.2017 - *Optical studies of BaWO<sub>4</sub>:Ce, Na*;
3. ON 4 seminar - IP PAS – 13.11.2018 - *Structural studies focused on Ca<sub>9</sub>R(VO<sub>4</sub>)<sub>7</sub> (R = La, Nd, Gd) whitlockites under elevated pressure*;
5. X Symposium of PhD Students PAS - Białobrzegi, 04.06.2018 – “*Optical studies of BaWO<sub>4</sub>:Ce,Na*”;
6. XI Symposium of Ph.D. Students PAS – Serock, 28-30.05.2019 - *Phase-transition and energy-transfer phenomena in modeled double perovskites doped with rare-earth ions*;
7. ON4 Seminar - IP PAS - 13.10.2020 – *Structural and optical studies of novel, doped A<sub>2</sub>CeWO<sub>6</sub> double perovskites (A=Ca, Sr or Ba<sup>2+</sup>)*.

## Projects

### Leader

1. 36-month project Preludium 17, NCN nr 2019/33/N/ST5/02317 (internal no. P-506/P) pt.: *Structural and optical studies of novel, doped A<sub>2</sub>CeWO<sub>6</sub> double perovskites*.

### Stipendist/Subcontractor

1. 31-month stipendist in project NCN OPUS-9, nr UMO-2015/17/B/ST5/01658 (internal no. P-441/0) pt.: „*Modification of solar energy through down-conversion on the basis of oxide materials doped with Ytterbium for photovoltaic applications*” under supervision of Prof. Andrzeja Suchocki.
2. Baltic TRAM; Project no. #R002; Financed by Interreg Baltic Sea Region Programme (85%) & IP PAS (15%); Supervisor – Prof. Krystyna Jabłońska; Partially employed as a subcontractor in between 04-10-2017 & 04-11-2018 regarding shungite research;
3. Subcontractor of PUR-GRAF NCBR project no. GRAF-TECH/NCBR/11/08/2013 regarding *Production of polyurethane nanocomposites doped with thermally reduced graphene oxide used towards shoe industry*; Project leader – Ph.D. hab. eng. Michał Strankowski; Time of employment 04.2015 to 04.2016;